

Bridgewater ex dono *Auctoris*
THE

DESCRIPTION AND VSE OF THE SECTOR.

For such as are studious of
Mathematicall practise.



LONDON,
Printed by WILLIAM IONES.
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THE
DESCRIPTION

SECTOR

For the use of the
Inventor

Printed by W. E. Lockhart
1853



HONORATISSIMO

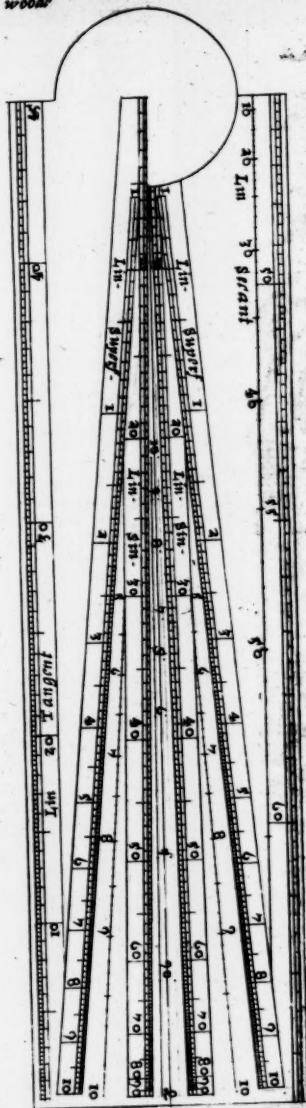
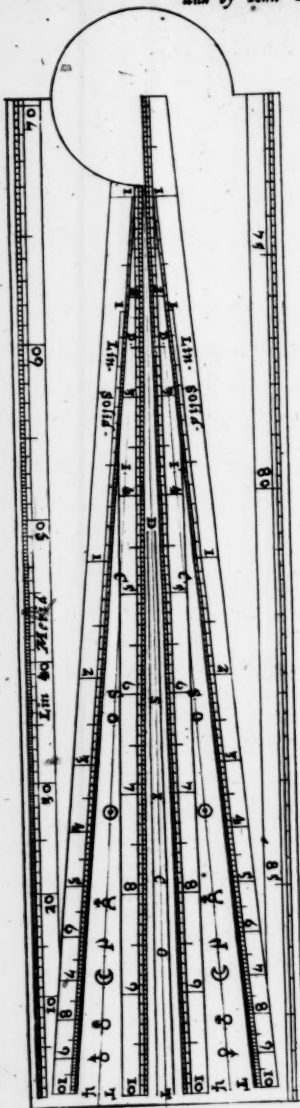
DOMINO *Dn. IOHANNI*
COMITI de BRIDGEWATER,
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BARONI de ELLESMERE,

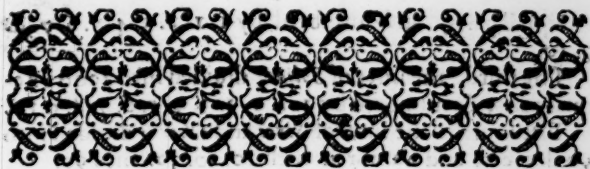
HVNC SECTOREM

D. D. D.

EDM. GYNTER.

This sector is mad by Elias Allen in Brass
and by John Tompshon in wood





THE
FIRST BOOKE
OF THE
SECTOR.

CHAP. I.

*The Description, the making, and the generall use
of the Sector.*



Sector in Geometrie, is a figure comprehended of two right lines containing an angle at the center, and of the circumference assumed by them. This *Geometrical instrument* hauing two legs containing all varietie of angles, & the distance of the feete, representing the subtenses of the circumference, is therefore called by the same name.

It containeth 12 feuerall lines or scales, of which 7 are generall, the other 5 more particular. The first is the scale of *Lines* diuided into 100 equall parts, and numbred by
1. 2. 3. 4. 5. 6. 7. 8. 9. 10.

The second, the lines of *Superficies* diuided into 100
B
unequall

vnequall parts, and numbred by 1. 1. 2. 3. 4. 5. 6. 7. 8. 9. 10.

The third, the lines of *Solids*, diuided into 1000 vnequall parts, & numbred by 1. 1. 1. 2. 3. 4. 5. 6. 7. 8. 9. 10.

The fourth, the lines of *Sines* and *Chords* diuided into 90 degrees, and numbred with 10. 20. 30. vnto 90.

These foure lines of *Lines*, of *Superficies*, of *Solids*, and of *Sines*, are all drawne from the center of the *Sector* almost to the end of the legs. They are drawne on both the legs, that euery line may haue his fellow. All of them are of one length, that they may answere one to the other. And euery one hath his parallels, that the eye may the better distinguish the diuisions. But of the parallels those onely which are inward most containe the true diuisions.

There are three other generall lines, which because they are infinite are placed on the side of the *Sector*. The first a line of *Tangents*, numbred with 10. 20. 30. 40. 50. 60. signifying so many degrees from the beginning of the line, of which 45 are equall to the whole line of *Sines*, the rest follow as the length of the *Sector* will beare.

The second, a line of *Secants*, diuided by pricks into 60 degrees, whose beginning is the same, with that of the line of *Tangents*, to which it is ioyned.

The third, is the *Meridian* line, or line of *Rumbs*, diuided vnequally into degrees, of which the first 70 are almost equall to the whole line of *Sines*, the rest follow vnto 84 according to the length of the *Sector*.

Of the particular lines inserted among the generall, because there was voyd space, the first are the lines of *Quadrature* placed betweene the lines of *Sines*, and noted with 10. 9. 8. 7. 6. 5. 4. 3. 2. 1.

The second, the lines of *Segments* placed betweene the lines of *Sines* and *Superficies* diuided into 50 parts, and numbered with 5. 6. 7. 8. 9. 10.

The third, the lines of *Inscribed bodies in the same Sphere*, placed betweene the scales of *Lines*, and noted with D. S. 1. C. O. T.

The

The fourth, the lines of *Equated bodies*, placed between the lines of *Lines* and *Solids*, and noted with *D.I.C.S.O.T.*

The fifth, are the lines of *Metalls*, inserted with the lines of *Equated bodies* (there being room sufficient) and noted with these Characters. \odot . \oint . \mathcal{H} . \mathcal{D} . \mathcal{Z} . \mathcal{S} . \mathcal{U} .

There remaine the edges of the *Sector*, and on the one I haue set a line of *Inches*, which are the twelfth parts of a foote English: on the other a lesser line of *Tangents*, to which the *Gnomon* is *Radius*.

2. Of the making of the Sector.

Let a *Ruler* be first made either of brasse or of wood, like vnto the former figure, which may open and shut vpon his center. The head of it may be about the twelfth part of the whole length, that it may beare the moueable foote, and yet the most part of the diuisions may fall without it. Then let a moueable *Gnomon* be set at the end of the moueable foote, and there turne vpon an *Axle*, so as it may sometime stand at a right angle with the feete, and sometimes be inclosed within the feet. But this is well knowne to the workeman.

For drawing of the lines. Vpon the center of the *Sector*, and semidiameter somewhat shorter then one of the feet, draw an occult arke of a circle, crossing the closure of the inward edges of the *Sector* about the letter *T*.

In this arke, at one degree on either side from the edge, draw right lines from the Center, siting them with Parallels and diuide them into an hundred equall parts, with subdivisions into 2.5. or 10. as the line will beare, but let the numbers set to them, be onely 1.2.3.4.&c.vnto 10. as in the example. These lines so diuided, I call the lines or scales of *Lines*; and they are the ground of all the rest.

In this Arke at 5 degrees on either side, from the edge neere *T*, drawe other right lines from the Center, and sit them with Parallels. These shall serue for the lines of *Solids*.

The inscription of the lines.

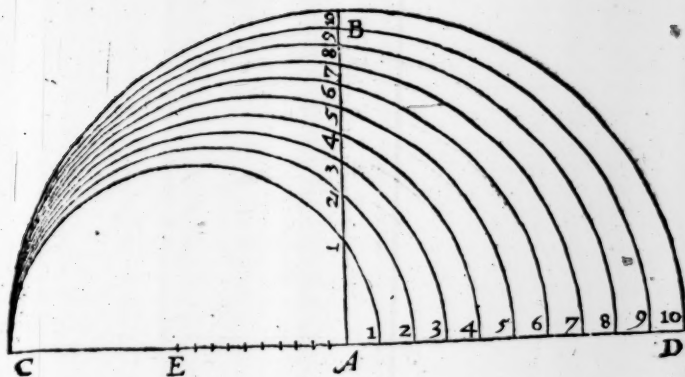
Then on the other side of the *Sector* in like manner, vpon the Center & equall Semidiameter, drawe another like Arke of a circle: & heere againe at one ~~more~~ degree on either side frō the edge neere the letter *Q* draw right lines from the Center, and fit them with parallells. These shall serue for the lines of *Sines*.

At 5 Degrees on either side from the edge neere *Q* drawe other right lines from the center, and fit them with parallels: these shall serue for the lines of *Superficies*.

These foure principall lines being drawne, and fitted with parallels, we may drawe other lines in the middle betweene the edges and the lines of *Lines*, which shall serue for the lines of *inscribed bodies*, and others betweene the edges and the *Sines* for the lines of *quadrature*. And so the rest as in the example.

3 *To diuide the lines of Superficies.*

Seeing like *Superficies* doe hold in the proportion of their *homologall* sides duplicated, by the 29 Pro. 6 lib. *Euclid*. If you shall find meane proportionals between the whole side, and each hundred part of the like side, by the 13 Pro. 6 lib. *Euclid*, all of them cutting the same line, that line so cut shall containe the diuisions required, wherefore vpon the center *A* and Semidiameter equall to the line of *Lines*, describe a Semicircle *ACB D*, with *AB* perpendicular to the diameter *CD*. And let the Semidiameter *AD* be diuided as the line of *Lines* into an hundred parts, & *A E* the one halfe of *AC* diuided also into an hundred parts, so shall the diuisions in *A E* be the centers from whence you shall describe the Semicircles *C 10*. *C 20*. *C 30*. &c. diuiding the lin *AB* into an hundred vnequall parts: and this line *AB* so diuided shall be the line of *Superficies*, and must be transferred into the *Sector*. But let the numbers set to them be onely 1. 2. 3. vnto 10. as in the example.



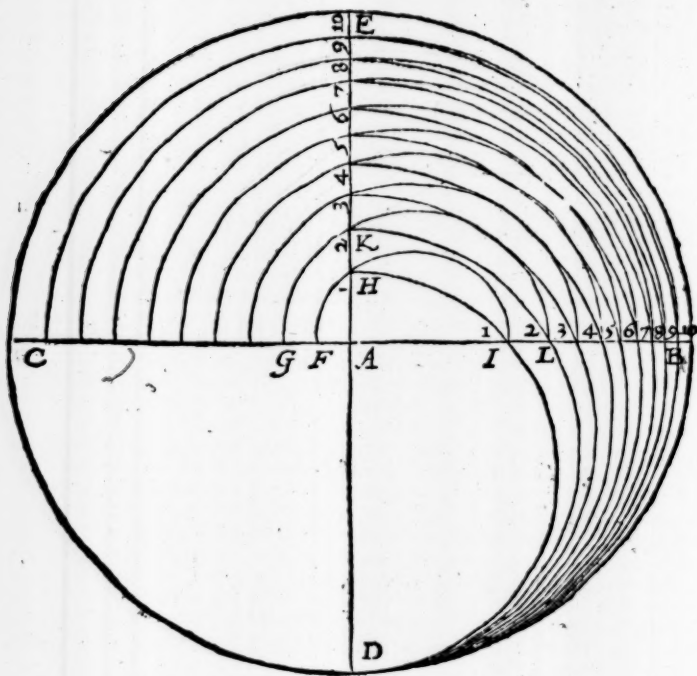
4 To divide the lines of Solids.

Seeing like Solids do hold in the proportion of their *homologall* sides triplicated, if you shall finde two meane proportionalls between the whole side & each thousand part of the like side : all of them cutting the same two right lines, the former of those lines so cut, shall containe the diuisions required.

Wherefore vpon the center *A* & Semidiameter equall to the line of *Lines*, describe a circle and diuide it into 4 equall parts *C E B D*, drawing the crosse diameters *C B*, *E D*. Then diuide the semidiameter *A C*, first into 10 equall parts, and betweene the whole line *AD* & *AP* the tenth part of *AC*, seeke out two meane proportionall lines *AI* and *AH*. againe betweene *AD* and *AG* being two tenth parts of *AC*, seeke out two meane proportionalls *AL* and *AK*, and so forward in the rest. So shall the line *AB* be diuided into 10 vnequall parts.

B 3

Secondly



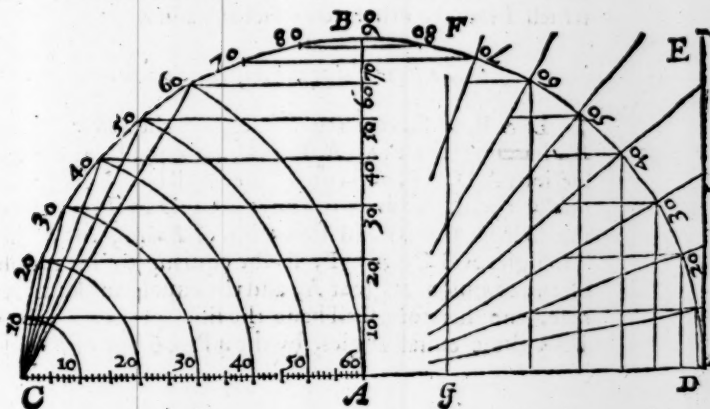
Secondly, divide each tenth part of the line *AC* into 10 more, and betwene the whole line *AD*, and each of them, seeke out two meane proportionalls as before: So shall the line *AB* be divided now into an hundred vn-equall parts.

Thirdly, If the length will beare it, subdiuide the line *AC* once againe, each part into ten more: and betwene the whole line *AD* and each subdiuision, seeke two meane

meane proportionalls as before. So should the line *AB* be now diuided into 1000 parts. But the ruler being short, it shall suffice, if those 10 which are nearest the center be exprest, the rest be vnderstood to be so diuided, though actually they be diuided into no more then 5 or 2. and this line *AB* so diuided shall be the line of *Solids*, and must be transferred into the *Sector*: But let the numbers set to them be onely 1. 2. 3. 4. 5. 6. vnto 10. as in the example.

5 To diuide the lines of Sines and Tangents on the side of the Sector.

Vpon the center *A*, and semidiameter equall to the line of *Lines*, describe a semicircle *ABCD*, with *AB* perpendicular to the diameter *CD*. Then diuide the quadrants *CB*, *BD*, each of them into 90. and subdiuide each degree into 2 parts: For so, if streight lines be drawne parallel to the diameter *CD*, through these 90, and their subdiuisions they shall diuide the perpendicular *AB* vnequally into 90.



And this line *A B* so diuided shall be the line of *Sines*, and must be transferred into the *Sector*. The numbers set to them are to be 10. 20. 30. &c. vnto 90 as in the example.

If now in the poynt *D*, vnto the diameter *C D*, we shall raise a perpendicular *D E*, and to it drawe streight lines from the center *A*, through each degree of the quadrant *D B*. This perpendicular so diuided by them shall be the line of *Tangents*, & must be transferred vnto the side of the *Sector*. The numbers set to them, are to be 10. 20. 30. &c. as in the example.

If betweene *A* and *D*, another streight line *G F*, be drawne parallell to *D E*, it will be diuided by those lines from the center in like sort as *D E* is diuided, and it may serue for a lesser line of *Tangents*, to be set on the edge of the *Sector*.

These lines of *Sines* and *Tangents*, may yet otherwise be transferred into the *Sector* out of the line of *Lines*, (or rather out of a diagonall Scale equall to the line of *Lines*) by tables of *Sines* and *Tangents*. In like manner may the lines of *Superficies*, be transferred by tables of square rootes; and the line of *Solids*, by tables of cubique rootes: which I leaue to others to extract at leasure.

6 To shew the ground of the Sector.

Let *A B*, *A C*, represent the leggs of the *Sector*: then seeing these two *A B*, *A C*, are equall, and their sections *A D*, *A E*, also equall, they shall be cut proportionally: and if we draw the lines *B C*, *D E*, they will be parallell by the second Pro. 6 lib. of *Euclid*, and so the Triangles *A B C*, *A D E*, shalbe equiangle; by reason of the common angle at *A*, and the equall angles at the base, and therefore shall haue the sides proportionall about those equall angles, by the 4 Pro. 6 lib. of *Euclid*.

The



The side A D, shalbe to the side A B, as the basis D E, vnto the parallell basis B C, and by conuerſion A B, ſhall be vnto A D, as B C, vnto D E: and by permutation A D, ſhall be vnto D E, as A B, to B C.&c. So that if A D, be the fourth part of the ſide A B, then D E, ſhall alſo be the fourth part of his parallell basis B C. The like reaſon holdeth in all other ſections.

7 To ſhew the generall uſe of the Sector.

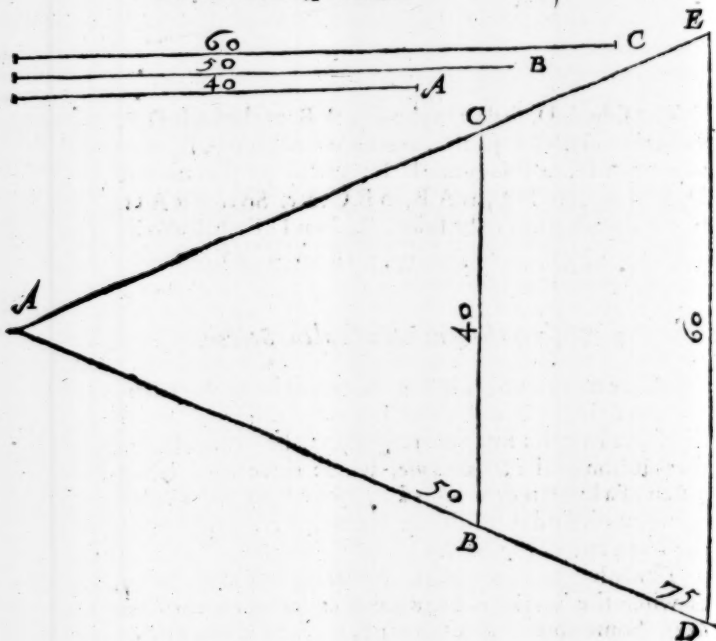
Here may ſome cōcluſions be wrought by the Sector, euen then when it is ſhut, by reaſon that the lines are all of one length: but generally the uſe hereof conſiſts in the ſolution of the *Golden rule*, where three lines being giuen of a known denomination, a fourth proportionall is to be found. And this ſolution is diuerſe in regard both of the *lines*, and of the *entrance* into the worke.

The ſolution in regard of the *lines* is ſometimes *ſimple*, as when the worke is begun and ended vpon the ſame *lines*. Sometimes it is *compound*, as when it is begun on one kind of *lines*, and ended on another. It may be begun vpon the lines of *Lines*; & finiſhed vpon the lines of *Superficies*. It may begin on the *Sines*, and end on the *Tangents*.

C

The

The solution in regard of the *entrance* into the worke, may be either with a *parallel* or else *laterall* on the side of the Sector, I cal it *parallel entrance*, or entring with a parallel, when the two lines of the first denomination are applied in the parallels, and the third line, and that which is sought for, are on the side of the Sector. I call it *laterall entrance*, or entring on the side of the Sector, when the two lines of the first denomination are one the side of the Sector, and the third line and that which is to be found out, doe stand in the parallels.

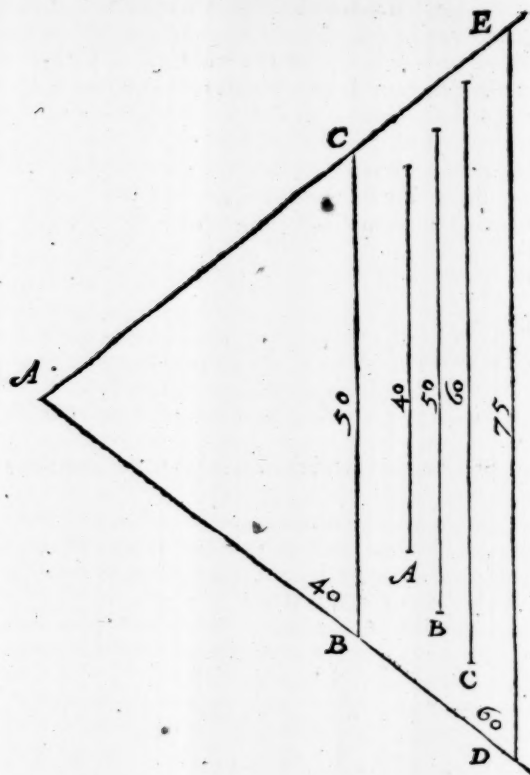


As for example, let there be given three lines A, B, C, to which I am to find a fourth proportionall. let A, measured in the line of *lines*, be 40, B 50, and C 60, and suppose the question be this. If 40 *Months* give 50 *pounds*, what shal 60? Here are lines of two denominatiōs, one of *months*, another of *pounds*, and the first with which I am to enter must be that of 40 *months*. If then I would enter with a *parallell*, first I take A, the line of 40, and put it over as a *parallell* in 50, reckoned in the line of *lines*, on either side of the *Sector* from the center, so as it may be the Base of an *Isofcheles* triangle B A C, whose sides A B, A C are equal to B, the line of the second denomination.

Then the *Sector* being thus opened, I take C the line of 60, betweene the feete of the compasses, and carrying them *parallell* to B C, I finde them to crosse the lines A B, A C, on the side of the *Sector* in D and E, numbred with 75, wherefore I conclude the line A D, or A E, is the fourth proportionall and the correspondend number 75 which was required.

But if I would enter on the side of the *Sector*, then would I dispose the lines of the first denomination A and C, in the line of *Lines*, on both sides of the *Sector*, in A B, A C, & in A D, A E, so as they should all meete in the center A, and then taking C the line of the second denomination put it over as a *parallell* in B C, that it may be the Basis of the *Isofcheles* triangle B A C, whose sides A B, A C, are equall to A, the first line of the first denomination, for so the *Sector* being thus opened, the other *parallell* from D to E, shall be the fourth proportionall which was required, and if it be measured with the other lines, it shal be 75, as before.

In both these manner of operations, the two first lines do serue to opē the *Sector* to his due angle, the difference betweene them is especially this, that in *parallell entrance*, the two lines of the first denomination, are placed in the *parallels* B C, D E, & in *latterall entrance* they are placed on both sides of the *Sector*, in A B, A D and in A C, A E.



Now in *simple solution* which is begun and ended, vpon the same kinde of lines, it is allone which of the two latter lines be put in the secōd or third places. As in our exāple we may say, as 40 are to 50, so 60 vnto 75, or else as 40 are to 60, so 50 vnto 75. And hence it cōmeth that we may enter both with a *parrallell*, & on the sides two manner of wayes at either entrance, and so the most part of questions may

may be wrought 4 seuerall wayes, though in the propositions following, I mention onely that which is most conuenient. Thus much for the generall vse of the *Setter*, which being considered and well vnderstood, there is nothing hard in that which followeth.

CHAP. II.

The vse of the Scale of Lines.

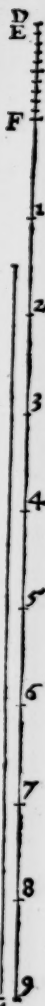
1. To set downe a Line, resembling any giuen parts or fraction of parts.

THe lines of *Lines* are diuided actually into 100 parts, but we haue put onely 10 numbers to them. These we would haue to signifie either themselves alone, or ten times themselves, or an hundred times themselves, or a thousand times themselves, as the matter shall require. As if the numbers giuen be no more then 10, then we may thinke the lines onely diuided into 10 parts according to the numbers set to them. If they be more then 10, and not more then 100, then either line shall containe 100 parts, and the numbers set by them shall be in value 10. 20. 30. &c. as they are diuided actually. If yet they be more then 100, then euery part must be thought to be diuided into 10, and either line shall be 1000 parts, and the numbers set to them shall be in value 100. 200. 300, and so forward still increasing themselves by 10. This being presupposed, we may number the parts and fraction of parts giuen in the line of *lines*; and taking out the distance with a paire of compasses, set it by, for the line so taken shall resemble the number giuen.

In this manner may we set downe a line resembling 75, if either we take 75 out of the hundred parts, into which one of the line of *lines* is actually diuided, and note it in A, or $7\frac{1}{2}$ of the first 10 parts, and note it in B, or onely $\frac{1}{2}$ of one of those hundred parts, and note it in C. Or

C 3

if CBA



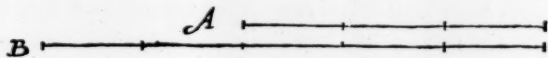
if this be either to great or to small, we may run a Scale at pleasure, by opening the compasse to some small distance, and running it ten times over; then opening the compasse to these ten, run them over nine times more, & set figures to them as in this example, and out of this we may take what parts we will as before.

To this end I have diuided the line of inches on the edge of the *Sector*, so as one inch containeth 8 parts, another 9, another 10, &c. according as they are figured, and as they are distant from the other end of the *Sector*, that so we might haue the better estimate.

2 To encrease a line in a given proportion.

3 To diminish a line in a given proportion.

TAKE the line giuen with a paire of compasses, and open the *Sector*, so as the feete of the compasses may stand in the points of the number giuen, then keeping the *Sector* at this angle, the parallell distance of the points of the number required, shall giue the line required.



Let *A*, be a line giuen to be increased in the proportion of 3 to 5. First I take the line *A*, with the compasses, and open the *Sector* till I may put it over in the poynts of 3 and 3, so the parallell betweene the poynts of 5 & 5, doth giue me the line *B*, which was required.

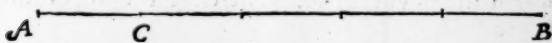
In like manner, if *B*, be a line giuen to be diminished in the proportiō of 5 to 3, I take the line *B* & to it open the *Sector* in the poynts of 5, so the parallell betweene the points of 3, doth giue me the line *A*, which was required.

If this manner of worke doth not suffice, we may multiply or diuide the numbers giuen by 1, or 2, or 3, &c. And so worke by their numbers *equimultiples*, as for 3 and

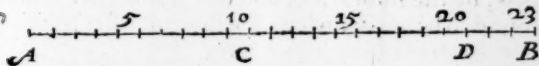
and 5, wee may open the *Sector* in 6 and 10, or else in 9 and 15, or else in 12 and 20, or in 15 and 25, or in 18 and 30, &c.

4 To divide a line into parts given.

Take the line giuen, and open the *Sector* according to the length of the said line in the points of the parts, wherevnto the line should be diuided, then keeping the *Sector* at this angle, the parallell distance betweene the points of 1 and 1 shall diuide the line giuen into the parts required.



Let A B, be the line giuen to be diuided into five parts, first I take this line A B, and to it open the *Sector* in the points of 5 and 5, so the parallell betweene the points of 1 and 1, doth giue me the line A C, which doth diuide it into the parts required.

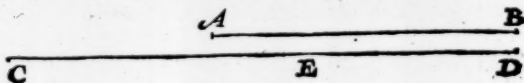


Or let the like line A B, be to be diuided into twenty three parts. First I take out the line and put it vpon the *Sector* in the points of 23, then may I by the former proposition diminish it in A C, C D, in the proportion of 23, to 10, and after that diuide the line A C into 10, &c. As before.

5 To finde a proportion betweene two or more right lines giuen.

Take the greater line giuen, and according to it open the

the *Sector* in the points of 100 and 100, then take the lesser lines seuerally, & carry them parallell to the greater, till they stay in like points, so the number of points wherein they stay, shall shew their proportion vnto 100.



Let the lines giuen be *AB, CD*, first I take the line *CD*, & to it open the *Sector* in the points of 100, and 100, then keeping the *Sector* at this angle, I enter the lesser line *AB*, parallell to the former, and find it to crosse the lines of *Lines* in the poynts of 60. Wherefore the proportion of *AB* to *CD*, is as 60 to 100.

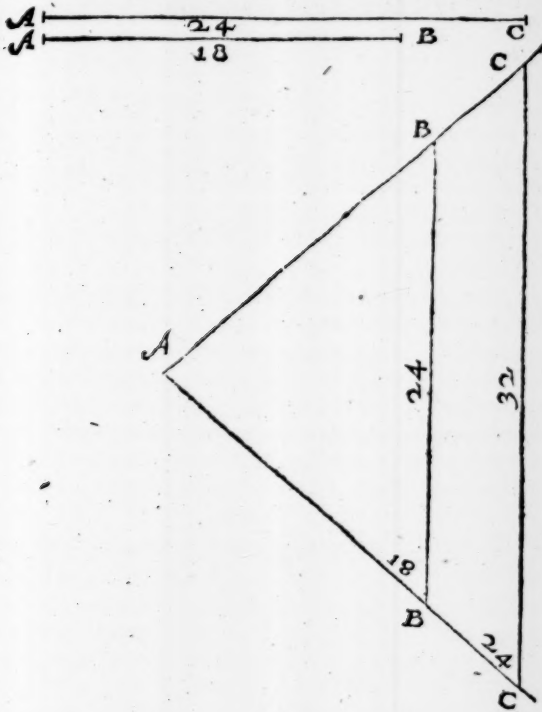
Or if the line *CD*, be greater then can be put ouer in the poynts of 100, then I admit the lesser line *AB*, to be 100, & cutting off *CE* equal to *AB*, I find the proportion of *CE*, vnto *ED*, to be as 100, almost to 67; wherefore this way y^e proportiō of *AB* vnto *CD*, is as 100 vnto almost 167.

This proposition may also not vnfitly be wrought by any other number, that admits seuerall diuisions, and namely, by the numbers of 60. And so the lesser line will be found to be 36, which is as before in lesser numbers, as 3 vnto 5. It may also be wrought without opening the *Sector*. For if the lines between which we seek a proportion, be applyed to the lines of *Lines*, (or any other Scale of equall parts) there will be such proportiō found between them, as betweene the lines to which they are equall.

6 *Two lines being giuen to finde a third
incontinuell proportion.*

First place both the lines giuen, on both sides of the *Sector* from the Center, and marke the termes of their extension, then take out the second line againe, and to it open the *Sector*, in the termes of the first line, so keeping the

the *Sector* at this angle, the parallell distance betweene the termes of the second line, shall be the third proportionall.

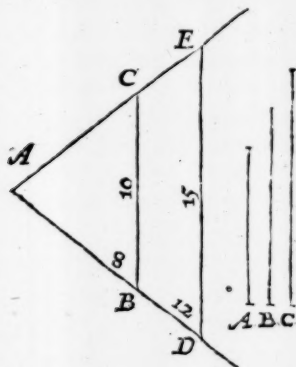


Let the two lines giuen be AB, AC , which I take out and place on both sides of the *Sector*, so as they all meete in the center A , let the termes of the first line be B and B , the termes of the second C and C . Then doe I take out AC the second line againe, and to it open the *Sector* in the termes BB . So the parallell betweene C and C doth giue me the third line in continuall proportion. For as AB is vnto AC , so BB , equall to AC , is vnto CC .

7 Three lines being given to finde the fourth
in discontinuall proportion.

Here the first line and the third are to be placed on both sides of the *Sector* from the center, then take out the second line, and to it open the *Sector* in the termes of the first line. For so keeping the *Sector* at this angle, the parallell distance between the termes of the third line, shall be the fourth proportionall.

Let the three lines given be A, B, C.



First I take out A and C, and place them on both sides of the *Sector*, in A B, A C, and A D, A E, laying the beginning of both lines at the center A, then do I take out B the second line, according to it I open the *Sector* in B and C, the termes of the first line: so the parallell distance between D and E, doth giue me the fourth proportionall which was required.

As in *Arithmetique*, it sufficeth if the first and third number giuen be of one denomination, the second & the fourth which is required be of another. For one and the same denomination is not required necessarily in them all. So in *Geometrie*, it sufficeth if the sides A B, A D, resembling the first

first and third lines giuen be measured in one Scale, and the parallells B C, D E be measured in another. Wherefore knowing the proportion of A the first line, and C the third line, by the first *prop.* before. Which is here as 8 to 12, & descending in lesser numbers is as 4 to 6, or as 2 to 3, or ascending in greater numbers, as 16 vnto 24, or 18 to 27, or 20 to 30, or 30 to 45, or 40 to 60 &c. If the *Sector* be opened in the points of 8 and 8, to the quantity of B, the second line giuen, then a parallell betweene 12 and 12, shall giue D E, the fourth line required. So likewise if it be opened in 4 and 4, then a parallell betweene 6 and 6, or if in 16 and 16, then a parallell betweene 24 and 24 shall giue the same D E. And so in the rest.

8 *To diuide a line in such sort as another line is before diuided.*

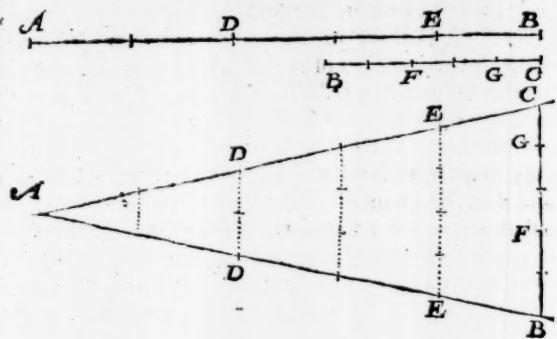
First take out the line giuen, which is already diuided, and laying it on both sides of the *Sector* from the center, marke how farre it extendeth. Then take out the second line which is to be diuided, and to it open the *Sector* in the termes of the first line. This done, take out the parts of the first line, and place them also on the same side of the *Sector* from the center. For the parallells taken in the termes of these parts, shalbe the correspondent parts in the line which is to be diuided.

Let *AB*, be a line diuided in *D* and *E*, and *BC*, the line which I am to diuide in such sort, as *AB* is diuided.

First I take out the line *AB*, and place it on the line of *Lines* in *AB*, *AC*, both from the center *A*, then take I out the second *BC*, and to it open the *Sector* in *B* and *C*, the termes of the first line. The *Sector* thus opened to his due angle, I take out *AD* and *AE*, the parts of the first line *AB*, and place them also on both the sides of the *Sector* in *AD*, *AE*, so the parallell *DD*, giueth me *BF*, and the parallell *E*, giueth me *BG*, and now the line *BC*, is diuided in *F* & *G*, as is the other line *AB*, in *D* and *E*, which is that which was

D 2

requi-



required.

If the line *AB*, were longer then one of the sides of the Ruler, then should I finde what proportion it hath to his parts *AD*, *AE*, and that knowne I may worke as before in the former proposition.

*9 Two numbers being given to finde a third
in continuall proportion.*

First reckon the two numbers giuen on both sides of the lines of *Lines* from the center, and marke the termes to which either of them extendeth, then take out a line resembling the second number againe, and to it open the *Sector* in the termes of the first number, forso keeping the *Sector* at this angle, the parallell distance betweene the termes of the second laterall number, being measured in the same Scale, from whence his parallell was taken, shall giue the third number proportionall.

Let the two numbers giuen be 18, 24, these being resembled in lines, the worke will be in a manner all one, with that in the sixt *Prop.* and so the third proportionall number will be found to be 32.

10. Three numbers being given to finde a fourth
in discontinuall proportion.

THe solution of this propolition, is in a manner all one with that before in the seventh *Prop.* onely there may be some difficulty in placing of the numbers. To avoyd this, we must remember that three numbers being given, the question is annexed but to one, and this must allwayes be placed in the third place, that which agrees with this third number in denomination, shalbe the first number, and that which remaineth the second number. This being considered, reckon the first, and third numbers, which are of the first denomination on both sides of the lines of *Lines* from the center, and marke the termes to which either of them extendeth, then take out a line resembling the second number, and to it open the *Sector* in the termes of the first number, for so keeping the *Sector* at this angle, the parallell distance between the termes of the third laterall number, being measured in the same Scale from whence his parallell was taken, shall giue the fourth number proportionall.

As if a question were proposed in this manner, 10 yards cost 8 £, how many yards may we buy for 12 £? heere the question is annexed to 12; and therefore it shall be the third number, and because 8 is of the same denomination, it shall be the first number, then 10 remaining, it must be the second number, so will they stand in this order, 8, 10, 12. These being resembled in lines, the worke will be in a manner the same, with that in the seventh *Prop.* and the fourth proportionall number will be found to be 15. For as 8 are to 10, so 12 vnto 15.

And this holdeth in direct proportion, where, as the first number is to the second, so the third to the fourth. So that if the third number be greater then the first, the fourth will be greater then the second, or if the third number be lesse then the first, the fourth will be lesse then the second, but in *reciprocall* proportion, commonly called the *Back rule*,

where by how much the first number is greater then the third, so much the second will be lesse then the fourth, or by how much the first number is lesse then the third, so much the second will be greater then the fourth. The manner of working must be contrary, that is; the *Sector* is to be opened in the termes of the third number, and the parallell resembling the number required, is to be found betweene the termes of the first number, the rest may be obserued as before, as for example.

If twelve men would raise a frame in ten dayes, in how many dayes would eight men raise the same frame? Here, because the fewer men would require the longer time, though the numbers be 12, 10, 8, yet the fourth proportionall will be found to be 15.

75.

50.

So if 60 yards, of three quarters of a yard in bredth, would hang round about a roome, and it were required to know how many yards of halfe a yard in bredth, would serue for the same roome. The fourth proportionall would be found to be 90.

So if to make a foote superficiall, 12 inches in bredth doe require 12 inches in length, and the bredth being 16 inches, it were required to knowe the length. Here, because the more bredth, the lesse length, the fourth proportionall will be found to be 9.

So if to make a Solid foote, a base of 144 inches, require 12 inches in height, and a base given being 216 inches, it were required to knowe how many inches it shall haue in height. The fourth proportionall would be found to be 8.

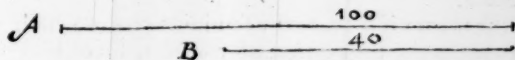
This last proposition of finding a fourth proportionall number, may be wrought also by the lines of Superficies, and by the lines of Solids.

CHAP, III.

The use of the lines of Superficies.

I To finde a proportion betweene two or more like Superficies.

TAKE one of the sides of the greater *Superficies* giuen, and according to it open the *Sector* in the points of 100 and 100, in the lines of *Superficies*, then take the like sides of the lesser *Superficies* seuerally, and carry them parallell to the former, till they stay in like points, so the number of points wherein they stay, shall shew their proportion vnto 100.

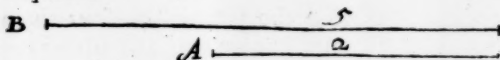


Let *A* and *B*, be the sides of like *Superficies*, as the sides of two squares, or the diameters of two circles, first I take the side *A*, and to it open the *Sector* in the points of 100, then keeping the *Sector* to this angle, I enter the lesser side *B*, parallel to the former, and finde it to crosse the lines of *Superficies* in the points of 40, wherefore the proportion of the *Superficies*, whose side is *A*, to that whose side is *B*, is as 100 vnto 40, which is in lesser numbers, as 5 vnto 2.

This proposition might haue beene wrought by 60, or any other number that admits seuerall diuisions. It may also be wrought without opening the *Sector*, for if the sides of the *Superficies* giuen, be applied to the lines of *Superficies* beginning alwayes at the center of the *Sector*, there will be such proportion found betweene them, as betweene the number of parts whereon they fall.

- 2 To augment a Superficies in a given Proportion.
- 3 To diminish a Superficies in a given Proportion.

Take the side of the *Superficies*, and to it open the *Sector* in the points of the numbers given; then keeping the *Sector* at that angle, the parallell distance between the points of the number required, shall giue the like side of the *Superficies* required.



Let *A* be the side of a Square to be augmented in the proportion of 2 to 5. First I take the side *A*, and put it ouer in the lines of *Superficies*, in 2 and 2; so the parallell between 5 and 5, doth giue me the side *B*, on which if I should make a Square, it would haue such proportion to the square of *A*, as 5 vnto 2.

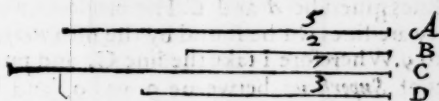
In like maner if *B* were the semidiameter of a circle to be diminished in the proportion of 5 vnto 2, I would take out *B*, and put it ouer in the lines of *Superficies*, in 5 and 5; so the parallell between 2 and 2, would giue me *A*; on which Semidiameter if I should make a circle, it would be lesse then the circle made vpon the Semidiameter *B*, in such proportion as 2 is lesse then 5.

For varietie of worke the like caution may be here obserued to that which we gaue in the third *Prop.* of *Lines*.

- 4 To adde one like Superficies to another.
- 5 To subtract one like Superficies from another.

First, the proportion betweene like sides of the *Superficies* given, is to be found by the first *Prop.* of *Superficies*, then adde or subtract the numbers of those proportions, and accordingly augment or diminish by the former *Prop.*

As

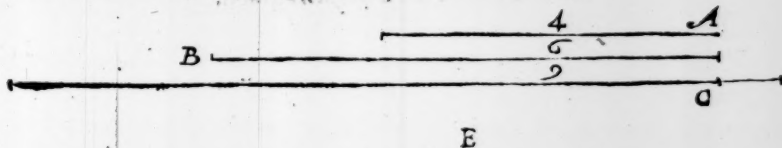


As if *A* and *B* were the sides of two Squares, and it were required to make a third Square equall to them both. First the proportion between the squares of *A* and *B*, would be found to be as 100 vnto 40, or in the lesser numbers as 5 to 2; then because 5 and 2 added do make 7, I augment the side *A* in the proportion of 5 to 7, and produce the side *C*, on which if I make a square, it will be equall to both the squares of *A* and *B*, which was required.

In like maner *A* and *B* being the sides of two Squares, if it were required to subtract the square of *B* out of the square of *A*, and to make a square equall to the remainder, here the proportion being as 5 to 2, because 2 taken out of 5, the remainder is 3, I would diminish the side *A* in the proportion of 5 to 3, and so I should produce the side *D*, on which if I make a square, it will be equall to the remainder when the square of *B* is taken out of the square of *A*, that is, the two squares made vpon *B* and *D*, shall be equall to the first square made vpon the side *A*.

¶ To find a meane proportionall betweene two lines giuen.

First find what proportion is betweene the lines giuen, as they are lines, by the fifth *Prop.* of *Lines*, then open the *Scale* in the lines of *Superficies*, according to his number, to the quantitie of the one, and a parallell taken betweene the points of the number belonging to the other line shall be the meane proportionall.



Let the lines giuen be *A* and *C*. The proportion between them as they are lines wil be found by the fifth *Prop. of Lines* to be as 4 to 9. Wherefore I take the line *C*, and put it ouer in the lines of *Superficies* betweene 9 and 9, and keeping the *Sector* at this angle, his parallell betweene 4 and 4 doth giue me *B* for the meane proportionall. Then for prooffe of the operation I may take this line *B*, and put it ouer between 9 and 9: so his parallel between 4 and 4, shall giue me the first line *A*. Whereby it is plaine that these three lines do hold in continuall proportion; and therefore *B* is a meane proportionall betweene *A* and *C* the extremes giuen.

Vpon the finding out of this meane proportion depend many Corollaries, as

To make a Square equall to a Superficies giuen.

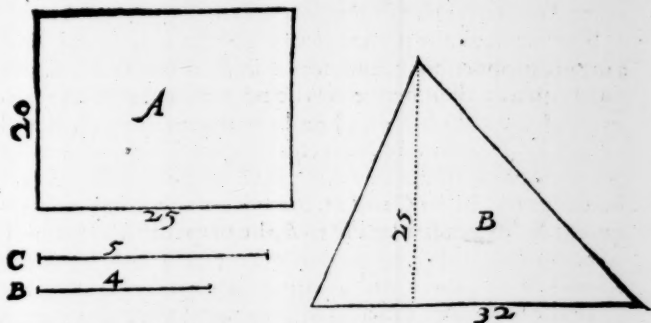
IF the *Superficies* giuen be a rectangle parallelogram, a meane proportionall betweene the two vnequal sides shall be the side of his equall square.

If it shall be a triangle, a meane proportion betweene the perpendicular and halfe the base shal be the side of his equal square. If it shall be any other right-lined figure, it may be resolued into triangles, and so a side of a square found equall to euery triangle; and these being reduced into one equall square, it shall be equall to the whole right-lined figure giuen.

To finde a proportion betweene Superficies, though they be unlike one to the other.

IF to euery *Superficies* we find the side of his equall square, the proportion betweene these squares, shall be the proportion betweene the *Superficies* giuen.

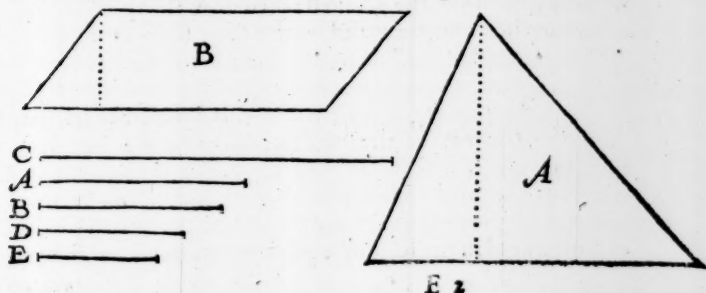
Let



Let the *Superficies* given, be the oblonge *A*, and the triangle *B*. First between the vnequall sides of *A*, I find a meane proportionall, and note it in *C*: this is the side of a square equall vnto *A*. Then betweene the perpendicular of *B*, and halfe his base, I finde a meane proportionall, and note it in *B*: this is the side of a Square equall to *B*: but the proportion between the squares of *C* and *B*, will be found by the first *Prop.* of *Superficies* to be as 5 to 4: and therefore this is the proportion betwene those giuen *Superficies*.

To make a *Superficies* like to one *Superficies*
and equall to another.

Let the one *Superficies* giuen be the triangle *A*, and the other the *Rhomboides* *B*; and let it be required to make an



other *Rhomboides* like to *B*, and equall to the triangle *A*.

First between the perpendicular and the base of *B*, I find a meane proportionall, and note it in *B*, as the side of his equall square: then betweene the perpendicular of the triangle *A*, and halfe his base, I find a meane proportionall, and note it in *A*, as the side of his equall square. Wherefore now as the side *B* is to the side *A*, so shall the sides of the *Rhomboides* giuen be to *C* and *D*, the sides of the *Rhomboides* required, & his perpendicular also to *E*, the perpendicular required.

perpend =

Having the sides and the perpendicular, I may frame the *Rhomboides* vp, and it will be equall to the triangle *A*.

If the *Superficies* giuen had been any other right-lined figures, they might haue been resolued into triangles, and then brought into squares as before.

Many such Corollaries might haue been annexed, but the meanes of finding a meane proportionall being knowne, they all follow of themselves.

7 To finde a meane proportionall betweene two numbers giuen.

First reckon the two numbers giuen on both sides of the Lines of *Superficies*, from the center, and mark the termes whereunto they extend; then take a line out of the Line of *Lines*, or any other scale of equall parts resembling one of those numbers giuen, and put it ouer in the termes of his like number in the lines of *Superficies*; for so keeping the *Sector* at this angle, the parallell taken from the termes of the other number and measured in the same scale from which the other parallell was taken, shall here shew the meane proportionall which was required.

Let the numbers giuen be 4 and 9. If I shall take the line *A*, in the *Diagram* of the sixt *Prop.* resembling 4 in a scale of equall parts, and to it open the *Sector* in the termes of 4 and 9, in the lines of *Superficies*, his parallell betweene 9 and 9 doth giue me *B* for the meane proportionall. And this measured in the scale of equall parts doth extend to 6, which

which is the meane proportionall number between 4 and 9.

For as 4 to 6, so 6 to 9.

In like maner if I take the line *C*, resembling 9 in a scale of equall parts, and to it open the *Sector* in the termes of 9 and 9, in the lines of *Superficies*, his parallell between 4 and 4 doth giue me the same line *B*, which will proue to be 6, as before, if it be measured in the same scale whence *D* was taken.

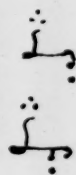
8 To find the square roote of a number.

9 The roote being giuen to find the square number of that roote.

IN the extraction of a square roote it is vsuall to set prickes vnder the first figure, the third, the fifth, the seuenth, and so forward, beginning from the right hand toward the left, and as many prickes as fall to be vnder the square number giuen, so many figures shall be in the roote: so that if the number giuen be lesse then 100, the roote shall be only of one figure; if lesse then 10000, it shall be but two figures; if lesse then 1000000, it shall be three figures, &c.

Thereupon the lines of *Superficies* are diuided first into an hundred parts, and if the number giuen be greater then 100, the first diuision (which before did signifie only one) must signifie 100, and the whole line shall be 10000 parts: if yet the number giuen be greater then 10000, the first diuision must now signifie 10000, and the whole line be esteemed at 1000000 parts: and if this be too little to expresse the number giuen, as oft as we haue recourse to the beginning, the whole line shall increase it selfe an hundred times.

By this meanes if the last prick to the left hand shall fall vnder the last figure, which will be as oft as there be odde figures, the number giuen shall fall out betweene the center of the *Sector* and the tenth diuision: but if the last prick shall fall vnder the last figure but one, which will be as oft as there be euen figures, then the number giuen shall fall out betweene the tenth diuision and the end of the *Sector*.



This being considered, when a number is given and the square roote is required, take a paire of compasses and setting one foote in the center, extend the other to the terme of the number given in one of the lines of *Superficies*; for this distance applied to one of the Lines of *Lines*, shall shew what the Square roote is, without opening the *Sector*.

Thus 64 doth giue a roote of 8, and 862 a roote of almost 29, and 1296 a roote of 36, and 7056 a roote of 84, and 62500 a roote of 250, and 714000 a roote of about 845, and so in the rest. 360

On the contrary, a number given may be squared, if first we extend the compasses to the number given in the lines of *Lines*, and then apply the distance to the *Lines* of *Superficies*, as may appeare by the former examples.

10 *Three numbers being given to find the fourth in a duplicated proportion.*

IT is plaine by the 19 and 20 *Prop. 6. Lib. of Euclid*, that like *Superficies* do hold in a duplicated proportion of their homologall sides, whereupon a question being moued concerning *Superficies* and their sides. It is visuall in Arithmetick that the proportion be first duplicated before the question be resolved, which is not necessarie in the vse of the *Sector*, only the numbers which do signifie *Superficies* must be reckoned in the lines of *Superficies*, and they which signifie the sides of *Superficies*, in the lines of *Lines*, after this maner.

If a question be made concerning a *Superficies*, the two numbers of the first denomination must be reckoned in the lines of *Lines*, and the *Sector* opened in the termes of the first number to the quantitie of a line out of the scale of *Superficies* resembling the second number; so his parallels taken betweene the termes of the third number, being measured in the same scale of *Superficies*, shall giue the Superficiall number which was required.

As if a Square, whose side is fortie perches in length, shall con-

containe ten acres in the *Superficies*, and it be required to know how many acres the Square should contain, whose side is sixtie perches.

Here if I tooke 10 out of the line of *Superficies*, and put it ouer in 40 in the lines of *Lines*, his parallell between 60 and 60 measured in the line of *Superficies*, would be $22\frac{1}{2}$; and such is the number of acres required. For Squares do hold in a duplicated proportion of their sides; wherefore when the proportion of their sides is as 4 to 6, and 4 multiplied into 4 become 16, and 6 multiplied into 6 become 36, the proportion of their squares shall be as 16 to 36; and such is the proportion of 10 to $22\frac{1}{2}$.

If a field measured with a statute perch of $16\frac{1}{2}$ foote, shall containe 288 acres, and it be required to know how many acres it would containe if it were measured with a woodland perch of 18 foote.

Here because the proportion is reciprocall, if I tooke 288 out of the line of *Superficies*, and put it ouer in 18, in the lines of *Lines*, his parallell between $16\frac{1}{2}$ and $16\frac{1}{2}$ measured in the line of *Superficies*, would be 242; and such is the number of acres required.

For seeing the proportion of the sides is as $16\frac{1}{2}$ to 18, or in lesser numbers as 11 to 12, and that 11 multiplied into 11 become 121, and 12 into 12 become 144, the proportion of these *Superficies* shall be as a 121 to 144, and so haue 288 to 242, in reciprocall proportion.

On the contrary, if a question be proposed concerning the side of a *Superficies*, the two numbers of the first denomination must be reckoned in the lines of *Superficies*, and the *Scissor* opened in the termes of the first number, to the quantitie of a line, out of the line of *Lines*, or some Scale of equall parts, resembling the second number; so his parallell taken between the termes of the third number being measured in the same scale with the second number, shal giue the fourth number required.

As if a field contained 288 acres when it was measured with a statute perch of $16\frac{1}{2}$, and being measured with another

ther perch, was found to containe 242 acres, it were required to know what was the length of the perch with which it was so measured.

Here because the proportion is reciprocal, if I tooke $16\frac{1}{2}$ out of the line of *Lines*, and put it over in 242 in the lines of *Superficies*, his parallell betweene 288 and 288, being measured in the line of *Lines*, would be 18, and such is the length of the perch in foote whereby the field was last measured.

For seeing the proportion of the acres is as 288 vnto 242, or in the least numbers as 144 to 121, and that the roote of 244 is 12, and the root of 121 is 11, the proportion of roots and consequently of the perches shall be as 12 to 11, and so are $16\frac{1}{2}$ to 18, in *reciprocall* proportion.

If 360 men were to be set in forme of a long square, whose sides shall haue the proportion of 5 to 8; and it were required to know the number of men to be placed in front and file: if the sides were onely 5 and 8, there should be but 40 men; but there are 360: therefore working as before, I find that

As 40 to the square of 5,
so 360 to the square of 15.

As 40 to the square of 8,
so 360 to the square of 24.

and so 15 and 24 are the sides required.

If 1000 men were lodged in a square ground, whose side were 60 paces, and it were required to know the side of the square wherein 5000 might be so lodged, here working as before, I should find that

As 1000 are to the square of 60:
so 5000 to the square of 134.

And such very neare is the number of paces required.

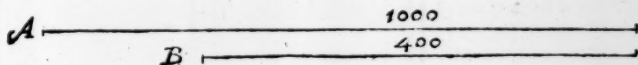
CHAP. IV.

The use of the lines of Solids.

To finde a proportion betweene two or more like Solids:

IN the Sphere, in regular, parallell, and other like bodies, whose sides next the equall angles are proportionall, the worke is in a manner the same, with that in the first *Prop.* of *Superficies*, but that it is wrought on other lines.

Take one of the sides of the greater *Solid*, & according to it open the *Sector* in the points of 1000 and 1000, in the lines of *Solids*, then take the like sides of the lesser *Solids* severally, and carry them parallell to the former, till they stay in like points, so the number of points wherein they stay, shall shew their proportion to 1000.



Let *A* and *B*, be the like sides of like Solids, either the diameters, or semidiameters of two spheres, or the sides of two cubes, or other like. First I take the side *A*, and to it open the *Sector* in the points of 1000, then keeping the *Sector* at this angle, I enter the lesser side *B*, parallell to the former, and finde it to crosse the line of *Solids* in the points of 400, and such is the proportion betweene the Solids required, which in lesser number is as 5 to 2.

This proposition might haue beene wrought by 60, or any other number that admits severall diuisions.

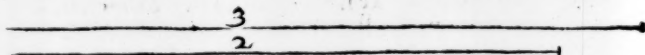
It may also be wrought without opening the *Sector*, for if the sides of the Solids giuen, be applied to the lines of *Solids*, beginning allwayes at the center of the *Sector*, there will be such proportion betweene them, as betweene the numbers of parts whereon they fall.

2 *To augment a Solid in a given proportion.*

3 *To diminish a Solid in a given proportion.*

TAKE the side of the Solid giuen, and to it open the *Sector*, in the points of the number giuen: then keeping the *Sector* at that angle, the parallell distance bet weene the points of the number required, shall giue the like side of the Solid required.

If it be a *parallelopipedon*, or some irregular Solid, the other like sides may be found out in the same manner, and with them the Solids required, may be made vp with the same angles.



Let *A* be the side of a cube, to be augmented in the proportion of 2 to 3. First I take the side *A*, and put it ouer in the lines of *Solids* in 2 and 2, so the parallell betweene 3 and 3, doth giue me the side *B*, on which if I make a cube, it will haue such proportion to the cube of *A*, as 3 to 2.

In like manner, if *B* were the diameter of a Sphere, to be diminished in the proportion of 3 to 2. I would take out *B*, and put it ouer in the lines of *Solids*, in 3 and 3, so the parallell betweene 2 and 2, would giue me *A*: to which diameter if I should make a Sphere, it would be lesse then the Sphere, whose diameter is *B*, in such proportion as 2 is lesse then 3.

Here also for variety of worke, may the like caution be obserued to that which we gaue in the third *Prop. of Lines*.

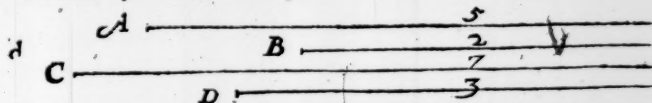
4 *To adde one like Solid to another.*

5 *To subtract one like Solid from another.*

FIRST the proportion betweene the sides of the like Solids giuen, is to be found by the first *Prop. of Solids*: then adde

or

or subtract those proportions, and accordingly augment or diminish by the former *Prop.*



As if *A* and *B* were the sides of two cubes, and it were required to make a third cube equall to them both : first the proportion betweene the sides *A* and *B*, would be found to be as 100 to 40, or in lesser termes as 5 to 2. Then because 5 and 2 being added do make 7, I augment the side *A* in the proportion of 5 to 7, and produce the side *C*, on which if I make a cube, it will be equall to both the cubes of *A* and *B*, which was required.

In like maner *A* and *B* being the sides of two cubes, if it were required to subtract the cube of *B* out of the cube of *A*, and to make a cube equal to the remainder. Here the proportion being as 5 to 2, because 2 taken out of 5, the remainder is 3, I should diminish the side *A* in the proportion of 5 to 3, and so I should haue the side *D*, on which if I make a cube, it will be equall to the remainder when the cube of *B* is taken out of the cube of *A*, that is the two cubes made vpon *B* and *D*, shall be equall to the first cube made vpon the side *A*.

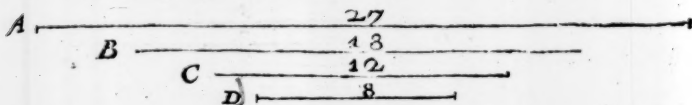
6 To find two meane proportionall lines betweene two extreme lines giuen.

First I find what proportion is betweene the two extreme lines giuen as they are lines, by the fifth *Prop.* of *Lines*, then open the *Setſor* in the lines of *Solids*, to the quantitie of the former extreme, and a parallell betweene the points of the number belonging to the other extreme, shall be that meane proportionall which is next the former extreme. This done, open the *Setſor* againe to this meane proportionall in the points of the former extreme, and the parallell distance

F 2

be-

betweene the points of the latter extreme, shall be the other meane proportionall required.



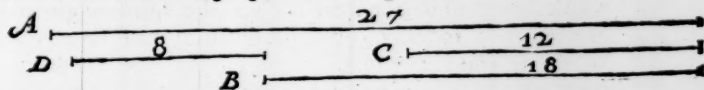
Let the two extreme lines giuen be A and D, the proportion betweene them, as they are lines, will be found to be as 27 to 8. Wherefore I take the line A, and put it ouer in the lines of *Solids* betweene 27 and 27, and keeping the *Settor* at this angle, his parallell betweene 8 and 8, doth giue me B, the meane proportionall next vnto A. Then put I ouer this line B, betweene the aforesaid 27 and 27, and his parallell betweene 8 and 8 doth giue me the line C, the other meane proportionall which was required.

Againe, for proofof the operation I put ouer this line C in the aforesaid 27 and 27, and his parallell between 8 and 8 doth giue me the very line D: whereby it is plaine that these foure lines do hold in continuall proportion; and so B and C are found to be the meane proportionals betweene A and D the extremes giuen.

7 To find two meane proportionall numbers between two extreme numbers giuen.

First reckon the numbers giuen on both sides of the lines of *Solids*, beginning from the center, and marking the termes whereto they extend: then take a line out of the line of *Lines*, or any other scale of equall parts resembling the former of those numbers, and put it ouer in the lines of *Solids*, betweene the points of his like number, and a parallell betweene the points belonging to the other extreme, measured in the scale from whence the other parallell was taken, shall giue that meane proportionall number which is next the former extreme. This done, open the *Settor* againe to this meane proportionall in the points of the former extreme, and

and the parallell distance betweene the points of the latter extreame, measured in the same scale as before, shall there shew the other meane proportionall required.



Let the two extreame numbers giuen be 27 and 8; if I shall take the line A, resembling 27 in a scale of equall parts, and to it open the *Sector* in 27 and 27, in the line of *Solids*, his parallell betweene 8 and 8 doth giue me B for his next meane proportionall, and this measured in the former scale doth extend to 18. Then put I ouer this line B between the aforesaid 27 and 27, and his parallell between 8 and 8 doth giue me C for the other meane proportionall, and this measured in the former scale doth extend to 12. Againe, for prooffe of my worke, I put ouer this line C between 27 and 27, as before, and his parallell betweene 8 and 8 doth giue me D, which measured in the former scale doth extend to 8, which was the latter extreame number giuen; whereby it is plaine that these foure numbers do hold in continuall proportion: and therefore 18 and 12 are meane proportionals betweene 27 and 8, which was required.

8 To finde the cubique roote of a number.

9 The roote being giuen to finde the cube number of that roote.

231. solid Inches
in a gallon.

IN the extraction of a cubique root, it is vsuall to set pricks vnder the first figure, the fourth, the seuenth, the tenth, and so forward, omitting two, and pricking the third from the righthand toward the left; and as many pricks as fall to be vnder the cubique number, so many figures shall be in the roote. So that if the number giuen be lesse then 1000, the roote shall be only of one figure; if lesse then 1000000, it shall be but of two figures; if aboue these, and lesse then 1000000000, it shall be but three figures; &c. whereupon

the lines of *Solids* are diuided, first into 1000 parts, and if the numbers giuen be greater the 1000, the first diuision (which before did signifie onely one) must signifie 1000, and the whole line shall be 1000000 : if yet the number giuen be greater then 1000000, the first diuision must now signifie 1000000, and the whole line be esteemed at 1000000000 parts, and if these be to little to expresse the numbers giuen, as oft as wee haue recourse to the begining, the whole line shall encrease it selfe a thousand times..

By these meanes, if the last pricke, to the left hand, shall fall vnder the last figure, the number giuen shall be reckoned at the beginning of the lines of *Solids*, from 1 to 10, and the first figure of the roote shall be alwayes either 1, or 2. If the last pricke shall fall vnder the last figure but one, then the number giuen shall be reckoned in the middle of the line of *Solids*, betweene 10 and 100, and the first figure of the roote shall be alwayes either 2, or 3, or 4. But if the last pricke shall fall vnder the last figure but two, then the number giuen, shall be reckoned at the end of the line of *Solids*, betweene 100, and 1000.

This being considered when a number is giuen, and the cubique roote required : Set one foote of the compalles in the center of the *Sector*, extend the other in the line of *Solids*, to the points of the number giuen : for this distance applied to one of the line of *Lines*, shall shew what the cubique roote is, without opening the *Sector*.

So the nearest roote of 8490000, is about 204.

The nearest roote of 84900000, is about 439.

The nearest roote of 849000000, is about 947.

On the contrary, a number may be cubed, if first we extend the compalles to the number giuen, in the line of *Lines*, and then apply the distance to the lines of *Solids*; as may appear by the former examples.

10 Three numbers being given to finde a fourth in a triplicated proportion.

AS like *Superficies* do hold in a duplicated proportion, so like solids in a triplicated proportion of their homologall sides: and therefore the same worke is to be observed here on the lines of *Solids*, as before in the lines of *Superficies*; as may appear by these two examples.

If a cube whose side is 4 inches, shall be 7 pound weight, and it be required to know the weight of a cube whose side is 7 inches; here the proportion would be,

As 4 are to a cube of 7:
so 7 to a cube of $37\frac{1}{8}$.

And if I tooke 7 out of the lines of *Solids*, and put it over in 4 and 4, in the lines of *Lines*, his parallell between 7 and 7 measured in the lines of *Solids*, would be $37\frac{1}{8}$; and such is the weight required.

If a bullet of 27 pound weight haue a diameter of 6 inches, and it be required to know the diameter of the like bullet, whose weight is 125 pounds; here the proportion would be,

As the cubique root of 27 is vnto 6:
so the cubique root of 125 is vnto 10.

And if I tooke 6 out of the line of *Lines*, and put it over in 27 and 27 of the lines of *Solids*, his parallell betweene 125 and 125 measured in the line of *Lines*, would be 10; and such is the length of the diameter required.

The end of the first booke.

THE HISTORY OF THE

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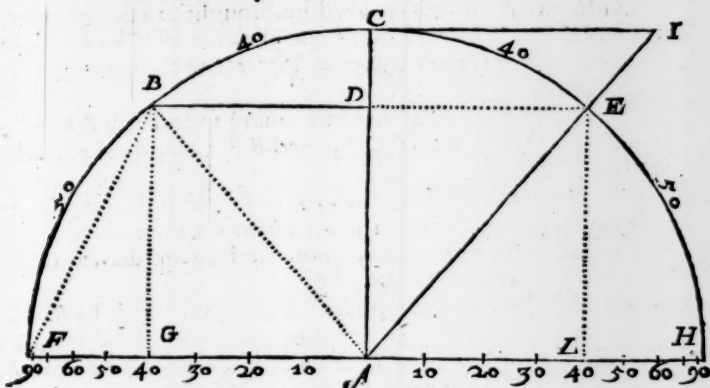
THE SECOND BOOKE OF THE SECTOR,

Containing the vse of the Circular
Lines.

CHAP. I.

*Of the nature of Sines, Chords, Tangents and
Secants, fit to be knowne before hand
in reference to right-line Triangles.*

IN the *Canon of Triangles*, a circle is commonly di-
vided into 360 *degrees*, each *degree* into 60 *minutes*,
each *minute* into 60 *seconds*.



A semicircle therefore is an arke of 180 gr.

G

A

A quadrant is an arke of 90 gr.

The measure of an angle is the arke of a circle, described out of the angular point, intercepted betwene the sides sufficiently produced.

So the measure of a right angle is alwayes an arke of 90 gr. and in this example the measure of the angle BAD is the arke BC of 40 gr; the measure of the angle BAG , is the arke BF of 50 gr.

The complement of an arke or of an angle doth commonly signifie that arke which the giuen arke doth want of 90 gr: and so the arke BF is the cōplement of the arke BC ; & the angle BAF , whose measure is BF , is the complement of the angle BAC ; and on the contrary.

The complement of an arke or angle in regard of a semi-circle, is that arke which the giuen arke wanteth to make vp 180 gr: and so the angle EAH is the complement of the angle EAF , as the arke EH is the complement of the arke FE , in which the arke CE is the excess above the quadrant.

The proportions which these arkcs (being the measures of angles) haue to the sides of a triangle, cannot be certaine, vnlesse that which is crooked be brought to a straight line; and that may be done by the application of *Chords*, *Right Sines*, *versed Sines*, *Tangents* and *Secants*, to the semidiameter of a circle.

A *Chorde* is a right line subtending an arke: so BE is the chorde of the arke BCE , and BF a chorde of the arke BF .

A *right Sine* is halfe the chorde of the double arke, viz. the right line which falleth perpendicularly from the one extreme of the giuen arke, vpon the diameter drawne to the other extreme of the said arke.

So if the giuen arke be BC , or the giuen angle be BAC , let the diameter be drawne through the center A vnto C ; and a perpendicular BD be let downe from the extreme B , vpon AC ; this perpendicular BD shall be the *right sine* both of the arke BC , and also of the angle BAC : and it is also

also the halfe of the chord B E, subtending the arke B C E, which is double to the giuen arke B C. In like maner, the semidiameter F A, is the *right sine* of the arke F C, and of the right angle F A C; for it falleth perpendicularly vpon A C, and it is the halfe of the chord F H.

This whole Sine of 90 gr. is hereafter called *Radius*; but the other *Sines* take their denomination from the degrees and minutes of their arks.

Sinus versus, the *versed sine* is a segment of the diameter, intercepted betweene the *right sine* of the same arke, and the circumference of the circle. So D C is the *versed sine* of the arke C B, and G F the *versed sine* of the arke B F, and G H the *versed sine* of the arke B H.

A *Tangent* is a right line perpendicular to the diameter, drawne by the one extreme of the giuen arke, and terminated by the *secant* drawne from the center through the other extreme of the said arke.

A *Secant* is a right line drawne from the center, through one extreme of the giuen arke, till it meete with the *tangent* raised from the diameter at the other extreme of the said arke.

So if the giuen arke be C E, or the giuen angle be C A E, let the diameter be drawne through the center A to C, and in C to A C, be raised a perpendicular C I. Then let another line be drawne from the center A through E, till it meet with the perpendicular C I in I; the line C I is a *Tangent*, and A I is the *Secant* both of the arke C E, and of the angle C A E.

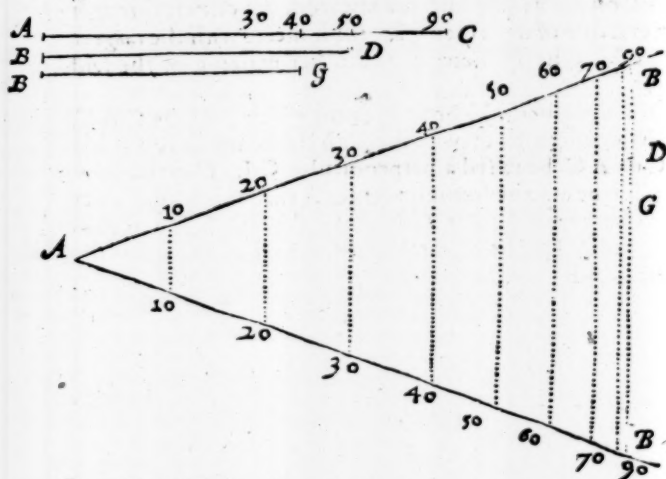
CHAP. II.

Of the generall use of Sines and Tangents.

- 1 The Radius being knowne to find the right sine of any arke or angle.

IF the Radius of the circle giuen be equall to the laterall Radius, that is, to the whole line of *Sines* on the *Sector*, there needs no farther worke, but to take the other sines also out of the side of the *Sector*. But if it be either greater or lesser, then let it be made a parallell Radius, by applying it ouer in the lines of *Sines*, betweene 90 and 90; so the parallell taken from the like laterall sines, shall be the *sine* required.

As if the giuen Radius be AC , and it were required to find the sine of 50 *Gr.* & his complement agreeable to that radius,



Let AB , AB represent the lines of *sines* on the *Sector*, and let BB , the distance betweene 90 and 90, be equall to the giuen

giuen radius *AC*. Here the lines *A 40*, *A 50*, *A 90*, may be called the *laterall sines* of 40, 50, & 90; in regard of their place on the sides of the *Sector*. The lines betweene 40 and 40, betweene 50 and 50, betweene 90 and 90, may be called the *parallell sines* of 40, 50, and 90; in regard they are parallell one to the other. The whole line of 90 *Gr.* here standing for the semidiameter of the circle, may be called the Radius. And therefore if *AC* be put ouer in the line of *Sines* in 90 and 90, and so made a *parallell radius*, his parallell sine betweene 50 and 50, shall be *BD*, the sine of 50 required. And because 50 taken out of 90, the complement is 40; his *parallell sine* betweene 40 and 40, shall be *BG*, the sine of the complement which was required.

2 *The right sine of any arke being giuen
to finde the Radius.*

Turne the sine giuen into a parallell sine, and his parallell *Radius* shall be the *Radius* required.

As if *BD* were the giuen sine of 50 *Gr.* and it were required to finde the Radius: let *BD* be made a parallell sine of 50 *Gr.* by applying it ouer in the lines of *Sines*, betweene 50 and 50; so his parallell Radius betweene 90 and 90 shall be *AC*, the Radius required.

3 *The Radius of a circle, or the right Sine of any arke
being giuen, and a streight line resembling a Sine,
to find the quantitie of that unknowne Sine.*

Let the Radius or right sine giuen be turned into his parallell; then take the right line giuen, and carrie it parallell to the former, till it stay in like *Sines*: so the number of degrees and minutes where it stayeth, shall giue the quantitie of the Sine required.

As if *BD* were the giuen sine of 50 *Gr.* and *BG* the streight line giuen: first I make *BD* a parallell sine of 50 *Gr.*; then keeping the *Sector* at this angle, I carie the line *BG*

parallell, and find it to stay in no other but 40 and 40; and therefore 40 gr. is his quantitie required.

4 The Radius or any right Sine being giuen, to finde the versed sine of any arke.

IF the arke, whose *versed sine* is required, be lesse then the Quadrant, take the sine of the complement out of the radius; and the remainder shall be the *sinus versus*, the versed sine of that arke.

As if A B being the laterall Radius, it were required to find to find the versed sine of 40 gr; here the sine of the complement is A 50, and therefore B 50 is the *versed sine* required. Or if I reckon from B, at the end of the Sector, toward the center, the distance from 90 to 80, is the versed sine of 10 gr; from 90 to 70, the versed sine of 20 gr; from 90 to 60, is the versed sine of 30 gr; and so in the rest.

If A D be the giuen sine of 50 gr. and it be required to find the *versed sine* of 50 gr; here because A D is vncquall to the laterall sine of 50 gr, I make it a parallell. And first I find the radius A C; then the sine of the complement A 40, which being taken out of A C, leaueth C 40 for the versed sine of 50 gr. which was required.

But if the arke, whose versed sine is required, be greater then the quadrant, his versed sine also is greater then the Radius, by the right sine of his excelsse about 90 gr.

As if A C being the Radius giuen, it were required to find the versed sine of 130 gr: here the excelsse about 90 gr. is 40 gr; and therefore the versed sine required is equall to the Radius A C and A 40, both being set together.

5 The Diameter or Radius being giuen to finde the Chords of euery arke.

The sines may be fitted many wayes to serue for chords.

1 A sine being the halfe of the chord of the double arke, if the sine be doubled, it giueth the chord of the double ark,

48 *The generall use of Sines and Tangents.*

each degree of the sines into two, that so they might shew how far the halfe degrees do reach in the sines, and yet stand for whole degrees when they are used as chords.

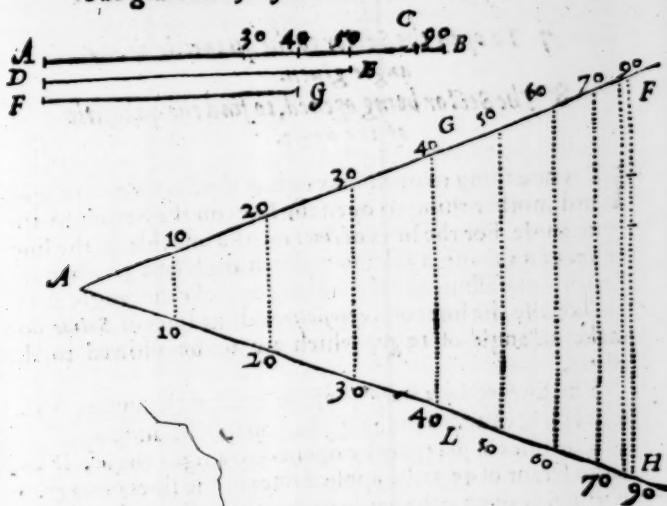
Wherefore if the Radius of the circle given be equall to the laterall semiradius (the sine of 30 Gr. and chord of 60 Gr.) there needs no farther work then to take the sine of 10 Gr. for a chord of 20 Gr. and a sine of 15 Gr. for a chord of 30 Gr. &c.

But if the Radius of the circle given be either greater or lesser then the laterall semiradius, take the diameter of it, and make it a parallell chord of 180 Gr. by applying it ouer the lines of *Sines* between 90 and 90: or take the Radius or Semidiameter which is equall to the chord of 60 Gr. and make it a parallell Radius of 60 Gr. by applying it ouer in the sines of 30 and 30, and keepe the Sector at this angle. The parallels taken from the laterall chords shall be the chords required.

As if the diameter of a circle given were the line *AB*, and it were required to find the chord of 80 gr: first I make *AB* a parallell chord of 180 Gr. or the halfe of it a parallell chord of 60 Gr; so his parallell *LG* doth giue me *FG* the chord of 80 Gr. which was required.

3 Seeing that as the sine of the complement of the halfe arke is vnto the Radius, so the sine of the same whole arke is vnto the chord of it: if we seeke but for one single chord, we may finde it without either doubling the sines, or doubling the number. For applying ouer the Radius given in the line of the complement of halfe the arke required, his parallell sine shall be the chord required.

As if the semidiameter of the circle given were *AC*, and it were required to find the chord of 40 Gr: the halfe of 40 Gr. is 20 Gr. the complement of 20 Gr. is 70 Gr. Wherefore I make *AC* a parallell sine of 70 Gr. and his parallell sine *GL* doth giue me *FG* the chord of 40 Gr. agreeable to the semidiameter *AC*.



6 The chord of any arke being given to find the diameter and Radius.

TUrne the chord given vnto a parallell chord, and his parallell semiradius shall be the semidiameter, and the parallell radius shall be the diameter.

As if FG be the chord of 80 gr . I put this ouer in G and L , the sine of 40 , and chord of 80 gr . and the parallell chord of 180 gr . giueth me AB the diameter required.

Or if I turne the chord given into a parallell line of the same quantitie, his parallell line of the complement of halfe the arke, doth giue me the semidiameter.

As if FG be the given chord of 40 gr . I put it ouer in G and L , the lines of 40 gr ; then because the halfe of 40 gr . is 20 gr . and the complement of 20 gr . is 70 gr . I take out the parallell sine of 70 gr . and it giueth me AC for the semidiameter, agreeable to that chord of 40 gr .

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7 The

The generall use of Sines and Tangents

- 7 To open the Sector to the quantitie of any
angle given.
8 The Sector being opened, to find the quantitie
of the angle.

IT is one thing to open the edges of the Sector to an angle, and another thing to open the lines on the Sector to the same angle. For the lines of *lines* on the one side, & the lines of *sines* on the other side, do make an angle of 2 gr. when the Sector is close shut, and the edges doe make no angle at all. So likewise the lines of *Superficies* and the lines of *Solids* doe make an angle of 10 gr, which are to be allowed to the edges.

The lines of *lines* may be opened to a right angle, if the whole line of 100 parts be applied ouer in 80 and 60.

The lines of *sines* may be opened to a right angle, if the large secant of 45 gr. be applied ouer in the sines of 90 gr. or if the sine of 90 gr. be applied ouer in the sines of 45 gr. or if the sine of 45 gr. be applied ouer in the sines of 30 gr.

If it be required to open those lines to any other angle, take out the chord thereof, and apply it ouer in the *semiradius*, and those lines shall be opened to that angle.

As if it were required to open the Sector in the lines of *sines* to an angle of 40 gr, take out the chord of 40 gr, and so it open the Sector in the chord of 60 gr; so shall the lines of *sines* be opened to the angle required. Or if the same chord of 40 Gr. be applied ouer betweene 50 and 50, in the lines of *lines*, they shall also be opened to the same angle. If it be applied ouer in 25 of the lines of *Superficies*, or 125 in the lines of *Solids*, they also shall be opened to the same angle: because the chord of 60 Gr. or sine of 30 Gr. and 50 in the lines of *lines*, and 25 in the lines of *Superficies*, and 125 in the *Solids*, are all of the same length with the semiradius.

Or if the *Semiradius* be applied ouer betweene the sine of 30 Gr. and the sine of the complement of the angle required, it will open the lines of *Sines* to that angle.

As

The generall use of Sines and Tangents. 37

As if the semiradius be applied over in the sines of 30 Gr. and the sine of 50 Gr. it shall open the lines of *Sines* to an angle of 40 Gr.

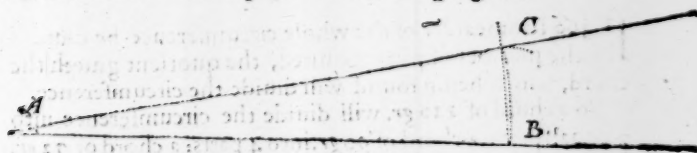
On the contrary, if the *Sector* be opened to an angle, and it be required to know the quantitie thereof, open the compasses to the semiradius, and setting one foote in the sine of 30 Gr. turne the other toward the other line of *sines*, and it shall fall there in the complement of the angle; if it fall on 50 Gr. the angle is 40 Gr; if on 60 Gr. the angle is 30 Gr. &c.

Or take over the parallell chord of 60 Gr. and measure it in the laterall chord; and it shall there shew the quantitie of the angle. As if the *Sector* being opened to an angle, I should take over the parallell of 30 Gr. of the sines, and 60 Gr. of the chords, and measure it in the laterall chords, find it to be 40 Gr; the angle comprehended between the lines of *Sines* is 40 Gr. but the angle between the edges of the *Sector* is 2 Gr. lesse, and therefore but 38 Gr.

9 To finde the quantitie of any angle given.

IF out of the angular point, to the quantitie of the *Semiradius*, be described an occult arke that may cut both sides of the angle, the chord of this arke measured in the laterall chord, shall give the quantitie of the angle.

Let the angle given be *BAC*: first I take the *Semiradius* with the compasses, and setting one foote in *A*, I cut the sides of the angle in *B* and *C*; then I take the chord *BC*, and measure it in the laterall chord, and I find it to be 11 Gr. and 15 M. and such is the quantitie of the angle given.



Or if the arke be described out of the angular point at any other distance, let the semidiameter be turned into a parallel

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parallel

parallel chord of 60 Gr. then take the chord of this arke, and carrie it parallel till it crosse in like chords: so the place where it stayeth shall giue the quantitie of the angle.

As in the former example, if I make the semidiameter AB a parallel chord of 60 Gr. and then keeping the Sector at that angle, carrie the chord BC parallel, till it stay in like chords; I shall finde it to stay in no other but 11 Gr. 15 M. and such is the angle BAC .

10 *Vpon a right line and a point giuen in it, to make an angle equall to any angle giuen.*

First out of the point giuen describe an arke, cutting the same line: then by the 5. Prop. afore, find the chord of the angle giuen agreeable to the semidiameter, and inscribe it into this arke: so a right line drawne through the point giuen, and the end of this chord, shall be the side that makes vp the angle.

Let the right line giuen be AB , and the point giuen in it be A , and let the angle giuen be 11 gr. 15 m. Here I open the compasses to any semidiameter AB , (but as oft as I may conveniently to the laterall semiradius) and setting one foot in A , I describe an occult arke BC ; then I seeke out the chord of 11 gr. 15 m. and taking it with the compasses, I set one foote in B , the other crosseth the arke in C , by which I draw the line AC , and it makes vp the angle required.

11 *To diuide the circumference of a circle into any parts required.*

If 360 the measure of the whole circumference be diuided by the number of parts required, the quotient giueth the chord, which being found will diuide the circumference.

So a chord of 120 gr. will diuide the circumference into 3 equall parts; a chord of 90 gr. into 4 parts; a chord of 72 gr. into 5 parts; a chord of 60 gr. into 6 parts; a chord of 51 gr. 26. into 7 parts; a chord of 45 gr. into 8 parts; a chord of 40 gr. into

into 9 parts; a chord of 36 gr. into 10 parts; a chord of 32 gr. 44 m. into 11 parts; a chord of 30 gr. into 12 parts.

In like manner if it be required to diuide the circumference of the circle, whose semidiameter is AB , into 32: first I take the semidiameter AB , and make it a parallell chord of 60 gr.; then because 360 gr. being diuided by 32, the quotient will be 11 gr. 15 m. I find the parallell chord of 11 gr. 15 m. and this will diuide the circumference into 32.

But here the parts being many, it were better to diuide it first into fewer, and after to come ouer it againe. As first to diuide the circumference into 4, and then each 4 parts into 8, or otherwise, as the parts may be diuided.

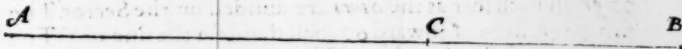
12 To diuide a right line by extreme and meane proportion.

THE line to be diuided by extreme and meane proportion, hath the same proportion to his greater segment, as in figures inscribed in the same circle, the side of an hexagon a figure of six angles, hath to a side of a decagon a figure of ten angles: but the side of a hexagon is a chord of 60 gr. and the side of a decagon is a chord of 36 gr.

Let AB be the line to be diuided: if I make AB a parallell chord of 60 gr. and to this semidiameter find AC a chord of 36 gr. this AC shall be the greater segment, diuiding the whole line in C , by extreme and meane proportion. So that,

As AB the whole line is vnto AC the greater segment: so AC the greater segment vnto CB the lesser segment.

Or let AC be the greater segment giuen: if I make this a parallell chord of 36 gr. the correspondent semidiameter shall be the whole line AC , and the difference CB the lesser segment.



Or let CB be the lesser segment giuen: if I make this a parallell chord of 36 gr. the correspondent semidiameter

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shall

✓

shall be greater segment AC , which added to CB , giueth the whole line AB .

To auoid doubling of lines or numbers, you may put ouer the whole line in the *Sines* of 72 gr. and the parallell line of 36 gr. shall be the greater segment.

Or if you put ouer the whole line in the *sines* of 54 gr. the parallell line of 30 gr. shall be the greater segment, and the parallell line of 18 gr. shall be the lesser segment.

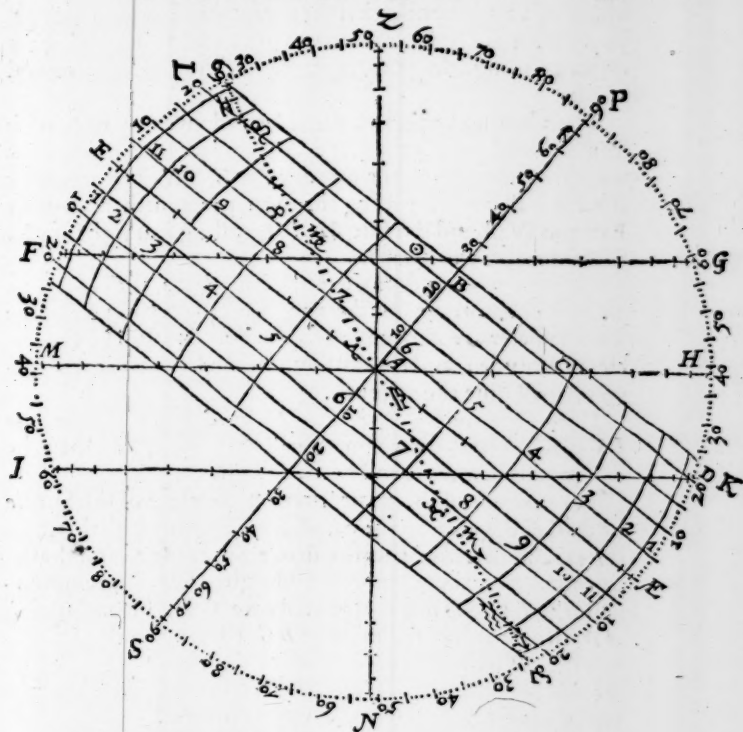
CHAP. III.

Of the proiection of the Sphere in Plano.

THe Sphere may be proiected in *Plano* in streight lines, as in the *Analemma*, if the semidiameter of the circle giuen be diuided in such sort as the line of *Sines* on the Sector.

As if the Radius of the circle giuen were AE , the circle thereon described may represent the plane of the generall meridian, which diuided into foure equal parts in E, P, \mathcal{A}, S , and crossed at right angles with $E\mathcal{A}$ and PS , the diameter $E\mathcal{A}$ shall represent the equator, and PS the circle of the houre of 6. And it is also the axis of the world, wherein P stands for the North pole, and S for the South pole. Then may each quarter of the meridian be diuided into 90 gr. from the equator towards the poles. In which if we number 23 gr. the greatest declination of the Sun from E to 69 Northwards, from \mathcal{A} to \mathcal{W} Southward, the line drawne from 69 to \mathcal{W} shall be the ecliptique, and the lines drawne parallell to the equator through \mathcal{S} and \mathcal{W} shall be the tropiques.

Hauiug these common sections with the plane of the meridian, if we shall diuide each diameter of the Ecliptique into 90 gr. in such sort as the *Sines* are diuided on the Sector. The first 30 gr. from A toward 69 , shall stand for the sine of V . The 30 gr. next following for \mathcal{V} . The rest for $II, \mathcal{S}, \mathcal{N}, \&c.$ in their order. So that by these meanes we haue the place of the Sun for all times of the yeare.



If againe we diuide AP , AS , in the like sort, and set to the numbers 10. 20. 30. &c. vnto 90 gr. the lines drawne through each of these degrees parallell to the equator, shall shew the declination of the Sunne, and represent the parallels of latitude.

If farther we diuide AE , AS , and his parallels in the like sort, and then carefully draw a line through each 15 gr. so as it makes no angles; the lines so drawne shall be *ellipticall* and represent the houre-circles. The meridian PE the houre

houre of 12 at noone; that next vnto it drawne through 75 gr. from the center the houres of 11 and 1, that which is drawne through 60 gr. from the center the houres of 10 and 2. &c.

Then hauing respect vnto the latitude, we may number it from *E* Northward vnto *Z*, and there place the zenith: by which and the center the line drawne *ZAN* shall represent the verticall circle, passing through the zenith and nadir East and West, and the line *MAH* crossing it at right angles shall represent the horizon.

These two being diuided in the same sort as the ecliptique and the equator, the line drawne through each degree of the semidiameter *AZ*, parallell to the horizon, shall be the circles of altitude, and the diuisions in the horizon and his parallels shall giue the azimuth.

Lastly, if through 18 gr. in *AN*, be drawne a right line *IK* parallell to the horizon, it shall shew the time when the day breaketh, and the end of the twilight.

For example of this proiection, let the place of the Sunne be the last degree of ϵ , the parallell passing through this place is *LD*, and therefore the meridian altitude *ML*, and the depression below the horizon at midnight *HD*: the semidiurnall arke *LC*, the seminocturnall arke *CD*, the declination *AB*, the ascensionall difference *BC*, the amplitude of ascension *AC*. The difference betweene the end of twilight and the day breake is very small; for it seemes the parallell of the Sunne doth hardly crosse the line of twilight.

If the altitude of the Sunne be giuen, let a line be drawne for it parallell to the horizon; so it shall crosse the parallell of the Sunne, and there shew both the azimuth and the houre of the day. As if the place of the Sunne being giuen as before, the altitude in the morning were found to be 20 gr. the line *FG* drawne parallell to the horizon through 20 gr. in *AZ*, would crosse the parallell of the Sunne in \odot . Wherefore *F \odot* sheweth the azimuth, & *L \odot* the quantitie of houres from the meridian. It seemes to be about halfe an houre past 6 in the morning, and yet more then halfe a point short

short of the East.

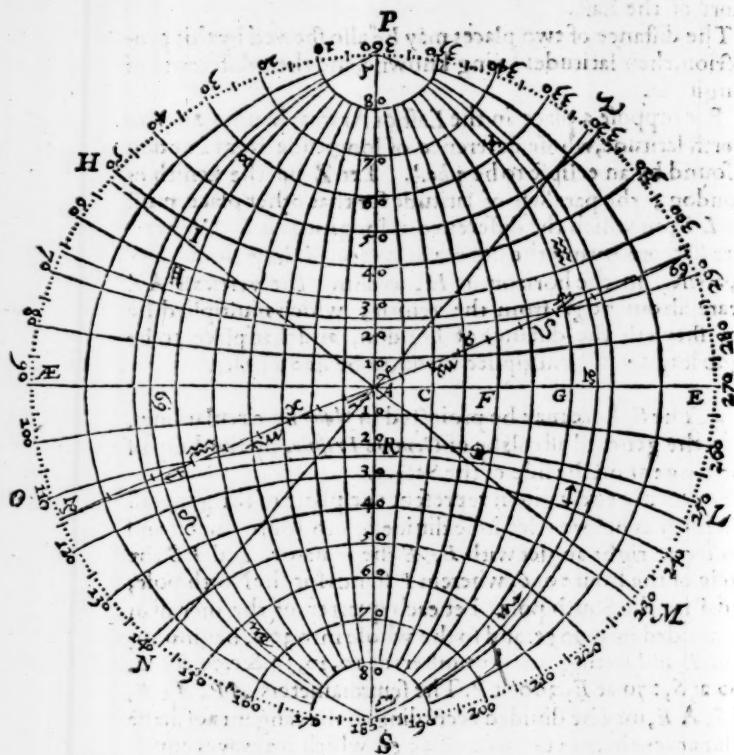
The distance of two places may be also shewed by this projection, their latitudes being knowne, and their difference of longitude.

For suppose a place in the East of Arabia, having 20 gr. of North latitude, whose difference of longitude from London is found by an eclipse to be 560. $\frac{1}{2}$. Let Z be the zenith of London, the parallell of latitude for that other place must be LD , in which the difference of longitude is $L\odot$. Wherefore \odot representing the site of that place, I draw through \odot a parallell to the horizon MH , crossing the verticall AZ neare about 70 gr. from the zenith, which multiplied by 20, sheweth the distance of London, and that place to be 1400 leagues. Or multiplied by 60, to be 4200 miles.

2 The Sphere may be projected in *plano* by circular lines, as in the generall astrolabe of *Gemma Frisius*, by the help of the tangent on the side of the Sector.

For let the circle giuen represent the plane of the generall meridian as before; let it be diuided into foure parts, and crossed at right angles with $E\odot E$ the equator, and PS the circle of the houre of 6, wherein P stands for the North pole, and S for the South pole. Let each quarter of the meridian be diuided into 90 gr. and so the whole into 360, beginning from P , and setting to the numbers of 10, 20, 30. &c. 90 at E , 180 at S , 270 at E , 360 at P . The semidiameters AP , $A\odot E$, AS , AE , may be diuided according to the tangents of halfe their arkes, that is a tangent of 45 gr. which is alwayes equall to the Radius, shall giue the semidiameter of 90 gr; a tangent of 40 gr. shall giue 80 gr. in the semidiameter: a tangent of 35 gr. shall giue 70. &c. So that the semidiameters may be diuided in such sort as the tangent on the side of the Sector, the difference being onely in their numbers.

Hauiug diuided the circumference and the semidiameters, we may easily draw the meridians and the parallels by the helpe of the Sector.



The meridians are to be drawne through both the poles *P* and *S*, and the degrees before, graduated in the equator. The distance of the center of each meridian from *A* the center of the plane, is equall to the tangent of the same meridian, reckoned from the generall meridian *PAE*, and the semidiameter equall to the secant of the same degree.

As for example, if I should draw the meridian *PBS*, which is the tenth from *PAE*, the tangent of 10 gr. giueth me *AC*, and the secant of 10 gr. giueth me *SC*, whereof *C* is the center

center of the meridian PBS , and CS his semidiameter: so AF a tangent of 20 gr. sheweth F to be the center of PDS , the twentieth meridian from $PAES$, and AG a tangent of 23 gr. 30 M. sheweth G to be the center of $P69S$, &c.

The parallels are to be drawne through the degrees, in AP , AS , and their correspondent degrees in the generall meridian. The distance of the center of each parallell from A the center of the plane, is equall to the secant of the same parallell from the pole, and the semidiameter equall to the tangent of the same degree. As if I should draw the parallell of 80 gr. which is the tenth from the pole S , first I open the compasses vnto AC the tangent of 10 gr. and this giueth me the semidiameter of this parallel, whose center is a little from S , in such distance as the secant SC is longer then the radius SA .

The meridians and parallels being drawne, if we number 23 gr. 30 m. from E to \odot Northward, from \odot to Ψ Southward, the line drawne from \odot to Ψ shall be the ecliptique: which being diuided in such sort as the semidiameter AP , the first 30 gr. from A to \odot shall stand for the sine of γ ; the 30 gr. next following for δ ; the rest for π . \odot . Ω . &c. in their order.

If farther we haue respect vnto the latitude, we may number it from E Northward vnto Z , and there place the zenith, by which and the center, the line drawne ZAN shall represent the verticall circle, and the line MAH crossing it at right angles, shall represent the horizon; and these diuided in the same sort as AP , the circles drawne through each degree of the semidiameter AZ , parallell to the horizon, shall be the circles of altitude: and the circles drawne through the horizon and his poles, shall giue the azimuths.

For example of this projection, let the place of the Sunne be in the beginning of π , the parallell passing through this place is $\pi \odot L$; and therefore the meridian altitude ML , and the depression below the horizon at midnight $H\pi$, the semi-diurnall arke $L \odot$, the seminocturnall arke $O \odot$, the declination AR , the ascensionall difference $R \odot$, the ampli-

tude of ascension $A\odot$.

Or if A be put to represent the pole of the world, then shall $PAESE$ stand for the equator, and $PSSW$ for the ecliptique, and the rest which before stood for meridians, may now serue for particular horizons, according to their severall eleuations. Then suppose the place of the Sunne giuen to be 24 gr. of φ , his longitude shall be PI , his right ascension PH , his declination HI . And if the place giuen be 19 gr. of Ω , his longitude shal be PK , his right ascension PN , his declination NK . Againe, the declination brought to the horizon of the place, shall there shew the ascensionall difference, amplitude of ascension, and the like conclusions of the globe. But I intend not here to shew the vse of the Astrolabe, but the vse of the Sector in projection.

And after this maner may a nocturnall be projected to shew the houre of the night, whereof I will set downe a type for the vse of Sea-men.



It consists as you see of two parts, the one is a plane, diuided equally according to the 24 houres of the day, and each houre into quarters or minutes, as the plane will beare: the line from the center to XII ; stands for the meridian, and XII stands for the houre of 12 at midnight. The other part is a rundle for such starres as are neare the North pole, together with the twelue moneths, and the dayes of each moneth fitted to the right ascension of the starres. Those that haue occasion to see the South pole, may do the like for the Southern constellations, and put them in a rundle on the back of this plane, and so it may serue for all the world.

The vse of this nocturnall is easie and ready. For looke vp to the pole, and see what starres are neare the meridian, then place the rundle to the like situation, so the day of the moneth will shew the houre of the night.

3 The Sphere may be projected in *plane* by circular lines, as in the particular Astrolabe of *Ioh. Stöbberlin*, by help of the tangent, as before.

For let the circle giuen represent the tropique of φ , let it be diuided into foure parts, and crossed at right angles with AC the equinoctiall colure, and MB the solstitiall colure, and generall meridian, the center P representing the pole of the world. Let each quarter be diuided into 90 gr. and so the whole into 360, beginning from A towards B . The meridian PM , or PB , may be diuided according to the tangent of halfe his arke. So as the arke from the North pole to the tropique of φ , being 90 gr. and 23 gr. 30 m. that is 113 gr. 30 m. and the halfe arke 56 gr. 45 m. the meridian shall be diuided into 90 gr. and 23 gr. 30 m. in such sort as the tangent of 56 gr. 45 m. on the side of the Sector is diuided into degrees and halfe degrees; of which PAE the arke of the equator 90 gr. from the pole, shall be giue by the tangent of 45 gr. And $P69$ the arke of the Summer tropique 66 gr. 30 m. from the pole, shall be giuen by the tangent of 33 gr. 15 m. And the circles drawne vpon the center P through E and 69 , shall be the equator, and the Summer tropique.

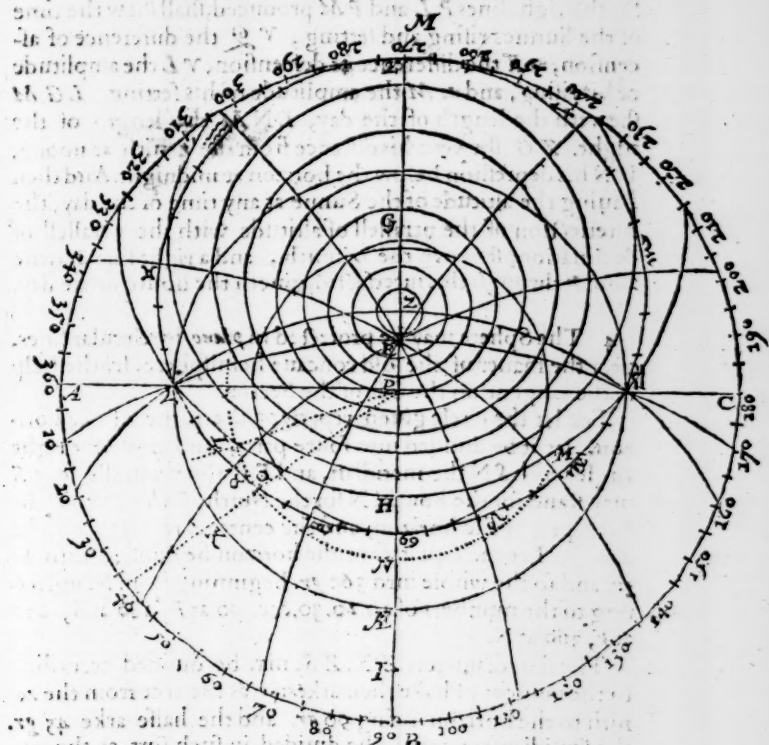
Having the equator and both the tropiques, the eclip-

tique $\gamma \text{ } \approx \text{ } \psi$ shall be drawne from the one tropique to the other, through the interfection of the equator and the equinoctiall colure. And it may be diuided first into the twelue Signes after this maner: PE the arke of the pole of the ecliptique $23 \text{ gr. } 30 \text{ m.}$ from the pole of the world, shall be giuen by the tangent of $11 \text{ gr. } 45 \text{ m.}$ The center of the circle of longitude passing through this pole $E \gamma$ and \approx , shall be found at D (somewhat below B) by the tangent of $66 \text{ gr. } 30 \text{ m.}$ Then through D draw an occult line parallel to AC , and diuide it on each side from D , in such sort as the tangent is diuided on the side of the *Sector*, allowing 45 gr. to be equall to DE . So the thirtieth degree from D toward the right hand, shall be the center of the circle of longitude passing through $E \delta$ and m . The sixtith degree, the center of II $E \zeta$. The thirtieth degree from D toward the left hand, the center of \times $E \text{m}$. The sixtith, the center of \approx $E \eta$. And the other intermediate degrees shall be the centers to diuide each Signe into 30 gr.

If farther we haue respect vnto the latitude, we may (the meridian being before diuided) number it from P Northward vnto H , and there place the North interfection of the meridian and horizon: then the complement of the latitude being numbred from P Southward vnto Z , shall there giue the zenith; and 90 gr. from Z Southward vnto F , shall there giue the South interfection of the meridian and horizon. The middle betweene F and H shall be G the center of the horizon $\gamma H \approx F$, passing through the beginning of γ and \approx , vnlesse there be some former error.

All parallels to the horizon may be found in like sort by their interfections with the meridian, and the middle betweene those interfections is alwayes the center.

The azimuths may be drawne as the circles of longitude were before. For the center of the first verticall $\gamma Z \approx$ will be found at I (somewhat neare vnto B) by the tangent of the latitude. And if through I we draw an occult line parallel to AC , and diuide it on each side from I , in such sort as the tangent is diuided on the side of the *Sector*, allowing 45 gr. to be equall



equal to $\frac{1}{2}Z$; these divisions shall be the centers, and the distance from these divisions unto Z , shall be the semidiameters whereon to describe the rest of the azimuths.

For example of this projection, let \odot the place of the Sun given be to 27° of \odot : a right line drawne from P through this place unto the equator, shall shew his right ascension γK , and his declination $K \odot$. Then may we on the center P and semidiameter $\odot P$, draw an occult parallell of declination, crossing the horizon in L and M , the meridian in G and N .

So

So the right lines $P L$ and $P M$ produced, shall shew the time of the Sunnes rising and setting, $\vee Q$ the difference of ascension, $\approx R$ the difference of descension, $\vee L$ the amplitude of his rising, and $\approx M$ the amplitude of his setting. $L G M$ sheweth the length of the day, $L N M$ the length of the night. $Z G$ sheweth his distance from the zenith at noone, $H N$ his depression below the horizon at midnight. And then hauing the altitude of the Sunne at any time of the day, the interfection of the parallell of altitude with the parallell of declination, sheweth the azimuth, and a right line drawne from P through this interfection, giueth the houre of the day.

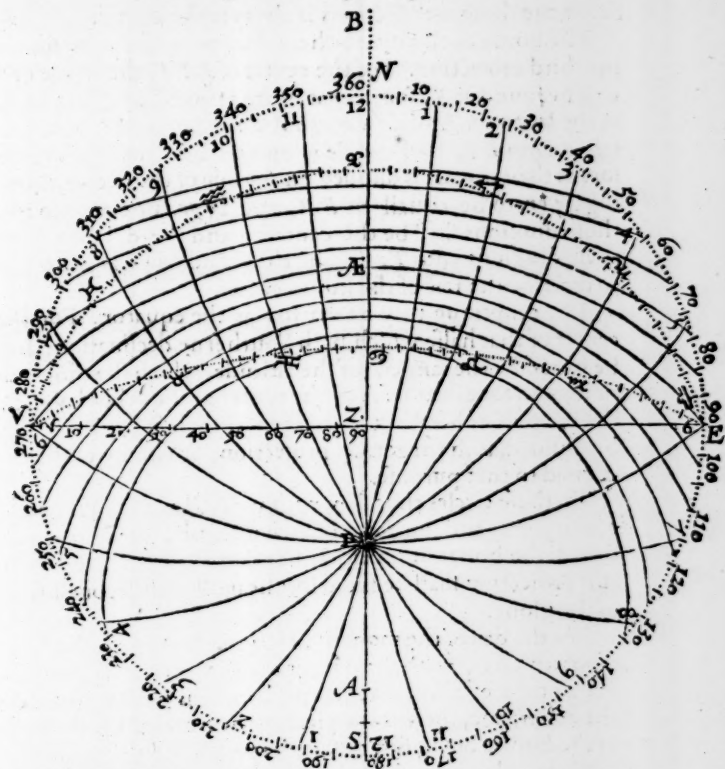
4 The Sphere may be projected in *plano* by circular lines, after the maner of the old concaue hemisphere, by the help of the tangent on the side of the Sector.

For let the circle giuen represent the plane of the horizon, let it be diuided into foure parts, and crossed at right angles with $S N$ the meridian, and $E V$ the verticall; so as S may stand for the South, N for the North, E the East, V the West part of the horizon, and the center Z representeth the zenith. Let each quarter of the horizon be diuided into 90 *gr.* and so the whole into 360 *gr.* beginning from N , and setting to the numbers of 10. 20. 30. &c. 90 at E , 180 at S , 270 at V , 360 at N .

The semidiameters $Z N$, $Z S$, may be diuided according to the tangent of halfe their arkes: So as the arke from the zenith to the horizon being 90 *gr.* and the halfe arke 45 *gr.* the semidiameters are to be diuided in such sort as the tangent of 45 *gr.* as was shewed before in the second projection. And if from Z we draw circles through each of these diuisions, they shall be parallels of altitude.

Then hauing respect vnto the latitude, we may (the meridian being before diuided) number it from Z to E , and there place the interfection of the meridian and equator. The complement of the latitude from Z vnto P , shall there giue the pole of the world, and 90 further from P shall there giue the other interfection of the meridian and equator.

The



The middle between these interfections shall be *A*, the center of the equator, passing through *E* and *V*, vnlesse there be some former error. The interfections of the tropiques depend on the equator. From *E* 23 gr. 30 m. farther shall be *W*, the interfection of the meridian & the Southerne tropique. From *E* 23 gr. 30 m. nearer shall be *U*, the interfection of the meridian and the Northerne tropique. The interfections of the other intermediat parallels, shall be giuen in like sort, by their degrees of distance from the equator, and the middle

K

bc.

betweene those intersections is alwayes the center.

The houre circles may be here drawne as the azimuths in the third projection. For the center of EPV , the houre of 6 will be found at B (somewhat neare vnto N) by the tangent of the latitude. And if through B we draw an occult line parallel vnto EV , and diuide it on each side from B , in such sort as the tangent is diuided on the side of the Sector, allowing 45 gr. to be equall to BP , and 15 gr. for euery houre: those diuisions shall be the centers, and the distance from these diuisions vnto P , shall be the semidiameters, whereon to describe the rest of the houre circles.

The ecliptique may be drawne as the equator. For the center of that halfe which hath Southerne declination, shall be giuen by the tangent of the altitude, which the Sun hath in his entrance into Ψ . And the center of the other halfe, by the tangent of his altitude, at his entrance into Ξ . And it may be diuided, as in the former projection, or else by tables calculated to that purpose.

To these circles thus drawne, if we shall adde the moneths of the yeare, and the dayes of each moneth, as we may well doe, at the horizon, on either side betweene the tropiques; this projection shall be fitted for the most vsfull conclusions of the globe.

For the day of the moneth being giuen, the parallell that shooteth on it, doth shew what declination the Sunne hath at that time of the yeare. And where this parallell crosseth the ecliptique, there is the place of the Sunne. Or the place of the Sunne being first giuen, the parallell which crosseth it shall at the horizon shew the day of the moneth. Either of these then being giuen, or onely the parallell of declination, we may follow it first vnto the horizon, there the distance of the end of the parallell from E or V , sheweth the amplitude; the same among the houre circles sheweth the time when the Sunne riseth or setteth. Then hauing the altitude of the Sunne at any time of the day, the intersection of the parallell of declination with the parallell of altitude, sheweth the houre of the day; and a right line drawne from Z through

through this intersection to the horizon, giue the azimuth.

Thus in either of these projections, that which is otherwise most troublesome, is easily done by the help of the *tangent* line: and what I haue said of this line, the same may be wrought by scale and numbers out of the table of Tangents.

CHAP. IV.

Of the resolution of right-line Triangles.

IN all Triangles there being six parts, viz. three angles and three sides, any three of them being giuen, the rest may be found by the Sector.

As in a Rectangle triangle,

I *To finde the base, both sides being giuen.*

Let the Sector be opened in the lines of *Lines* to a right angle, (as before was shewed *Cap. 2. Prop. 7.*) then take out the sides of the triangle, and lay them, one on one line, the other on the other line, so as they meete in the center, and marke how farre they extend. For the line taken from the termes of their extension, shall be the base required, viz. the side opposite to the right angle.

Or adde the squares of the two sides (as in *Prop. 4. Superf.*) and the side of the compound square shall be the base.

2 *To find the base by hauing the angles,
and one of the sides giuen.*

Take the side giuen, and turne it into the parallell line of his opposite angle; so the parallell Radius shall be the base.

3 *To find a side by hauing the base,
and the other side giuen.*

Let the Sector be opened in the lines of *lines* to a right angle,

angle, and the side giuen laid on one of those lines from the center; then take the base with a paire of compasses, and setting one foote in the terme of the giuen side, turne the other to the other line of the Sector, and it shall there shew the side required.

Or take the square of the side out of the square of the base (as in *Prop. 4. Superf.*) and the side of the remaining square shall be the side required.

*4 To find a side hauing the base
and the angles giuen.*

Take the base giuen, and make it a parallell Radius, so the parallell *sines* of the angles, shall be the opposite sides required.

*5 To find a side by hauing the other side
and the angles giuen.*

Take the side giuen, and turne it into his parallell *sine* of his opposite angle; so the parallell *sine* of the complement shall be the side required.

*6 To find the angles by hauing the base
and one of the sides giuen.*

First take out the base giuen, and laying it on both sides of the Sector, so as they may meete in the center, and marke how farre it extendeth. Then take out the laterall Radius, and to it open the Sector in the termes of the base. This done, take out the side giuen, and place it also on the same lines of the Sector from the center. For the parallell taken in the termes of this side, shall be the *sine* of his opposite angle.

Or take the base giuen, and make it a parallell Radius; then take the side giuen, and carrie it parallell to the base, till it stay in like *sines*: so they shall giue the quantitie of the

the opposite angle.

7 To finde the angles by having both the sides giuen.

Take out the greater side, and lay it on both sides of the Sector, so as they meete in the center, and marke how farre it extendeth. Then take the other side, and to it open the Sector in the termes of the greater side; so the parallell Radius shall be the tangent of the lesser angle. The third angle is alwayes knowne by the complement.

8 The Radius being giuen, to find the tangent, and secant of any arke.

9 The tangent of any arke being giuen, to find the tangent thereof, and the Radius.

10 The secant of any arke being giuen, to find the tangent thereof, and the Radius.

The tangent, and the secant, together with the Radius of euery arke, do make a right angle triangle; whose sides are the Radius and tangent, and the base alwayes the secant; and the angles alwayes knowne by reason of the giuen arkes. Wherefore the solution is the same with those before.

In any right-lined triangle whist soeuer,

11 To find a side by knowing the other two sides, and the angle contained by them.

Let the Sector be opened in the lines of lines to the angle giuen, then take out the sides of the triangle, & laying them the one on the one line, the other on the other, so as they meete in the center, marke how far they extend. For the line taken between the termes of their extension, shall be the third side required.

- 12 To find a side by having the other two sides,
and one of the adjacent angles, so it be
knowne which of the other angles
is acute or oblique.

Let the *Sector* be opened in the lines of *lines* to the angle giuen, and the adjacent side layd on one of those lines from the center; then take the other side with a paire of compasses, and setting one foote in the terme of the former giuen side, turne the other to the other line of the *Sector* which here representeth the side required, and it shall crosse it in two places; but with which of them is the terme of the side required, must be iudged by the angle.

As if in the triangle following, the side AC being giuen, and the side CD and the angle CAD $18\text{ gr. }40\text{ m.}$ it were required to find the side AD .

First I open the *Sector* in the lines of *lines* to an angle of $18\text{ gr. }40\text{ m.}$ and laying the adjacent side from the center A , it extendeth to 800 in C . Then I take the other side CD with the compasses, and setting one foote in C , and turning the other to the other line of the *Sector*, I find that it doth crosse it both in B and D ; so that it is vncertaine whither the side required be AB or AD , onely it may be iudged by the angle. For if the inward angle where they crosse be obtuse, the side required is the lesser; if it be acute, it is the greater.

- 13 To find a side by having the angles
and one of the other sides giuen.

Take the side giuen, and turne it into the parallell line of his opposite angle; so the parallell lines of the other angle shall be the opposite sides required.

14 *To find the proportion of the sides
by having the three angles.*

Take the laterall lines of the angles, and measure them in the line of *lines*. For the numbers belonging to those lines do giue the proportion of the sides.

15 *To finde an angle by knowing the
three sides.*

Let the two containing sides be layd on the lines of the *Sector* from the center, one on one line, and the other on the other; and let the third side, which is opposite to the angle required, be fitted ouer in their termes: so shall the *Sector* be opened in those lines to the quantitie of the angle required.

The quantitie of this angle is found as in *Cap. 2. Prop. 8.*

16 *To finde an angle by having two sides
and one adjacent angle.*

First take out the side opposite to the angle giuen, and laying it on both sides of the *Sector*, so as they meete in the center, marke how farre it extenderh; then take out the laterall line of the angle, and to it open the *Sector* in the termes of the first side: this done, take out the other side giuen, and place it also on the same lines of the *Sector* from the center, for the parallels taken in the termes of this side, shall be the line of the angle opposite to the second side.

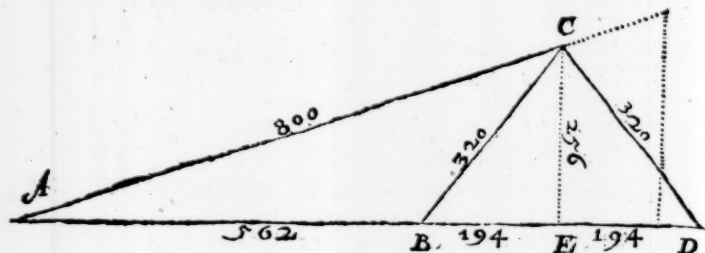
Or take out the side opposite to the angle giuen, and make it a parallell line of that angle; then take the other side giuen and carrie it parallell to the former, till it stay in like lines: so they shall giue the quantitie of the angle opposite to the second side.

17 To finde an angle by having two sides,
and the angle contained by them.

First find the third side by the 11. Prop. and then the angles may be found by the 15. or 16. Prop.

For practise in each of these cases, we may use the examples following, wherein CEA , CEB , CED are rectangle in E ; the rest consist of oblique angles.

CAB	18	gr.	40	m.
ABC	126		52	
ACB	34		28	
ACD	108		12	
ADC	53		8	
BCD	73		44	



For observation of angles, the Sector may have sights set on the moueable foote; so that by looking through them, the edges of the Sector may be applied to the sides of the angle.

For.

For measuring of the sides of lesser triangles, any scale may suffice, either of feet, or inches, or lesser parts. But for greater triangles, especially for plotting of grounds, I hold it fit to use a chaine of foure perches in length, diuided into an hundred links. For so the length being multiplied into the breadth, the five last figures giue the content in roods and perches by this Table; the other figures toward the left hand, doe shew the number of acres directly.

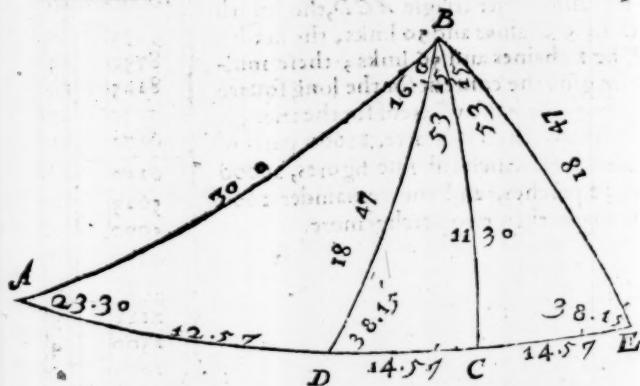
As if in the former triangle *ACD*, the length *AD* be 9 chaines and 50 links, the breadth *CE* be 2 chaines and 56 links; these multiplied giue the content for the long square 2. 43200, the halfe whereof for the triangle is 1. 21600, that is 1 acre, 21600 parts of 100000, of which last five figures, 20000 giue 32 perches, and the remainder 1600 giue better then two perches more.

Links	R	P
100000	4	0
90000	3	24
80000	3	8
70000	2	32
60000	2	16
50000	2	0
40000	1	24
30000	1	8
20000		32
10000		16
9375		15
8750		14
8125		13
7500		12
6875		11
6250		10
5625		9
5000		8
4375		7
3750		6
3125		5
2500		4
1875		3
1250		2
625		1

CHAP. V.

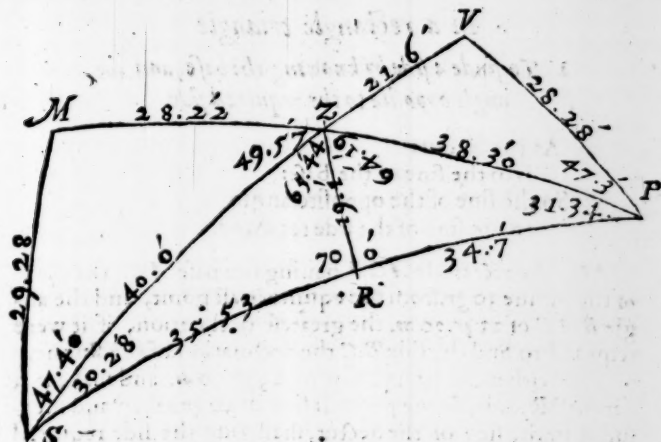
Of the resolution of Spherical triangles.

For our practise in spherical triangles, let A be the equinoctial point, AB an arke of the ecliptique representing the longitude of the Sunne in the beginning of γ , BC an arke of the declination from the Sunne to the equator, and AC an arke of the equator representing the right ascension.



Let BD be an arke of the horizon representing the amplitude of the Sunnes rising from the East, and BE an arke of the horizon for his setting from the West: so DC shall be the difference of ascension, and CE the difference of descension, AD the oblique ascension, and AE the oblique descension of the same place of the Sunne in our latitude at Oxford of $51^{\circ} 45'$ whose complement $38^{\circ} 15'$ is the angle at E and D . The triangles ACB , DCB , ECB , are rectangle in C : the other ADB , AEB , consist every way of oblique angles.

Or



Or to fit an example nearer to the latitude of *London*. Let ZPS represent the zenith pole and Sun, ZP being 38° gr. $30'$ m. the complement of the latitude, PS 70° gr. the complement of the declination, and ZS 40° gr. the complement of the Sun's altitude. The angle at Z shall shew the azimuth, and the angle at P , the hour of the day from the meridian. Then if from Z to PS we let downe a perpendicular ZR , we shall reduce the oblique triangle into two rectangle triangles ZRP , ZRS . Or if from S to ZP we let downe a perpendicular SM , we shall reduce the same ZPS into two other triangles, SMZ , SMP , rectangle at M : whatsoever is said of any of these triangles, the same holdeth for all other triangles in the like cases.

For the resolution of each of these, there be severall wayes. I onely chuse those which are fittest for the *Sector*, wherein if that be remembred which before is shewed in the generall use of the *Sector* concerning laterall and parallell entrance, it may suffice onely to set downe the proportion of the three parts given to the fourth required, and so I shew first by the *sines* alone.

In a rectangle triangle

- 1 To finde a side by knowing the base, and the angle opposite to the required side.

As the Radius

is to the sine of the base:

So the sine of the opposite angle
to the sine of the side required.

As in the rectangle ACB , having the base AB , the place of the Sunne 30 gr. from the equinoctiall point, and the angle BAC of 23 gr. 30 m. the greatest declination, if it were required to find the side BC the declination of the Sunne.

Take either the laterall sine of 23 gr. 30 m. and make it a parallell Radius; so the parallell sine of 30 gr. taken and measured in the side of the Sector, shall give the side required 11 gr. 30 m. Or take the sine of 30 gr. and make it a parallell Radius; so the parallell sine of 23 gr. 30 m. taken and measured in the laterall lines, shall be 11 gr. 30 m. as before.

So in the triangle ZPS having ZP 38 gr. 30 m. and the angle P 31 gr. 34 m. given, we shall find the perpendicular ZR to be 19 gr. 1 m; or having PS 70 gr. and the said angle P 31 gr. 34 m. given, we may finde the perpendicular SA to be 29 gr. 28 m.

- 2 To finde a side by knowing the base and the other side.

As the sine of the complement of the side given
is to the Radius:

So the sine of the complement of the base
to the sine of the complement of the side required.

So in the rectangle ACB , having AB 30 gr. and BC 11 gr. 30 m. given, the side AC will be found 27 gr. 54 m.

Or in the rectangle ZRP , having ZP 38 gr. 30 m. and ZR 19 gr. 1 m. given, the side RP will be found 34 gr. 7 m.

P 20

3 *To find a side by knowing the two oblique angles.*

As the sine of either angle
to the sine of the complement of the other angle:
So is the Radius
to the sine of the complement of the side opposite
to the second angle.

So in the rectangle ACB , having CAB for the first angle $23\text{ gr. }30\text{ m.}$ and ABC for the second $69\text{ gr. }21\text{ m.}$ the side AC will be found $27\text{ gr. }54\text{ m.}$ Or making ABC the first angle, and CAB the second, the side BC will be found $11\text{ gr. }30\text{ m.}$

4 *To find the base by knowing both the sides.*

As the Radius
to the sine of the complement of the one side:
So the sine of the complement of the other side,
to the sine of the complement of the base required.

So in the rectangle ACB having $AC\ 27\text{ gr. }54\text{ m.}$ and $BC\ 11\text{ gr. }30\text{ m.}$ the base AB will be found 30 gr.

5 *To find the base by knowing the one side, and the angle opposite to that side.*

As the sine of the angle given,
to the sine of the side given:
So is the Radius
to the sine of the base required.

So in the rectangle BCD , knowing the latitude and the declination, we may find the amplitude, as having BC the side of the declination $11\text{ gr. }30\text{ m.}$ and BDC the angle of the complement of the latitude $38\text{ gr. }15\text{ m.}$ the base BD which is the amplitude, will be found to be $18\text{ gr. }47\text{ m.}$

6 To find an angle by the other oblique angle, and the side opposite to the inquired angle.

As the Radius

to the sine of the complement of the side:

So the sine of the angle giuen,

to the sine of the complement of the angle required.

So in the rectangle ACB , hauing the angle BAC $23\text{ gr. } 30\text{ m.}$ and the side AC $27\text{ gr. } 54\text{ m.}$ the angle ABC will be found $69\text{ gr. } 21\text{ m.}$

7 To finde an angle by the other oblique angle, and the side opposite to the inquired angle.

As the sine of the complement of the side

to the side of the complement of the angle giuen:

So is the Radius

to the sine of the angle required.

So in the rectangle ACB , hauing BAC $23\text{ gr. } 30\text{ m.}$ and BC $11\text{ gr. } 30\text{ m.}$ the angle ABC will be found $69\text{ gr. } 21\text{ m.}$

8 To finde an angle by the base, and the side opposite to the inquired angle.

As the sine of the base

is to the Radius:

So the sine of the side

to the sine of the angle required.

So in the rectangle BCD , hauing BD $18\text{ gr. } 47\text{ m.}$ and BC $11\text{ gr. } 30\text{ m.}$ the angle BDC will be found $38\text{ gr. } 15\text{ m.}$

These eight Propositions haue been wrought by the *sines* alone; those which follow require ioynt help of the *tangent*.

And forasmuch as the *tangent* could not well be extended beyond $63\text{ gr. } 30\text{ m.}$ I shall set downe two wayes for the resolution of each Proposition; if the one will not hold, the other may.

9 To finde a side by having the other side, and the angle opposite to the inquired side.

- 1 As the Radius
to the sine of the side given:
So the tangent of the angle,
to the tangent of the side required.
- 2 As the sine of the side given,
is to the Radius:
So the tangent of the complement of the angle,
to the tangent of the complement of the side required.

So in the rectangle ACB, having the right side AC 27 gr. 54 m, and the angle B A C 23 gr. 30 m. the side B C will be found to be 11 gr. 30 m.

10 To find a side, by having the other side, and the angle adjacent next to the inquired side.

- 1 As the tangent of the angle,
to the tangent of the side given:
So is the Radius
to the sine of the side required.
- 2 As the tangent of the complement of the side,
to the tangent of the complement of the angle:
So is the Radius
to the sine of the side required.

This and the like, where the tangent standeth in the first place, are best wrought by parallell entrance. And so in the rectangle B C D, having B C the side of declination 11 gr. 30 m. and B D C the angle of the complement of the latitude 38 gr. 15 m. the side D C, which is the ascensionall difference, will be found 14 gr. 57 m.

By the ascensionall difference is given the time of the Sunnes rising and setting, and length of the day; allowing

an houre for each 15 gr. and 4 minutes of time for each seuerall degree. As in the example the difference betweene the Sunnes ascension in a right sphere, which is alwayes at 6 of the clocke, and his ascension in our latitude being 14 gr. 57 m. it sheweth that the Sunne riseth very neare an houre before 6, because of the Northerne declination; or after 6, if the Sunne be declining to the Southward.

II To finde a side by knowing the base, and the angle adjacent next to the inquired side.

- 1 As the Radius
to the sine of the complement of the angle:
So is the tangent of the base,
to the tangent of the side required.
- 2 As the sine of the complement of the angle
is to the Radius:
So the tangent of the complement of the base,
to the tangent of the complement of the side required.

So in the rectangle A C B, knowing the place of the Sun from the next equinoctiall point, and the angle of his greatest declination, we may find his right ascension: viz. the base A B 30 gr. and the angle B A C 23 gr. 30 m. being giuen, the right ascension A C will be found 27 gr. 54 m.

12 To finde the base by knowing the oblique angles.

- As the tangent of the one angle,
to the tangent of the complement of the other angle:
So is the Radius
to the sine of the complement of the base.

So in the rectangle A C B, hauing B A C 23 gr. 30 m. and A B C 69 gr. 21 m. the base A B will be found 30 gr.

13 *To finde the base, by one of the sides, and the angle adjacent next that side.*

1 As the Radius
is to the sine of the complement of the angle:
So the tangent of the complement of the side,
to the tangent of the complement of the base.

2 As the sine of the complement of the angle
is to the Radius:
So the tangent of the side giuen,
to the tangent of the base required.

So in the rectangle ACB , hauing AC 27 gr. 54 m. and BAC 23 gr. 30 m. the base AB will be found 30 gr. 0 m.

14 *To find an angle, by knowing both the sides.*

1 As the Radius
is to the sine of the side next the inquired angle:
So the tangent of the complement of the opposite side,
to the tangent of the complement of the angle required.

2 As the sine of the side next the inquired angle,
is to the Radius:
So the tangent of the opposite side,
to the tangent of the angle required.

So in the rectangle ACB , hauing AC 27 gr. 54 m. and BC 11 gr. 30 m. the angle at A will be found 23 gr. 30 m. and the angle at B 69 gr. 21 m.

15 *To finde an angle, by the base, and the side adjacent to the inquired angle.*

1 As the tangent of the complement of the side,
to the tangent of the complement of the base:

M

So

So is the Radius
to the sine of the complement of the angle required.

- 2 As the tangent of the base,
to the tangent of the side:

So is the Radius,
to the sine of the complement of the angle required.

So in the rectangle BCD, having the base BD 18 gr. 47 m.
and the side BC 11 gr. 30 m, the angle D B C between them
will be found 53 gr. 15 m.

16 *To find an angle, by knowing the other
oblique angle, and the base.*

- 1 As the Radius,
to the sine of the complement of the base:
So the tangent of the angle given,
to the tangent of the complement of the angle required.

- 2 As the sine of the complement of the base,
is to the Radius:
So the tangent of the complement of the angle given,
to the tangent of the angle required.

So in the rectangle A C B, having the angle at A 23 gr.
30 m. and the base A B 30 gr. the angle A B C will be found
69 gr. 21 m.

These sixteen cases are all that can fall out in a rectangle
triangle: those which follow do hold

In any sphericall triangle whatsoever

- 17 *To find a side opposite to an angle given, by knowing
one side, and two angles, whereof one is op-
posite to the given side, the other
to the side required.*

As the sine of the angle opposite to the side giuen,
is to the sine of that side giuen :

So the sine of the angle opposite to the side required,
to the sine of the side required.

So in the triangle ABE , hauing the place of the Sunne,
the latitude, and the greatest declination, we may finde the
amplitude. As hauing AB 30 gr. BAE 23 gr. 30 m. and AEB
38 gr. 15 m. the side BE which is the amplitude, will be
found 18 gr. 47 m.

18 *To finde an angle opposite to a side giuen, by hauing
one angle and two sides, the one opposite to
the giuen angle, the other to
the angle required.*

As the sine of the side opposite to the angle giuen,
is to the sine of that angle giuen:

So the sine of the side opposite to the angle required,
to the sine of the angle required.

So in the triangle ZPS , hauing the azimuth, and lati-
tude, and declination, we may find the houre of the day. As
hauing PZS 130 gr. 3 m. PS 70 gr. and ZS 40 gr. the an-
gle ZPS , which sheweth the houre from the meridian shall
be found 31 gr. 34 m.

19 *To find an angle by knowing the three sides.*

This proposition is most vsfull, but most difficult of all
others: as in Arithmetique, so by the *Seller*, yet may it be per-
formed seuerall wayes.

1 According to *Regiomontanus* and others.

As the sine of the lesser side next the angle required,
to the difference of the versed sines of the base and diffe-
rence of the sides:
So is the Radius
to a fourth proportionall.

M 2

Then

Then as the sine of the greater side next the angle required
is to that fourth proportionall :

So is the Radius
to the versed sine of the angle required.

So in the triangle ZPS , having the side PS , the complement of the declination $70\text{ gr. }0\text{ m.}$ the side ZP the complement of the latitude $38\text{ gr. }30\text{ m.}$ and the base ZS the complement of the altitude 40 gr. the angle of the houre of the day ZPS will be found $31\text{ gr. }34\text{ m.}$ which is $2\text{ h. }6\text{ m.}$ from the meridian.

For the base being $40\text{ gr. }0\text{ m.}$ and the difference of the sides $38\text{ gr. }30\text{ m.}$ and $70\text{ gr. }0\text{ m.}$ being $31\text{ gr. }30\text{ m.}$ the difference of their versed sines will be the same with the distance between the right sine of 50 gr. and $58\text{ gr. }30\text{ m.}$ This difference I take out, and make it a parallell sine of the lesser side $38\text{ gr. }30\text{ m.}$ so the parallell Radius will be the fourth proportionall. Then coming to the second operation, I make this fourth proportionall a parallel sine of the greater side of $70\text{ gr. }0\text{ m.}$ and take out his parallell Radius. For this measured from 90 gr. toward the center, will be the versed sine of $31\text{ gr. }34\text{ m.}$

In the like sort in the same triangle ZPS , having the same complements given, the angle PZS which is the azimuth from the North part of the meridian, will be found $130\text{ gr. }3\text{ m.}$ For here the base opposite to the angle required being 70 gr. and the difference of the sides $38\text{ gr. }30\text{ m.}$ and 40 gr. being $1\text{ gr. }30\text{ m.}$ the difference of their versed sines will be the same with the distance between the right sines of 20 gr. and $88\text{ gr. }30\text{ m.}$ This difference I take, and make it a parallell sine of the lesser side $38\text{ gr. }30\text{ m.}$ so the parallell Radius will be the fourth proportionall. Then coming to the second operation, I make this fourth proportionall a parallell sine of the greater side 40 gr. and take out his parallell Radius. For this measured from 90 gr. beyond the center in the lines of sines stretched forth at their full length, will be the versed sine of $130\text{ gr. }3\text{ m.}$

2 I may finde an angle by knowing three sides, by that which I haue elsewhere demonstrated vpon *Barth. Pitiscus*,
and

and that at one operation in this manner.

As the sine of the greater side

is to the secant of the complement of the other side:

So the difference of sines of the complement of the base;
and the arke compounded of the lesser side with the
complement of the greater,
to the versed sine of the angle required.

So in the same triangle ZPS , having the same complements giuen, the angle at P , which sheweth the houre from the meridian, will be found as before $31\text{ gr. }34\text{ m.}$

For the sides being $38\text{ gr. }30\text{ m.}$ and $70\text{ gr. }0\text{ m.}$ I take the secant of the complement of $38\text{ gr. }30\text{ m.}$ and make it a parallell sine of 70 gr. ; then keeping the Sector at this angle, I consider that the complement of 70 gr. being 20 gr. added vnto $38\text{ gr. }30\text{ m.}$ the compounded side (which is here the meridian altitude) will be $58\text{ gr. }30\text{ m.}$; and that the base being 40 gr. the difference of sines of the compounded side and the complement of the base will be (as before) the distance between the sines of 50 gr. and $58\text{ gr. }30\text{ m.}$ Wherefore I take out this difference, and lay it on both the lines of sines from the center: so the parallell taken in the termes of this difference, and measured from 90 gr. toward the center, doth giue the versed sine of $31\text{ gr. }34\text{ m.}$

The other angles PZS , PSZ , may be found in the same sort; but hauing the sides and one angle, it will be sooner done by that which we shewed before in the 18. *Prop.*

20 To find a side by knowing the three angles.

If for the greater angle we take his complement to 180 gr. the angles shall be turned into sides, and the sides into angles, & the operation shall be the same, as in the former *Prop.*

21 To finde a side, by hauing the other two sides,
and the angle comprehended.

This proposition being the conuerse of the nineteenth,

may be wrought accordingly; but the best way both for it and those which follow, is to resolve them into two rectangles, by letting downe a perpendicular, as was shewed in the first *Prop.*

So in the triangle ZPS , having ZP the complement of the latitude, and PS the complement of the declination, with ZPS the angle of the houre from the meridian, we may find ZS the complement of the altitude of the Sunne.

For having let downe the perpendicular ZR by the first *Prop.* we have two triangles, ZRP , ZRS , both rectangle at R . Then may we find the side PR , either by the second, or tenth, or eleventh *Prop.*; which taken out of PS , leaueth the side RS : with this RS and ZR we may find the base ZS by the fourth *Prop.*

Or having let downe the perpendicular SM , we have two rectangle triangles SMZ , $SM P$. Then may we find MP , from which if we take ZP , there remaineth MZ : but with MZ and SM , we may find the base ZS .

22 To find a side, by having the other two sides, and one of the angles next the inquired side.

So in the triangle ZPS , having ZP the complement of the latitude, and PS the complement of the declination, with PZS the angle of the azimuth, we may finde ZS the complement of the altitude of the Sunne.

For having ZP , and the angle at Z , we may to SZ produced, let downe a perpendicular PV . Then we have two rectangle triangles, PVZ , PVS , wherein if we find the sides VZ , VS , and take the one out of the other, there will remain the side inquired ZS .

23 To find a side, by having one side, and the two angles next the inquired side.

So in the triangle ABD , having AB the place of the sun, and BAD the angle of the greatest declination, and ADB the

the angle of the equator with the horizon, we may find AD the oblique ascension.

For hauing let downe BC the perpendicular of declination, we haue two rectangle triangles, ACB , DCB . Then may we find AC the right ascension, and DC the ascensionall difference; and comparing the one with the other, there remaineth AD .

24 To find a side, by hauing two angles, and the side inclosed by them.

So in the triangle ZPS , hauing the angles at Z and P , with the side intercepted ZP , we may find the side PS . For hauing let downe the perpendicular PV , we haue two rectangles PVZ , PVS . Then may we find the angle VPZ , ei- by the seuenth, or fifteenth, or sixteenth *Prop.* which added to ZPS , maketh the angle VPS : with this VPS and PV , we may find the base PS , according to the 13 *Prop.*

25 To find an angle by hauing the other two angles and the side inclosed by them.

So in the triangle ZPS , hauing the angles at Z and P , with the side intercepted ZP , we may finde the other angle ZSP . For hauing let downe the perpendicular ZR , we haue two rectangles ZRP , ZRS . Then may we finde the angle PZR by the sixteenth *Prop.* and that compared with PZS , leaueth the angle RZS : with this RZS and ZR we may find the angle required ZSR , according to the sixth *Prop.*

26 To finde an angle, by hauing the other two angles, and one of the sides next the inquired angle.

So in the triangle ABD , hauing the angles at A and D , with the side AB , we may find the angle ABD . For hauing let downe the perpendicular BC , we haue two rectangles,
 ACB ,

ACB, DCB . Then may we find the angles ABC, DBC , and take DBC out of ABC ; for so there remaineth the angle required ABD .

27 *To finde an angle, by knowing two sides, and the angle contained by them.*

So in the triangle ZPS , hauing the sides ZP, PS , with the angle comprehended ZPS , we may find the angle PZS . For hauing let downe the perpendicular SM , we haue two rectangles SMZ, SMP . Then may we find the side MP , and taking ZP out of MP , there remaineth MZ : with this MZ and the perpendicular MS , we may finde the angle MZS , by the fourteenth *Prop.* This angle MZS , taken out of 180 gr. there remaineth PZS .

28 *To find an angle by knowing the two sides next it, and one of the other angles.*

So in the triangle ZPS , hauing the sides ZP and PS , with the angle PZS , we may find the angle ZPS . For hauing let downe the perpendicular PV , we haue two rectangles PVZ, PVS . Then may we find the angles VPZ, VPS , and taking VPZ out of VPS , there remaineth ZPS , which was required.

These 28 cases are all that can fall out in any spherickall triangle: if any do not presently vnderstand them, let them once more reade ouer the vse of the globes, and they shall soone become ealie vnto them.

CHAP. VI.

Of the use of the Meridian line
in Navigation.

THe *Meridian* line is here set on the side of the *Sector*, stretched forth at full length, on the same plane with the line of *lines* and *Solids*, and is divided vnequally toward 87 gr. (whereof 70 gr. are about one halfe) in such sort as the Meridian in the cart of *Mercators* projection. The use of it may be

1 To diuide a sea-chart according to Mercators projection.

If a degree of the equator on the sea-chart be equall to the hundred part of the line of *lines* in the *Sector*, the degrees of the *Meridian* vpon the *Sector*, shall giue the like degrees vpon the sea-chart: if otherwise they be vnequall, then may the meridians of the sea-chart be diuided in such sort as the line of *Meridians* is diuided on the *Sector*, by that which we shewed before in the 8. *Prop.* of the line of *lines*.

But to auoid error, I haue here set downe a Table, whereby the Meridian line may be diuided out of the degrees of the equator, supposing each degree to be subdivided into a thousand parts. By which Table, & the vsuall Table of *Sines*, *Tangents* and *Secants*, the proportions following may be also resolued arithmetically. For the maner of diuision, let the equator (or one of the parallels if it be a particular chart) be drawne, and diuided, and crossed with parallell meridians, as in the common sea-chart: then looke into the Table, and let the distance of 40 gr. in the meridian, from the equator, be equall to 43 gr. 711 parts of the equator; let 50 gr. in the meridian from the equator, be equall to 57 gr. 929 parts of the equator; and so in the rest.

M	Gr	Par	M	Gr	Par	M	Gr	Par	M	Gr	Par	M	Gr	Par
0	0	0	3	3	001	6	6	011	9	9	037	12	12	088
		100		3	101			6111		9	138		12	190
		200		3	201			6212		9	239		12	293
		300		3	301			6312		9	341		12	395
		400		3	402			6413		9	442		12	497
		500		3	502			6514		9	543		12	600
		600		3	602			6614		9	645		12	702
		70		3	702			6715		9	746		12	805
		800		3	803			6816		9	848		12	907
		900		3	903			6916		9	949		13	010
1	1	000	4	4	003	7	7	017	10	10	051	13	13	112
	1	100		4	103			7118		10	152		13	215
	1	200		4	204			7219		10	254		13	318
	1	300		4	304			7319		10	355		13	421
	1	400		4	404			7420		10	457		13	523
	1	500		4	504			7521		10	559		13	626
	1	600		4	605			7622		10	661		13	729
	1	700		4	705			7723		10	762		13	832
	1	800		4	805			7824		10	864		13	935
	1	900		4	906			7925		10	966		14	038
2	2	000	5	5	006	8	8	026	11	11	068	14	14	141
	2	100		5	106			8127		11	170		14	244
	2	200		5	207			8228		11	272		14	347
	2	300		5	307			8329		11	374		14	450
	2	400		5	408			8430		11	476		14	553
	2	500		5	508			8531		11	578		14	656
	2	601		5	609			8632		11	680		14	760
	2	701		5	709			8733		11	782		14	863
	2	801		5	810			8834		11	884		14	967
	2	901		5	910			8936		11	986		15	070
3	3	001	6	6	011	9	9	037	12	12	088	15	15	174

of the Meridian line.

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<i>M Gr Par</i>	<i>M Gr Par</i>	<i>M Gr Par</i>	<i>M Gr Par</i>	<i>M Gr Par</i>
15 15 174	18 18 303	21 21 486	24 24 734	27 27 058
15 277	18 408	21 593	24 844	28 171
15 381	18 513	21 701	24 953	28 283
15 485	18 619	21 808	25 063	28 396
15 588	18 724	21 915	25 173	28 508
15 692	18 830	21 023	25 282	28 621
15 796	18 935	22 130	25 392	28 734
15 900	19 041	22 238	25 502	28 847
16 004	19 146	22 345	25 613	28 959
16 107	19 251	22 453	25 723	29 072
16 16 211	19 356	22 561	25 833	29 186
16 316	19 463	22 669	25 943	29 299
16 420	19 569	22 777	26 054	29 413
16 524	19 675	22 885	26 164	29 526
16 628	19 781	22 993	26 275	29 640
16 732	19 887	23 101	26 386	29 753
16 836	19 993	23 210	26 497	29 867
16 941	20 100	23 318	26 608	29 981
17 045	20 206	23 427	26 719	30 095
17 150	20 312	23 535	26 830	30 300
17 17 255	20 419	23 643	26 941	30 324
17 359	20 525	23 752	27 052	30 438
17 464	20 632	23 861	27 164	30 553
17 568	20 738	23 970	27 275	30 667
17 673	20 845	24 079	27 387	30 782
17 778	20 952	24 188	27 499	30 897
17 883	21 059	24 297	27 610	31 012
17 988	21 165	24 406	27 722	31 127
18 093	21 272	24 515	27 834	31 242
18 198	21 379	24 624	27 946	31 357
18 18 303	21 21 486	24 24 734	27 28 058	30 31 473

M	Gr	Par	M	Gr	Par	M	Gr	Par	M	Gr	Par	M	Gr	Par
30	31	473	33	34	992	36	38	633	39	42	415	42	46	362
	31	588		35	111		38	757		42	544		46	496
	31	704		35	231		38	880		42	673		46	631
	31	820		35	350		39	004		42	802		46	766
	31	936		35	470		39	129		42	931		46	902
	32	052		35	590		39	253		43	061		47	037
	32	168		35	710		39	377		43	191		47	173
	32	284		35	830		39	502		43	320		47	309
	32	40		35	950		39	627		43	451		47	445
	32	517		36	071		39	752		43	581		47	581
31	32	633	34	36	191	37	39	877	40	43	711	43	47	718
	32	750		36	312		40	002		43	842		47	855
	32	867		36	433		40	128		43	973		47	992
	32	984		36	554		40	253		44	104		48	129
	33	101		36	675		40	379		44	235		48	267
	33	218		36	796		40	505		44	366		48	404
	33	336		36	917		40	631		44	498		48	542
	33	453		37	039		40	757		44	630		48	681
	33	571		37	161		40	884		44	762		48	819
	33	688		37	283		41	011		44	894		48	958
32	33	806	35	37	405	38	41	137	41	45	026	44	49	097
	33	924		37	527		41	264		45	159		49	236
	34	042		37	649		41	392		45	292		49	375
	34	161		37	771		41	519		45	425		49	515
	34	279		37	894		41	646		45	558		49	655
	34	397		38	017		41	774		45	691		49	795
	34	516		38	140		41	902		45	825		49	935
	34	635		38	263		42	030		45	959		50	076
	34	754		38	386		42	158		46	093		50	217
	34	873		38	509		42	287		46	227		50	358
33	34	992	36	38	633	39	42	415	42	46	362	45	50	499

of the Meridian line.

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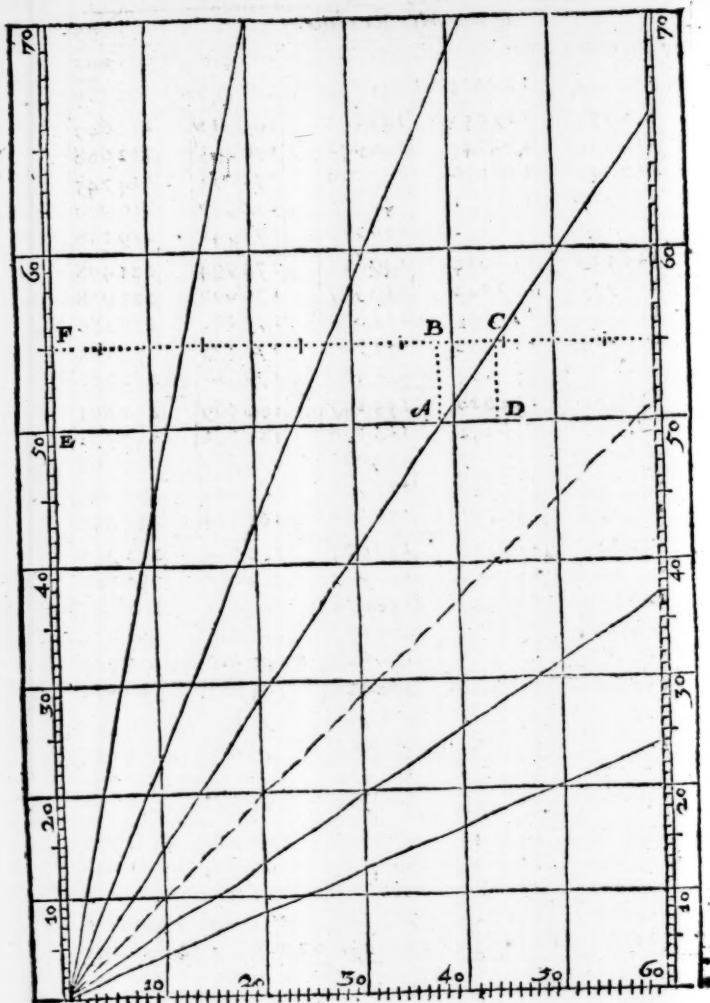
M Gr Par	M Gr Par	M Gr Par	M Gr Par	M Gr Par
45 50 499	48 54 860	51 59 481	54 64 412	57 69 711
50 641	55 010	59 640	64 582	69 895
50 783	55 160	59 800	64 753	70 080
50 925	55 310	59 960	64 924	70 263
51 068	55 460	60 120	65 096	70 449
51 210	55 611	60 280	65 268	70 635
51 353	55 762	60 441	65 440	70 821
51 496	55 913	60 602	65 613	71 008
51 639	56 065	60 763	65 786	71 195
51 783	56 217	60 925	65 960	71 383
46 51 927	49 56 369	52 61 088	55 66 134	58 71 572
52 071	56 522	61 250	66 308	71 761
52 215	56 675	61 413	66 483	71 950
52 360	56 828	61 577	66 659	72 140
52 505	56 981	61 740	66 835	72 331
52 650	57 135	61 904	67 011	72 522
52 795	57 289	62 069	67 188	72 714
52 941	57 444	62 234	67 365	72 906
53 087	57 598	62 399	67 543	73 099
53 233	57 754	62 564	67 721	73 292
47 53 380	50 57 909	53 62 730	56 67 900	59 73 486
53 526	58 065	62 897	68 079	73 680
53 673	58 221	63 063	68 258	73 875
53 821	58 377	63 231	68 438	74 071
53 968	58 534	63 398	68 618	74 267
54 116	58 691	63 566	68 799	74 464
54 264	58 848	63 734	68 981	74 661
54 413	59 006	63 903	69 163	74 859
54 562	59 164	64 072	69 345	75 057
54 711	59 322	64 242	69 528	75 256
48 54 860	51 59 481	54 64 412	57 69 711	60 75 456

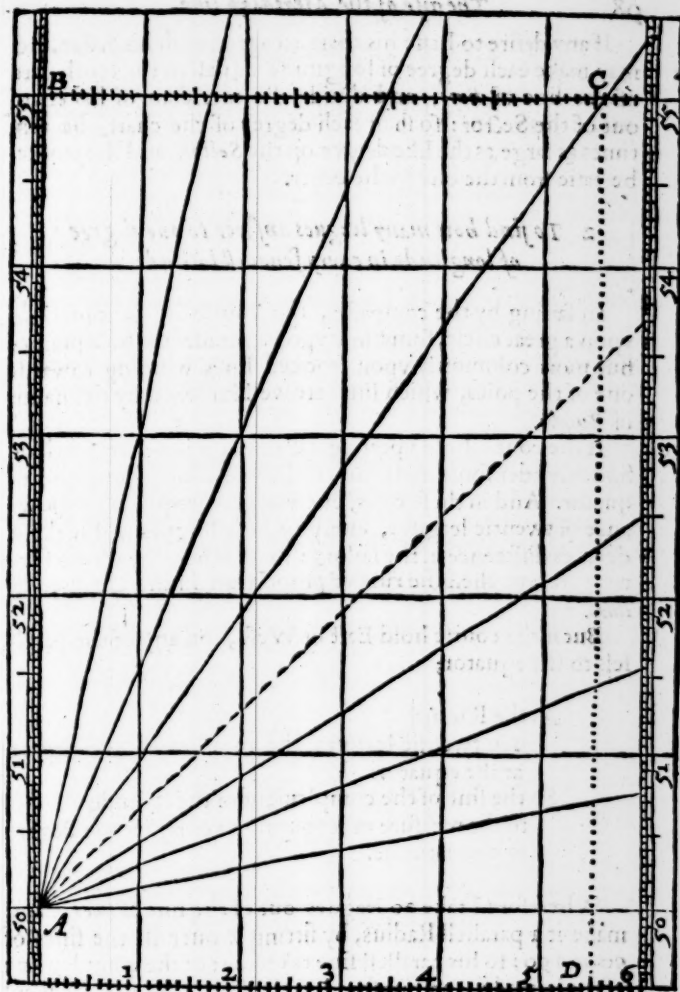
M	Gr	Par	M	Gr	Par	M	Gr	Par	M	Gr	Par	M	Gr	Par
60	75	456	63	81	749	66	88	725	69	96	575	72	105	579
	75	656		81	970		88	921		96	854		105	904
	75	857		82	191		89	219		97	135		106	230
	76	059		82	413		89	467		97	418		106	558
	76	261		82	635		89	716		97	701		106	888
	76	464		82	860		89	967		97	986		107	220
	76	667		83	084		90	218		98	272		107	553
	76	871		83	310		90	470		98	560		107	888
	77	076		83	536		90	723		98	849		108	226
	77	281		83	763		90	978		99	139		108	565
61	77	487	64	83	990	67	91	232	70	99	431	73	108	906
	77	694		84	219		91	489		99	724		10	249
	77	901		84	448		91	746		100	018		109	594
	78	109		84	678		92	005		100	314		109	941
	78	317		84	909		92	264		100	612		110	290
	78	526		85	141		92	525		100	910		110	641
	78	736		85	374		92	787		101	211		110	994
	78	947		85	607		93	050		101	513		111	349
	79	158		85	842		93	314		101	816		111	707
	79	370		86	077		93	579		102	121		112	066
62	79	583	65	86	313	68	93	846	71	102	427	74	112	428
	79	796		86	550		94	113		102	735		112	792
	80	010		86	788		94	382		103	044		113	158
	80	225		87	027		94	652		103	356		113	526
	80	441		87	267		94	923		103	668		113	897
	80	657		87	508		95	195		103	983		114	270
	80	874		87	749		95	468		104	299		114	645
	81	091		87	992		95	743		104	616		115	023
	81	310		88	235		96	019		104	936		115	403
	81	529		88	480		96	296		105	257		115	786
63	81	749	66	88	725	69	96	575	72	105	579	75	116	171

of the Meridian line.

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M	Gr.	Par	M	Gr.	Par	M	Gr.	Par	M	Gr.	Par	M	Gr.	Par	M	Gr.	Par
75	116	171	78	129	075	81	145	650	84	168	947	87	208	705			
	113	559		129	558		146	292		169	912		210	649			
	116	949		130	045		146	942		170	893		212	668			
	117	342		130	536		147	600		171	891		214	745			
	117	737		131	031		148	265		172	907		216	909			
	118	135		131	530		148	937		173	941		219	158			
	118	536		132	034		149	618		174	994		221	498			
	118	939		132	542		150	307		176	067		223	938			
	119	345		133	055		151	003		177	160		226	486			
	119	755		133	572		151	709		178	275		229	153			
76	120	160	79	134	094	82	152	423	85	179	411	88	231	950			
	120	581		134	620		153	147		180	569		234	891			
	121	000		135	151		153	878		181	752		237	991			
	121	420		135	687		154	620		182	960		241	268			
	121	843		136	228		155	372		184	194		244	744			
	122	270		136	775		156	132		185	454		248	445			
	122	700		137	326		156	903		186	743		252	402			
	123	133		137	883		157	685		188	062		256	652			
	123	570		138	445		158	478		189	411		261	243			
	124	009		139	012		159	281		190	793		266	235			
77	124	452	80	139	585	83	160	096	86	192	210	80	271	705			
	124	898		140	164		160	922		193	661		277	753			
	125	348		140	748		161	761		195	151		284	517			
	125	801		141	339		162	612		196	680		292	191			
	126	258		141	936		163	475		198	251		301	058			
	126	718		142	538		164	352		199	867		311	563			
	127	182		143	147		165	242		201	529		324	455			
	127	649		143	763		166	146		203	240		341	166			
	128	121		144	385		267	065		205	005		365	039			
	128	596		145	014		167	999		206	825		408	011			
78	129	075	81	145	650	84	168	947	87	208	705	90	Infinite				





If any desire to haue his chart to agree with his *Sector*, he may make each degree of longitude equall to the tenth part of the line of *lines*, and diuide the meridian of his chart out of the *Sector*: so shall each degree of the chart, be ten times as large as the like degree on the *Sector*, and the worke be easie from the one to the other.

2 *To find how many leagues answer to one degree of longitude in euery seuerall latitude.*

In sailing by the compasse, the course holds sometime vpon a great circle, sometime vpon a parallel to the equator; but most commonly vpon crooked lines winding towards one of the poles, which lines are well knowne by the name of *Rumbs*.

If the course hold vpon a great circle, it is either North or South, vnder some meridian, or East or West, vnder the equator. And in these cases, euery degree requires an allowance of twentie leagues, euery twentie leagues will make a degrees difference in the sailing: so that here needs no further precept then the rule of proportion in the Chapter of *lines*.

But if the course hold East or West, on any of the parallels to the equator,

As the Radius

is to twentie leagues, the measure of one degree at the equator:

So the sine of the complement of the latitude to the measure of leagues answering to one degree in that latitude.

Wherefore I take 20 leagues out of the line of *lines*, and make it a parallel Radius, by fitting it ouer in the sines of 90 and 90: so his parallel line taken out of the complement of the latitude, and measured in the line of *lines*, shall shew the number of leagues required.

Thus

Thus in the latitude of 18 gr. 12 m. we shall find 19 leagues answering to one degree of longitude, and 18 leagues in the latitude of 25 gr. 15 m. and as in this Table.

This may be done more readily without opening the Sector, by doubling the sine of the complement of the latitude, as may appear in the same example.

It may also be done by the line of meridians, either vpon the Sector, or vpon the chart. For if we open a paire of compasses to the quantitie of one degree of longitude in the equator, and measure it in the meridian line, setting one foot as much above the latitude given, as the other falleth beneath it, so that the latitude may be in the middle betweene the feet of the compasses, the number of leagues intercepted shall be that which was required.

But if the course hold vpon any of the *rumb*s, betweene a parallell of the equator and the meridian, we are to consider besides the quarter of the world to which we tend, which must be alwayes knowne.

Gr.	1	2
0	0	20
18	12	19
25	15	18
31	18	17
36	21	16
41	25	15
45	34	14
49	28	13
53	8	12
56	38	11
60	0	10
63	15	9
66	25	8
69	30	7
72	32	6
75	31	5
78	28	4
81	23	3
84	15	2
87	8	1

- 1 The difference of longitude at least in generall.
- 2 The difference of latitude, and that in particular.
- 3 The *rumb* whereon the course holds.
- 4 The distance vpon the *rumb*, which is the distance, which we are here to consider, and is alwayes somewhat greater then the like distance vpon a greater circle. And for these first I shew in generall this third Prop.

3 To finde how many leagues do answer to one degree of latitude in every severall *Rumb*.

As the sine of the complement of the *rumb* frō the meridian, is to 20 leagues the measure of one degree at the meridian: So the Radius

to the leagues answering to one degree vpon the *Rumb*.

Wherefore I take 20 leagues out of the line of *lines*, and make it a parallell line of the complement of the Rumb from the meridian; so his parallell Radius taken and measured in the line of *lines*, shall shew the number of leagues required.

Thus in the first Rumb from the meridian, we shall finde 20 *lgs* 39 *parts* answering to one degree of latitude, and 21 *lgs* 65 *parts* in the second Rumb, &c. as in this Table, where we subdiuide each league into a hundred parts, and shew besides what inclination the rumb hath to the meridian.

This may be done more readily without opening the *Sector*, by doubling the secant of the latitude, as may appeare in the same example.

It may also be done vpon the chart, if we take the distance vpon the Rumb betweene two parallels, and measure it in the meridian line, as farre aboue the greater latitude as beneath the lesser. For so the number of leagues intercepted, shall be that which was required.

This considered in generall, I shew more particularly in twelue *Prop.* following, how of these foure any two being giuen, the other two may be found, both by *Mercators* chart, and by this *Sector*.

Rumb in Quint.	Inclina- tio to the Meridian		Number of <i>lgs</i> .	
	Gr.	Ms.	Lgs	Par
1	2	49	20	02
	5	37	20	10
	8	26	20	22
1	11	15	20	39
2	14	4	20	62
	16	52	20	90
	19	41	21	24
	22	30	21	65
3	25	19	22	12
	28	7	22	68
	30	56	23	32
	33	45	24	05
4	36	34	24	90
	39	22	25	87
	42	11	26	99
	45	0	28	28
5	47	49	29	78
	50	37	31	52
	53	26	33	97
	56	15	36	00
6	59	4	38	90
	61	52	42	43
	64	41	36	78
	67	30	52	26
7	70	19	59	37
	73	7	68	90
	75	56	82	31
	78	45	102	52
8	81	34	136	30
	84	22	205	24
	87	11	407	60
8	90	0	Infinita.	

I By

**1 By one Latitude Rumb and distance to find
the difference of latitudes.**

As the Radius

to the sine of the complement of the Rumb from the me-
So the distance vpon the Rumb, (ridian:
to the difference of latitudes.

Let the place giuen be *A* in the latitude of 50 gr. *C* in a greater latitude, but vnknowne, the distance vpon the Rumb being 6 gr. betweene them, and the Rumb the third from the meridian.

First I take 6 gr. for the distance vpon the Rumb, out of the line of *sines*, and make it a parallell Radius, by putting it ouer in the lines of 90 and 90. Then keeping the *Sextor* at this angle, I take out the parallell line of 56 gr. 15 m. which is the sine of the complement of the third Rumb from the meridian, and measuring it in the line of *sines*, I find it to be 5 gr. and such is the difference of latitude required.

Or I may take out the sine of 56 gr. 15 m. for the complement of the third Rumb from the meridian, make it a parallell Radius; then keeping the *Sextor* at this angle, I take 6 gr. for the distance, either out of the line of *sines*, or any other scale of equall parts, or else out of the meridian line, and lay it on both sides of the *Sextor* from the center, either on the line of *sines* or *sines*: so the parallell taken from the termes of this distance, and measured in the same scale wherein the distance was measured, shall shew the difference of latitude to be 5 gr. as before.

But in shorter distances, such as fall within the compasse of a dayes sailing, this worke will hold much better. As may appeare by comparing the worke with the Table following: where the numbers in the front do signifie the leagues; the first in the side, the Rumb; and the rest in the middle, the difference of latitude.

Lgr	100	80	60	40	20	19	18	17	16	15
Rum	G.M	G.M	G.M	G.M	M	M	M	M	M	M
	5 0	4 0	3 0	2 0	60	57	54	51	48	45
	4 59	3 59	2 59	1 59	60	57	54	51	48	45
	4 58	3 58	2 59	1 59	60	57	54	51	48	45
	4 56	3 57	2 58	1 58	59	56	53	50	47	44
1	4 54	3 55	2 56	1 57	59	56	53	50	47	44
	4 51	3 53	2 55	1 56	58	56	52	50	47	43
	4 47	3 50	2 52	1 55	57	55	52	49	46	43
	4 42	3 46	2 49	1 53	56	54	51	48	45	42
2	4 37	3 42	2 46	1 51	55	53	50	47	44	41
	4 31	3 37	2 43	1 48	54	52	49	46	43	40
	4 25	3 32	2 39	1 46	53	50	48	45	42	39
	4 17	3 26	2 34	1 43	51	49	46	44	41	38
3	4 10	3 20	2 30	1 40	50	47	45	42	40	37
	4 1	3 13	2 25	1 36	48	46	43	41	39	36
	3 52	3 5	2 19	1 32	46	44	42	39	37	35
	3 42	2 58	2 13	1 28	44	42	40	38	36	33
4	3 32	2 50	2 7	1 25	42	40	38	36	34	32
	3 22	2 41	2 1	1 21	40	38	36	34	32	30
	3 10	2 32	1 54	1 16	38	36	34	32	30	28
	2 59	2 23	1 47	1 12	36	34	32	30	29	27
5	2 47	2 14	1 40	1 7	33	32	30	28	27	25
	2 34	2 3	1 32	1 2	31	29	28	26	25	23
	2 22	1 53	1 25	0 57	28	27	25	24	23	22
	2 8	1 43	1 17	0 52	26	24	23	22	21	19
6	1 55	1 32	1 8	0 46	23	22	21	20	18	17
	1 41	1 20	1 0	0 40	20	19	18	17	16	15
	1 27	1 9	0 52	0 35	17	16	16	15	14	13
	1 13	0 58	0 44	0 30	15	14	13	12	12	11
7	0 59	0 47	0 35	0 24	12	11	11	10	9	9
	0 44	0 36	0 26	0 18	9	8	8	7	7	7
	0 30	0 24	0 18	0 12	6	6	5	5	5	4
	0 15	0 12	0 9	0 9	3	3	3	3	2	2
8	0 0	0 0	0 0	0 0	0	0	0	0	0	0

and difference of latitudes.

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[illegible]

In the Chart let a meridian AB be drawne through A , and in A with AB make an angle of the Rumb BAC . Then open the compasses, according to the latitude of the places, to EF the quantitie of 6 gr. in the meridian, transferring them into the Rumb from A to C , and through C draw the parallell BC , crossing the meridian AB in B : so the degrees in the meridian from A to B , shall shew the difference of latitude to be 5 gr.

2 By the Rumb and both latitudes to find the distance vpon the Rumb.

As the sine of the complement of the Rumb from the meridian is to the Radius: (dian,
So the difference of latitudes,
to the distance vpon the Rumb.

As if the places giuen were A in the latitude of 50 gr. C in the latitude of 55 gr. and the Rumb the third from the meridian.

Here I may take 5 gr. for the difference of latitude out of the line of *lines*, and put it ouer in the sine of $56\text{ gr. } 15\text{ m.}$ for the complement of the third Rumb from the meridian. Then keeping the *Sector* at this angle, I take out the parallell Radius, and measuring it in the line of *lines*, I find it to be 6 gr. and such is the distance vpon the Rumb, which was required.

Or I may take the laterall Radius, and make it a parallell sine of $56\text{ gr. } 15\text{ m.}$ the complement of the Rumb from the meridian: then keeping the *Sector* at this angle, I take 5 gr. for the difference of latitude, either out of the line of *lines*, or out of some other scale of equall parts, and lay it on both sides of the *Sector* from the center, either on the line of *lines* or of *sines*: so the parallell taken from the termes of this difference, and measured in the same scale with the difference, shall shew the distance vpon the Rumb to be 6 gr. or 120 leagues.

Or

Or keeping the *Sector* at this angle, I may take the difference betweene 50 *gr.* and 55 *gr.* out of the *Meridian* line, and measuring it in the equator, I shall find it to be equall to 8 *gr.* 22 *p.* of the equator. Wherefore I take the parallell betweene 822 and 822 out of the line of *lines*, and measuring it in the line of *lines* I shall find it to be 989; which shewes that according to this projection, the distance vpon this third Rumb, answerable to the former distance of latitudes, will be equall to 9 *gr.* 89 *p.* of the equator.

Or the *Sector* remaining at this angle, I may take the difference betweene 50 *gr.* and 55 *gr.* out of the *Meridian* line, and lay it from the center on both sides of the *Sector*, either on the line of *lines* or of *sines*: so the parallell taken from the termes of this difference, shall be the very line of distance required, the same with *AC* or *EF* vpon the chart; which may serue for the better pricking downe of the distance vpon the Rumb, without taking it forth of the *Meridian* line, as in the former *Prop.*

Or if the Rumb fall nearer to the equator, that the laterall Radius cannot be fitted ouer in it, this proposition may be wrought by parallell entrance.

For if I first take out the sine of 56 *gr.* 15 *m.* and make it a parallell Radius, by fitting it ouer in the sines of 90 and 90; or in the ends of the lines of *lines*, and then take 5 *gr.* for the difference of latitudes out of the line of *lines*, and carrie it parallell to the former, I shall find it to crosse both lines of *lines* in the points of 6: and so it giues the same distance as before.

Or if the distance be small, it may be found by the former Table. For the Rumb being found in the side of the Table; and the difference of latitude in the same line; the top of the columnne wherein the difference of latitude was found, shall giue the number of leagues in the distance required.

Or we may finde this distance in the Table of Rumbs in the fifth *Prop.* following. For according to the example looke into the Table of the third Rumb for 5 *gr.* of latitude, and there we shall finde 6 *gr.* 01 *parts* vnder the title of distance.

P

So

So if the difference of latitude vpon the same Rumb were 50 *gr.* the distance would be 60 *gr.* 13 *parts*. If the difference of latitude vpon the same Rumb were onely $\frac{1}{2}$ of a degree, the distance would be onely 60 *parts*, such as 100 doe make a degree.

In the chart let a Meridian *AB* be drawne through *A*, and parallels of latitude through *A* and *C*; & then in *A* with *AB* make an angle of the Rumb *BAC*: so the distance taken from *A* to *C*, and measured in the Meridian line, according to the latitude of the places, shall be found to be 6 *gr.* or 120 leagues. And such is the distance required.

3 *By the distance and both latitudes
to find the Rumb.*

As the distance vpon the Rumb,
to the difference of latitudes:

So is the Radius (ridian,
to the sine of the complement of the Rumb from the Me-

As if the places giuen were *A* in the latitude of 50 *gr.* *C* in the latitude of 55 *gr.* the distance betweene them being 6 *gr.* vpon the Rumb. First I take 6 *gr.* for the distance vpon the Rumb, and lay it on both sides of the *Sector* from the center; then out of the same scale I take 5 *gr.* for the difference of latitude, and to it open the *Sector* in the termes of the former distance: so the parallell Radius taken and measured in the *sines*, doth giue 56 *gr.* 15 *m.* the complement whereof 33 *gr.* 45 *m.* is the angle of the Rumb's inclination to the Meridian, which was required.

In the chart let a meridian *AB* be drawne through *A*, and parallels of latitude both through *A* and *C*; then open the compasses according to the latitude of the places to *EF* the quantitie of 6 *gr.* in the meridian, and setting one foote in *A*, turne the other till it crosse the parallell *BC* in *C*, and draw the right line *AC*: so the angle *BAC* shall shew the inclination of the Rumb to the Meridian to be 33 *gr.* 45 *m.* as before.
These

These three last *Prop.* depend one on the other, and may be wrought as truly by the common sea-chart as by this of *Mercators* projection: and therefore in working them by the *Sector*, the distance and the difference of latitudes may as well or better be taken out of the line of *lines* (which here representeth the equator) or any other line of equall parts, as out of the enlarged degrees in the *meridian* line. But in the propositions following, the difference of longitude must be taken out of the equator; the difference of latitudes and distance vpon the *Rumb*, must alwayes be taken out of the *meridian* line; which I therefore call the proper difference, and proper distance.

4 By the longitude and latitude of two places
to find the *Rumb*.

As if the places giuen were *A* in the latitude of 50 gr. *C* in the latitude of 55 gr. and the difference of longitude betwene them were 5 gr. 30 m.

In the chart let meridians and parallels be drawn through *A* and *C*, and a straight line for the *Rumb* from *A* to *C*; then by that we shewed *Cap. 2. Prop. 9.* inquire the quantitie of the angle *BAC*, and it shall be found to be 33 gr. 45 m. which is the third *Rumb* from the Meridian. Wherefore the proportion holds for the *Sector*,

As *AB* the proper difference of latitude,
is to *BC* the difference of longitude:
So *AB* as Radius,
to *BC* the tangent of the *Rumb* from the Meridian.

According to this I take the proper difference of latitude from 50 gr. to 55 gr. out of the line of *meridians*, and lay it on both sides of the *Sector* from the center; then I take the difference of longitude 5 gr. $\frac{1}{2}$ out of the line of *lines*, and to it open the *Sector* in the termes of the former difference of latitudes; so the parallel Radius taken from betwene the 90 and 90, and measured in the greater *tangent* on the side of the *Sector*,

Flor, doth giue 33 *gr*. 45 *m*. for the Rumb required,

But if the Rumb fall nearer to the equator;

As *AD* the difference of longitudes,
is to *DC* the proper difference of latitudes:

So *AD* as Radius,

to *DC*, the tangent of the rumb from the equator.

According to this I take the former difference of latitudes from 50 *gr*. to 55 *gr*. out of the line of *Meridians*, and to it open the *Sector* in the termes of the difference of longitude reckoned in the line of *lines* from the center: so the parallell Radius taken and measured in the *tangent*, doth giue 56 *gr*. 15 *m*. for the rumb from the equator; which is the complement to the former 33 *gr*. 45 *m*: and so both wayes it is found to be the third rumb from the Meridian.

But if this rumb were to be found in the common sea-chart, it should seeme to be about 47 *gr*. which is more then the fourth rumb from the meridian.

5 By the Rumb and both latitudes to find the difference of longitude.

As if the places giuen were *A* in the latitude of 50 *gr*. and *C* in the latitude of 55 *gr*. and the rumb the third from the meridian.

In the chart, let a meridian be drawne through *A*, and a parallell of latitude through *C*; then in *A* with *AB* make the angle of the rumb from the meridian *BAC*, (as was shewed *Cap. 2. Prop. 10.*) So the degrees in the parallell betweene *B* and *C*, shall be found to be 5 *gr*. $\frac{1}{2}$, the difference of longitude which was required. Wherefore the proportion holds for the *Sector*.

As *AB* the Radius,

to *BC* the tangent of the rumb from the meridian:

So *AB* as proper difference of the latitudes,

to *BC* the difference of longitude.

According

According to this we may take the tangent of the Rumb which is here 33 gr. 45 m. from the meridian, out of the greater tangent on the side of the Sector; and putting it over between 90 and 90, make it a Radius: then keeping the Sector at this angle, take the proper difference of latitudes from 50 gr. to 55 gr. out of the line of Meridians, and lay it on both sides of the Sector from the center: so the parallel taken from the termes of this difference, and measured in the line of lines, shall shew the difference of longitude to be 5 gr. $\frac{1}{2}$.

Or if the Rumb fall nearer the equator.

As *DC* the tangent of the Rumb from the equator,
to *AD* the Radius:
So *DC* as proper difference of the latitudes,
to *AD* the difference of longitude.

According to this we may best work by parallell entrance, first taking 56 gr. 15 m. for the angle of the Rumb from the equator, out of the greater tangent, and make it a parallell Radius: then take the proper difference of latitudes out of the line of meridians, and carrie it parallell to the former: so we shall find it to crosse the line of lines in 5 gr. $\frac{1}{2}$. And this is the difference of longitude required, the same as before.

But if this difference were to be found by the common sea-chart, it should seeme to be only 3 gr. 20 m. which is more then 2 gr. lesse then the truth. And yet this error would be greater, if either the latitude be greater, or the Rumb fall nearer the equator: as may appeare by comparing the common sea-chart with the Tables following.

The first Runbe } from the Meridian, }			North and by East, South and by East,			North and by West, South and by West,		
La	Long.	Dist.	La	Long.	Dist.	La	Long.	Dist.
Gr.	Gr. P.	Gr. P.	Gr.	Gr. P.	Gr. P.	Gr.	Gr. P.	Gr. P.
0	0	0	30	6 26	30 54	60	15 01	61 18
1	20	1 02	31	6 49	31 61	61	15 41	62 20
2	40	2 04	32	6 72	32 63	62	15 83	63 21
3	60	3 06	33	6 96	33 65	63	16 26	64 23
4	80	4 08	34	7 20	34 67	64	16 71	65 25
5	1 00	5 10	35	7 44	35 69	65	17 17	66 27
6	1 20	6 12	36	7 68	36 71	66	17 65	67 29
7	1 40	7 14	37	7 92	37 73	67	18 15	68 31
8	1 60	8 16	38	8 17	38 75	68	18 67	69 33
9	1 80	9 18	39	8 43	39 77	69	19 21	70 35
10	2 00	10 20	40	8 70	40 78	70	19 78	71 37
11	2 20	11 22	41	8 96	41 80	71	20 37	72 39
12	2 40	12 24	42	9 22	42 82	72	21 00	73 41
13	2 61	13 25	43	9 50	43 84	73	21 66	74 43
14	2 81	14 27	44	9 76	44 86	74	22 36	75 45
15	3 02	15 29	45	10 04	45 88	75	23 10	76 47
16	3 22	16 31	46	10 33	46 90	76	23 90	77 49
17	3 43	17 33	47	10 62	47 92	77	24 75	78 51
18	3 64	18 35	48	10 91	48 94	78	25 67	79 53
19	3 85	19 37	49	11 21	49 96	79	26 67	80 55
20	4 06	20 39	50	11 52	50 98	80	27 76	81 57
21	4 27	21 41	51	11 83	52 0	81	28 97	82 59
22	4 49	22 43	52	12 15	53 2	82	30 32	83 61
23	4 70	23 45	53	12 47	54 4	83	31 84	84 63
24	4 92	24 47	54	12 81	55 6	84	33 61	85 62
25	5 14	25 49	55	13 16	56 8	85	35 69	86 67
26	5 36	26 51	56	13 50	57 10	86	38 24	87 69
27	5 58	27 53	57	13 86	58 12	87	41 52	88 71
28	5 80	28 55	58	14 23	59 14	88	46 15	89 73
29	6 03	29 57	59	14 62	60 16	89	54 06	90 75
30	6 26	30 59	60	15 01	61 18	90		

The second Runke from the Meridian.			North North-east. South South-east.			North North-west South South-west		
La	Long.	Dist.	La	Long.	Dist.	La	Long.	Dist.
Gr.	Gr. P.	Gr. P.	Gr.	Gr. P.	Gr. P.	Gr.	Gr. P.	Gr. P.
0	0	0	30	13 03	32 47	60	31 25	64 94
1	0 42	1 08	31	13 51	33 51	61	32 09	66 03
2	0 83	2 16	32	14 00	34 64	62	32 96	67 11
3	1 24	3 25	33	14 49	35 72	63	33 86	68 19
4	1 65	4 33	34	15 00	36 80	64	34 79	69 27
5	2 07	5 41	35	15 50	37 88	65	35 71	70 35
6	2 49	6 49	36	16 00	38 97	66	36 75	71 44
7	2 91	7 57	37	16 51	40 05	67	37 80	72 52
8	3 32	8 66	38	17 03	41 13	68	38 88	73 60
9	3 74	9 74	39	17 56	42 21	69	40 00	74 68
10	4 16	10 82	40	18 10	43 30	70	41 19	75 77
11	4 59	11 90	41	18 65	44 38	71	42 43	76 85
12	5 01	12 99	42	19 20	45 46	72	43 74	77 93
13	5 43	14 07	43	19 76	46 54	73	45 11	79 01
14	5 85	15 15	44	20 33	47 62	74	46 57	80 10
15	6 28	16 23	45	20 92	48 71	75	48 12	81 18
16	6 71	17 32	46	21 50	49 79	76	49 78	82 26
17	7 14	18 40	47	22 11	50 87	77	51 55	83 34
18	7 58	19 48	48	22 72	51 95	78	53 46	84 42
19	8 01	20 56	49	23 35	53 03	79	55 54	85 51
20	8 45	21 65	50	23 98	54 12	80	57 82	86 59
21	8 90	22 73	51	24 63	55 20	81	60 33	87 67
22	9 34	23 81	52	25 30	56 28	82	63 13	88 76
23	9 79	24 89	53	25 98	57 37	83	66 32	89 84
24	10 24	25 98	54	26 68	58 45	84	69 99	90 92
25	10 70	27 06	55	27 30	59 53	85	74 32	92 00
26	11 16	28 14	56	28 12	60 61	86	79 63	93 09
27	11 62	29 22	57	28 87	61 70	87	86 46	94 17
28	12 08	30 31	58	29 64	62 78	88	96 10	95 25
29	12 55	31 39	59	30 44	63 86	89	112 57	96 33
30	13 03	32 47	60	31 25	64 94	90		

The third Rumb
from the Meridian.

North-east by North,
South-east by South,

North-west by North,
South-west by South,

La	Long.	Diff.	La	Long.	Diff.	La	Long.	Diff.
Gr	Gr. P.	Gr. P.	Gr	Gr. P.	Gr. P.	Gr	Gr. P.	Gr. P.
0	0	0	30 21	03	36 08	60 50	42	72 16
1	0 66	1 20	31 21	80	37 28	61 51	78	73 36
2	1 33	2 40	32 22	58	38 49	62 53	18	74 56
3	2 00	3 61	33 23	38	39 69	63 54	63	75 77
4	2 67	4 81	34 24	18	40 89	64 56	12	76 97
5	3 34	6 01	35 25	00	42 09	65 57	68	78 17
6	4 01	7 22	36 25	82	43 30	66 59	29	79 37
7	4 68	8 42	37 26	64	44 50	67 60	96	80 58
8	5 36	9 62	38 27	48	45 70	68 62	71	81 78
9	6 03	10 82	39 28	34	46 90	69 64	53	82 98
10	6 71	12 03	40 29	21	48 11	70 66	44	84 19
11	7 39	13 23	41 30	09	49 31	71 68	45	85 39
12	8 07	14 43	42 30	98	50 51	72 70	55	86 59
13	8 76	15 64	43 31	88	51 71	73 72	77	87 79
14	9 44	16 84	44 32	80	52 92	74 75	12	89 00
15	10 13	18 04	45 33	74	54 12	75 77	62	90 20
16	10 83	19 24	46 34	69	55 32	76 80	30	91 40
17	11 53	20 45	47 35	67	56 52	77 83	15	92 61
18	12 23	21 65	48 36	66	57 73	78 86	25	93 81
19	12 93	22 85	49 37	67	58 93	79 89	60	95 01
20	13 64	24 05	50 38	69	60 13	80 93	27	96 22
21	14 35	25 26	51 39	74	61 33	81 97	32	97 42
22	15 07	26 46	52 40	82	62 54	82 101	85	98 62
23	15 80	27 66	53 41	91	63 74	83 106	97	99 82
24	16 53	28 86	54 43	03	64 94	84 112	90	101 03
25	17 26	30 07	55 44	19	66 15	85 119	90	102 23
26	18 00	31 27	56 45	37	67 35	86 128	45	103 43
27	18 75	32 47	57 46	58	68 55	87 139	47	104 64
28	19 50	33 67	58 47	82	69 75	88 155	00	105 84
29	20 26	34 88	59 49	11	70 96	89 181	58	107 04
30	21 03	36 08	60 50	42	72 16	90		

The fourth Number
from the Meridian.

North-east,
South-east,

North-west,
South-west.

La			Long.			Dist.			La			Long.			Dist.			La			Long.			Dist.		
Gr.	Gr.	P.	Gr.	Gr.	P.	Gr.	Gr.	P.	Gr.	Gr.	P.	Gr.	Gr.	P.	Gr.	Gr.	P.	Gr.	Gr.	P.	Gr.	Gr.	P.	Gr.	Gr.	P.
0		0			0		30	31	47	42	43		60	75	46		84	85								
1	1	00	1	41			31	32	63	43	84		61	77	49		86	27								
2	2	00	2	83			32	33	81	45	25		62	79	58		87	68								
3	3	00	4	24			33	34	99	46	67		63	81	75		89	09								
4	4	00	5	66			34	36	19	48	07		64	83	99		90	51								
5	5	01	7	07			35	37	41	49	50		65	86	31		91	91								
6	6	01	8	49			36	38	63	50	91		66	88	73		93	34								
7	7	02	9	90			37	39	88	52	33		67	91	23		94	75								
8	8	03	11	31			38	41	14	53	74		68	93	85		96	17								
9	9	04	12	73			39	42	42	55	15		69	96	58		97	58								
10	10	05	14	14			40	43	71	56	57		70	99	43		98	99								
11	11	07	15	56			41	45	03	57	98		71	102	43		100	41								
12	12	09	16	97			42	46	36	59	40		72	105	58		101	82								
13	13	11	18	38			43	47	72	60	81		73	108	91		103	24								
14	14	14	19	80			44	49	10	62	22		74	112	43		104	65								
15	15	17	21	21			45	50	50	63	64		75	116	17		106	06								
16	16	21	22	63			46	51	93	65	05		76	120	17		107	48								
17	17	25	24	04			47	53	38	66	46		77	124	45		108	89								
18	18	30	25	45			48	54	86	67	88		78	129	08		110	31								
19	19	36	26	87			49	56	37	69	29		79	134	10		111	72								
20	20	42	28	28			50	57	91	70	71		80	139	59		113	14								
21	21	49	29	70			51	59	48	72	12		81	145	65		114	55								
22	22	56	31	11			52	61	09	73	54		82	152	42		115	96								
23	23	64	32	53			53	62	73	74	95		83	160	10		117	38								
24	24	73	33	94			54	64	41	76	37		84	168	95		118	79								
25	25	28	35	35			55	66	13	77	78		85	179	41		120	21								
26	26	94	36	77			56	67	90	79	20		86	192	21		121	62								
27	28	06	38	18			57	69	71	80	61		87	208	71		123	04								
28	29	18	39	60			58	71	57	82	02		88	231	95		124	45								
29	30	32	41	01			59	73	49	83	44		89	271	71		125	86								
30	31	47	42	43			60	75	46	84	85		90													

Let <i>h</i> be Run be from the Meridian.				North-east and by East, South-east and by East.				North-west and by West, South-west and by West.			
La Long.		Diff.		La Long.		Diff.		La Long.		Diff.	
Gr.	Gr. P.	Gr.	P.	Gr.	Gr. P.	G.	P.	Gr.	Gr. P.	G.	P.
0	0	0	0	30	47	10	54	00	60	112	93
1	1 49	1	80	31	48	84	55	80	61	115	97
2	2 99	3	60	32	50	60	57	60	62	119	10
3	4 49	5	40	33	52	37	59	40	63	122	34
4	6 00	7	20	34	54	16	61	20	64	125	70
5	7 50	9	00	35	55	98	63	00	65	129	18
6	9 00	10	80	36	57	82	64	80	66	132	78
7	10 50	12	60	37	59	68	66	60	67	139	54
8	12 01	14	40	38	61	57	68	40	68	140	45
9	13 52	16	20	39	63	48	70	20	69	144	53
10	15 04	18	00	40	65	42	72	00	70	148	81
11	16 56	19	80	41	67	39	73	80	71	153	30
12	18 09	21	60	42	69	39	75	60	72	158	00
13	19 62	23	40	43	71	42	77	40	73	163	00
14	21 16	25	20	44	73	48	79	20	74	168	26
15	22 70	27	00	45	75	58	81	00	75	173	86
16	24 62	28	80	46	77	72	82	80	76	179	84
17	25 82	30	60	47	79	89	84	60	77	186	26
18	27 39	32	40	48	82	10	86	40	78	193	17
19	28 97	34	20	49	84	36	88	20	79	200	69
20	30 55	36	00	50	86	67	90	00	80	208	91
21	32 15	37	80	51	89	03	91	80	81	217	98
22	33 76	39	60	52	91	43	93	60	82	228	13
23	35 38	41	40	53	93	88	95	40	83	239	61
24	37 01	43	20	54	96	40	97	20	84	252	85
25	38 66	45	00	55	98	98	99	00	85	268	51
26	40 32	46	80	56	101	62	100	80	86	287	67
27	42 00	48	60	57	104	33	102	60	87	312	36
28	43 67	50	40	58	107	12	104	40	88	345	15
29	45 38	52	20	59	109	98	106	20	89	406	72
30	47 10	54	00	60	112	93	108	00	90		

The six Runbs from the Meridian.					East North-east, West North-west.					East South-east, West South-west.				
Lat.		Long.		Dist.	Lat.		Long.		Dist.	Lat.		Long.		Dist.
Gr.	G.	P.	Gr.	P.	Gr.	G.	P.	G.	P.	Gr.	G.	P.	G.	P.
0		0		0	30	75	98	78	39	60	182	18	156	78
1	2	41	2	61	31	78	78		81 00	61	187	07	155	40
2	4	83	5	23	32	81	61		83 62	62	192	13	162	01
3	7	25	7	84	33	84	48		86 23	63	197	36	164	62
4	9	66	10	45	34	87	37		88 84	64	202	77	167	24
5	12	06	13	06	35	90	30		91 46	65	208	38	169	85
6	14	51	15	68	36	93	27		94 07	66	214	20	172	46
8	16	94	18	29	37	96	27		96 68	67	220	25	175	08
7	19	37	20	90	38	99	31		99 30	68	226	57	177	69
9	21	81	23	52	39	102	40		101 91	69	233	15	180	30
10	24	26	26	13	40	105	53		104 52	70	240	06	182	92
11	26	71	28	74	41	108	71		107 14	71	247	27	185	53
12	29	17	31	36	42	111	93		109 75	72	254	90	188	14
13	31	65	33	97	43	115	20		112 36	73	262	92	190	75
14	34	14	36	58	44	118	53		114 97	74	271	43	193	37
15	36	63	39	20	45	121	92		117 59	75	280	46	195	98
16	39	13	41	81	46	125	36		120 20	76	290	11	198	59
17	41	65	44	42	47	128	87		122 81	77	300	46	201	21
18	44	18	47	03	48	132	44		125 43	78	311	62	203	82
19	46	73	49	65	49	136	09		128 04	79	323	73	206	43
20	49	29	52	26	50	139	81		130 65	80	337	00	209	05
21	51	87	54	87	51	143	60		133 27	81	351	64	211	66
22	54	47	57	49	52	147	47		135 88	82	368	00	214	27
23	57	08	60	10	53	151	44		138 49	83	386	51	216	89
24	59	71	62	71	54	155	50		141 10	84	407	89	219	50
25	62	36	65	33	55	159	66		143 72	85	433	13	222	11
26	65	04	67	94	56	163	93		146 33	86	464	05	224	73
27	67	74	70	55	57	168	31		148 95	87	503	88	227	34
28	70	46	73	17	58	172	80		151 56	88	560	00	229	95
29	73	20	75	78	59	177	42		154 17	89	636	08	232	56
30	75	98	78	39	60	182	18		156 78	90				

The seventh Runbe from the Meridian.			East and by North, West and by North,			East and by South, West and by South,		
La	Long.	Dist.	La	Long.	Dist.	La	Long.	Dist.
Gr.	Gr. P.	Gr. P.	Gr.	Gr. P.	Gr. P.	Gr.	Gr. P.	Gr. P.
0	0	0	30	158 23	153 77	60	379 35	307 55
1	5 02	5 12	31	164 06	158 90	61	389 56	312 67
2	10 05	10 25	32	169 96	164 02	62	400 10	317 80
3	15 08	15 38	33	175 92	169 15	63	410 98	322 93
4	20 12	20 50	34	181 95	174 28	64	422 26	328 05
5	25 16	25 63	35	188 04	179 40	65	433 94	333 18
6	30 21	30 75	36	194 22	184 53	66	446 03	338 30
7	35 27	35 88	37	200 48	189 65	67	458 66	343 43
8	40 34	41 00	38	206 82	194 78	68	471 80	348 55
9	45 42	46 13	39	213 24	199 90	69	485 52	353 68
10	50 52	51 26	40	219 76	205 03	70	499 89	358 81
11	55 63	56 38	41	226 37	210 16	71	514 94	363 93
12	60 77	61 51	42	233 08	215 28	72	530 79	369 06
13	65 92	66 63	43	239 90	220 41	73	547 52	374 18
14	71 09	71 76	44	246 84	225 53	74	565 22	379 31
15	76 28	76 88	45	253 89	230 66	75	584 03	384 43
16	81 50	82 01	46	261 05	235 79	76	604 13	389 56
17	86 75	87 14	47	268 36	240 91	77	625 67	394 69
18	92 02	92 26	48	275 80	246 04	78	648 91	399 81
19	97 31	97 39	49	283 40	251 16	79	674 15	404 94
20	102 64	102 51	50	291 13	256 29	80	701 75	410 06
21	108 01	107 64	51	299 03	261 41	81	732 25	415 19
22	113 42	112 77	52	307 11	266 54	82	766 30	420 32
23	118 87	117 89	53	315 37	271 69	83	804 86	425 44
24	124 35	123 02	54	323 82	276 79	84	849 38	430 57
25	129 87	128 14	55	332 48	281 92	85	901 98	435 69
26	135 44	133 27	56	341 36	287 04	86	966 31	440 82
27	141 05	138 40	57	350 47	292 17	87	1049 26	445 94
28	146 71	143 52	58	359 81	297 30	88	1166 11	451 07
29	152 44	148 65	59	369 45	302 42	89	1366 23	456 20
30	158 23	153 77	60	379 35	307 51	90		

The eight Rumbes of East and West, with the Longitude answering to one degree of distance, and the distance belonging to one degree of Longitude.

La	Long.	Dist.	La	Long.	Dist.	La	Long.	Dist.
Gr	Gr. P.	Parts.	Gr	Gr. P.	Parts.	Gr	Gr. P.	Parts.
0		0 100 00	30	1 25	86 60	60	2 00	50 00
1	1 00	99 98	31	1 17	85 71	61	2 06	48 48
2	1 00	99 94	32	1 18	84 80	62	2 13	46 94
3	1 00	99 86	33	1 19	83 86	63	2 20	45 40
4	1 00	99 75	34	1 21	82 90	64	2 28	43 83
5	1 00	99 62	35	1 22	81 91	65	2 37	42 26
6	1 01	99 45	36	1 24	80 90	66	2 46	40 67
7	1 01	99 25	37	1 25	79 86	67	2 56	39 07
8	1 01	99 02	38	1 27	78 80	68	2 67	37 46
9	1 01	98 76	39	1 29	77 71	69	2 79	35 83
10	1 02	98 48	40	1 31	76 60	70	2 92	34 20
11	1 02	98 16	41	1 33	75 47	71	3 07	32 55
12	1 02	97 81	42	1 35	74 31	72	3 24	30 90
13	1 03	97 43	43	1 37	73 13	73	3 42	29 23
14	1 03	97 03	44	1 39	71 93	74	3 63	27 56
15	1 03	96 59	45	1 41	70 71	75	3 86	25 88
16	1 04	96 12	46	1 44	69 46	76	4 13	24 19
17	1 04	95 63	47	1 47	68 20	77	4 44	22 49
18	1 05	95 10	48	1 49	66 91	78	4 81	20 79
19	1 06	94 55	49	1 52	65 60	79	5 24	19 08
20	1 06	93 97	50	1 55	64 28	80	5 76	17 36
21	1 07	93 35	51	1 59	63 93	81	6 39	15 64
22	1 08	92 72	52	1 62	61 56	82	7 18	13 91
23	1 09	92 05	53	1 66	60 18	83	8 20	12 18
24	1 09	91 35	54	1 70	58 77	84	9 57	10 45
25	1 10	90 63	55	1 74	57 35	85	11 47	8 71
26	1 11	89 88	56	1 79	55 92	86	14 33	61 97
27	1 12	89 10	57	1 84	54 46	87	19 11	5 23
28	1 13	88 29	58	1 89	52 99	88	28 65	3 49
29	1 14	87 46	59	1 94	51 50	89	57 36	1 74
30	1 15	86 60	60	2 00	50 00	90		0

These tables are calculated for each of the Rumbes. The first seven have three columnes, and of them the first containeth the degrees of Latitude, from the Equinoctiall to the Pole: the second doth giue the difference of Longitude; and the third the distance, both of them belonging to that Rumb and latitude.

As in the Table of the third Rumb; at the latitude of 50 Gr. I find vnder the title of *Longitude* 38 Gr. 69 parts, and vnder the title of *Distance* 60 Gr. 13 parts. This shewes that if the course held constantly on the third Rumb from the Equinoctiall to the Latitude of 50 Gr. the difference of Longitude would be 38 Gr. 69 parts of a 100, and the distance vpon the Rumb 60 Gr. 13 parts. For here I reckon the distance by degrees, rather then by leagues or miles, and subdiuide each degree into 100 parts, rather then into 60 minures, for the more ease in calculation, and withall to make the calculation to agree the better, both with this, and my *Crosse Staffe*, and other instruments.

The use of these Tables, for the finding of the difference of Longitude, is this. Turne to the table of the Rumb, and there see what longitude belongeth to either latitude, then take the one longitude out of the other, the remainder will be the difference of longitude required.

As in the former example, where the places given were A, in the latitude of 50 Gr. C in the latitude of 55 Gr. and the Rumb the third from the meridian: I looke into the table of the third Rumb and there find,

Latitude 50 gr.	Longitude 38 gr. 69 parts.
Latitude 55.	Longitude 44. 19.
Therefore the diff. of Longitude 5 50	

There is another use of these tables, for the describing of the Rumbes both on the *Globe*, and all sorts of *Charts*. For hauing drawne the circles of Longitude and latitude, and finding by the tables, the difference of longitude belonging to each Rumb and latitude: If we make a prick in the chart, at
euery

every degree of latitude, according to that difference of longitude, and draw lines through those prickles, so as they make no angles, the lines so drawne shall be the Rumbs required.

The use of the eight Rumb is something different from the rest. For there being here no change of latitude, I have set to each latitude, the difference of longitude, belonging to one degree of distance, and the distance belonging to one degree of longitude.

As if two places shall be 20 leagues, or one degree distant one from the other, in the latitude of 50 gr. the difference of longitude betweene them will be 1 gr. 55 parts. But if they differ one degree in longitude, the distance betweene them will be onely 64 parts, which fall short of 13 leagues, or at the most 6428 parts, such as 10000 do make a degree.

6 By the difference of longitude, Rumb, and one latitude, to find the other latitude.

As if the places given were *A*, in the latitude of 50 gr. *C* in a greater latitude but unknowne, the difference of longitude 5 gr. $\frac{1}{2}$, and the Rumb the third from the Meridian.

In the chart let *AB*, *DC*, meridians, be drawne through *A* and *C*, according to the difference of longitude, one 5 gr. $\frac{1}{2}$ from the other; and a parallell of latitude through *A*, crossing the meridian *CD* in *D*: then in *A*, with *AB*, make an angle of the Rumb *BAC*: so the degrees in the meridian betweene *D* and *C*, shall be found to be 5 gr. the proper difference of latitude which was required. Wherefore the proportion holds for the *Sector*,

As *AD* the Radius,

to *DC* the tangent of the Rumb from the equator:

So *AD* as difference of longitude,

to *DC* the proper difference of latitude.

According to this, I take 56 gr. 15 m. for the angle of the Rumb from the equator, out of the greater Tangent, and make

make it a parallell Radius. Then I reckon $5\text{ gr.}\frac{1}{2}$ in the line of *lines* from the center, for the difference of longitude. So the parallell taken from the termes of this difference, and measured in the line of *meridians*, shal reach from 50 gr. the latitude giuen, to 55 gr. which is the latitude required.

Or if the Rumb fall nearer to the meridian.

As BC the tangent of the Rumb from the meridian,
is to AB the Radius:

So BC as difference of longitude,
to AD the proper difference of latitude.

According to this we may best work by parallel entrance; first take $33\text{ gr.}45\text{ m.}$ for the angle of the Rumb from the meridian, out of the greater *Tangent*, and make it a parallell Radius; then take $5\text{ gr.}\frac{1}{2}$ for the difference of longitude out of the line of *lines*, and carrie it parallell to the former, till the feete of the compasses stay in like points: so the line between the center and the place of this stay, being taken and measured in the line of *meridians* from 50 gr. forward, shall shew the latitude required to be 55 gr. as in the former way.

The like may be found by the tables of Rumbs. For in the table of the third Rumb, at the latitude of 50 gr. I finde the longitude of $38\text{ gr.}69\text{ p.}$ to this if I adde $5\text{ gr.}50\text{ p.}$ for the difference of longitude giuen, the compound longitude will be $44\text{ gr.}19\text{ p.}$ and this answers to the latitude of 55 gr.

But if this difference of latitude were to be found by the common sea-chart, it should seeme to be $8\text{ gr.}13\text{ m.}$; and so the second latitude should be $58\text{ gr.}13\text{ m.}$ which is about 3 gr. more then the truth.

7 By one latitude, rumb, and distance, to find the difference of longitude.

As if the places giuen were A in the latitude of 50 gr. C in a greater latitude but vnknowne, the distance vpon the Rumb being 6 gr. betweene them, and the Rumb the third from the meridian.

In the chart, let a meridian AB , and a parallell AD be drawne through A ; and in A , with AB , make an angle BAC for the Rumb from the meridian; then open the compasses according to the latitude of the places to EF , the quantitie of 6 gr. in the meridian, transferring them into the Rumb from A to C , and through C draw another meridian DC , crossing the parallell drawne through A in D : so the degrees intercepted in the parallell from A to D , shall shew the difference of longitude required to be about $5\text{ gr.}\frac{1}{2}$. Wherefore the proportion holds for the Sector.

As AC the Radius, (meridian:
is to AD , as quall to BC , the sine of the Rumb from the
So AC as proper distance vpon the Rumb,
to AD the difference of longitude.

According to this I take the sine of $33\text{ gr.}45\text{ m.}$ for the angle of the Rumb from the meridian, and make it a parallell Radius; then keeping the Sector at this angle, I take 6 gr. for the distance out of the meridian line, according to the estimated latitudes of both places, and lay it on both sides of the Sector from the center: so the parallell taken from the termes of this distance, and measured in the lines of *lines*, shall shew the difference of longitude to be about $5\text{ gr.}\frac{1}{2}$.

In this, and some of the *Prop.* following, where there is but one latitude knowne, there may be sometimes an error of a minute or two, in the estimation of the proper distance, yet it may be rectified at a second operation.

This proposition may also be wrought by the Tables of Rumbs. For according to the example, in the Table of the third Rumb, at the latitude of 50 gr. I find the longitude of $38\text{ gr.}69\text{ p.}$ and the distance of $60\text{ gr.}13\text{ p.}$ to this I add 6 gr. for the distance giuen; so the compound distance will be $66\text{ gr.}13\text{ p.}$ and this answers to the longitude of $44\text{ gr.}19\text{ p.}$; then if I take the one longitude out of the other, the difference will be $5\text{ gr.}50\text{ p.}$ as before.

But if this difference were to be found by the common sea-chart, it should seeme to be onely $3\text{ gr.}20\text{ m.}$ which is

R

more

more then 2 gr. lesse then the truth.

8 By one latitude, Rumb, and difference of longitudes,
to find the distance.

As if the places given were *A*, in the latitude of 50 gr. *C* in a greater latitude but unknowne, the difference of longitude betweene them being 5 gr. $\frac{1}{2}$, and the Rumb the third from the meridian.

In the chart let *A B*, *D C*, meridians be drawne through *A* and *C*, according to the difference of longitude, and a parallell of latitude through *A*, crossing the meridian *DC* in *D*; then in *A*, with *A B*, make an angle of the Rumb *BAC*: so the distance on the Rumb from *A* to *C* taken and measured in the meridian, according to the estimated latitude of the places, shall be found to be 6 gr. Wherefore the proportion holds for the Sellar.

As *A D*, equall to *BC*, the sine of the Rumb from the meridian is to *AC* the Radius:

(dian,

So *A D* as difference of longitudes,

to *AC* the proper distance vpon the Rumb.

According to this, I take the laterall Radius, and make it a parallell sine of 33 gr. 45 m. which is here the angle of the Rumb from the meridian; then I reckon 5 gr. $\frac{1}{2}$ in the lines of lines from the center, for the difference of longitude: so the parallell taken from the termes of this difference, and measured in the line of meridians, according to the latitudes of the places, shall there shew the distance required to be about 6 gr. which are 120 leagues.

Or if the Rumb fall nearer to the meridian, that the laterall Radius cannot be fitted ouer in his sine, this *Prop.* must be wrought by parallell entrance, and so also it giues the same distance as before.

Or we may find this distance by the Table of Rumbs. For in the table of the third Rumb, at the latitude of 50 gr. I find the longitude of 38 gr. 69 p. and the distance of 60 gr. 13 p.

To

To this longitude here found, I adde 5 gr. 50 p. for the difference of longitude given: so the compound longitude will be 44 gr. 19 p. and this answers to the distance of 66 gr. 15 p. Then if I take the one distance out of the other, the remainder will be 6 gr. 02 p. for the distance required.

But if this distance were to be measured on the common sea-chart, it should seeme to be almost 10 gr. or at the least 197 leagues, about 77 leagues more then the truth.

9 By one latitude, distance, and difference of longitudes, to find the Rumb.

As if the places given were *A*, in the latitude of 50 gr. *C* in a greater latitude but vnkowne, the difference of longitude betweene them being 5 gr. $\frac{1}{2}$, and the distance 6 gr. vpon the Rumb.

In the chart let *AB*, *DC*, meridians, be drawne through *A* and *C*, and a parallell of latitude through *A*; then open the compasses according to the latitudes of the places, to *EF* the quantitie of 6 gr. in the meridian, and setting the one foote in *A*, the other foote shall crosse the other meridian in *C*; and if we draw the right line *AC*, the angle *BAC* shall shew the inclination of the Rumb to the meridian to be about 33 gr. 45 m. Wherefore the proportion holds for the Sector.

As *AC* the proper distance vpon the Rumb,
is to *AD* the difference of longitude:

So *AC* as Radius,

to *AD*, equall to *BC*, the sine of the Rumb from the meridian.

According to this, I take the proper distance 6 gr. out of the line of meridians, and lay it on both sides of the Sector from the center; then I take the difference of longitude 5 gr. $\frac{1}{2}$ out of the line of lines, and to it open the Sector in the terms of the former distance: so the parallell Radius taken from betweene 90 and 90, and measured in the sines, doth giue about 33 gr. 45 m. for the Rumb required.

But if this Rumb were to be found by the common sea-

chart, it should seeme to be about 66 gr. and so almost the sixt Rumb from the meridian.

10 *By the longitude and latitude of two places,
to find their distance vpon the Rumb.*

Let the *Sector* be opened in the lines of *lines*, vnto a right angle (as was shewed before *Cap. 2. Prop. 7.*) then take out the proper difference of latitude, and lay it on the one line, and the difference of longitude, and lay it on the other line, so as they may both meete in the center, marking how far they extend. For the line taken from the termes of their extension, and measured in the *meridian*, according to their latitudes, shall shew the distance required.

So if the places giuen were *A* and *C*, *A* in the latitude of 50 gr. *C* in the latitude of 55 gr. the proper difference of latitude shall be the line *AB*, and let *BC* the difference of longitude be 5 gr. $\frac{1}{2}$, we shall find that *AC* the distance vpon the Rumb is about 6 gr. which make 120 leagues.

For in the chart, let an occult meridian be drawne through *A*, and a parallell of latitude through *C*, crossing the former meridian in *B*, and a right line for the Rumb from *A* to *C*, so haue we a rectangle triangle *ABC*, whose base *AC*, taken and measured in the meridian from *E* below 50 gr. to *F*, as much about 55 gr. doth containe the quantitie of 6 gr.

In the same maner the *Sector* being opened to a right angle, in the lines of *lines*: if we take the difference of latitude out of the line of *meridians*, in his proper place from 50 gr. to 55 gr. and place it on one of the sides from the center, to resemble *AB*, then reckon the difference of longitude on the other perpendicular line from the center to 5 gr. $\frac{1}{2}$, in stead of *BC*, we shall haue the like rectangle triangle on the *Sector*, to that which we had before on the chart; and if we take out the base of it, and measure it in the line of *meridians* from below 50 gr. to as much about 55 gr. we shall finde as before, that it containeth about 6 gr. or 120 leagues.

But if this distance were to be measured on the common
sea-

sea-chart, it should seeme to be almost $7\text{ gr.}\frac{1}{2}$, or 145 leagues; which is 25 leagues more then the truth.

II By the latitude of two places, and the distance vpon the Rumb, to find the difference of longitude.

Let the *Settor* be opened in the lines of *lines* to a right angle, then take out the proper difference of latitudes, and lay it on one of the lines from the center, then take the proper distance with a paire of compasses, and setting one foote in the termes of the difference, turne the other foote to the other line of the *Settor*, and it shall there shew the difference of longitude required.

So if the places giuen were *A*, in the latitude of 50 gr. *C* in the latitude of 55 gr. with 6 gr. of distance one from another, we shall find their difference of longitude to be about $5\text{ gr.}\frac{1}{2}$.

For in the chart let a meridian *AB* be drawne for the one, and *BC*, *AD*, parallels of latitude for them both. Then open the compasses according to the latitude of the places, to *EF* the quantitie of 6 gr. in the *meridian*, and setting one foote in *A*, hauing latitude of 50 gr. turne the other to the parallell of 55 gr. and it shall there cut off the required difference of longitude *BC* $5\text{ gr.}\frac{1}{2}$.

In the same maner, the *Settor* being opened to a right angle, in the lines of *lines*: if we take the difference of latitude out of the line of *meridians* in his proper place from 50 gr. vnto 55 gr. and place it on one of the lines from the center; then take 6 gr. the distance vpon the Rumb out of the same line of *meridians*, according to the latitudes of the places, and set the one foote in the terme of the former difference, turning the other foote to the other perpendicular line, we shall finde that it will crosse it about $5\text{ gr.}\frac{1}{2}$ from the center: which is the difference of longitude required.

But if this difference of longitude were to be found by the common sea-chart, it would seeme to be only $3\text{ gr. } 20\text{ m.}$ which is more then $2\text{ gr. } 10\text{ m.}$ lesse then the truth.

12 By one latitude, distance and difference of longitudes,
to finde the difference of latitudes.

Let the *Sector* be opened in the line of *lines* to a right angle, and let the difference of longitude be reckoned in one of those lines from the center; then take the proper distance with a paire of compasses, and setting the one foote in the terme of the former difference, turne the other foote to the other line of the *Sector*, and it shal thence cut off a line, equal to the proper difference of latitude required.

So if the places given were *A* and *C*, *A* in the latitude of 50 gr. *C* in a greater latitude but vnknowne, the difference of longitude betweene them 5 gr. $\frac{1}{2}$, and the distance vpon the Rumb 6 gr. or 120 leagues, we shall find the difference of latitude to be 5 gr.

For in the chart, let occult meridians be drawne through *A* and *C*, and a parallell of latitude through *A*; then open the compasses according to the estimated latitudes of the places to *E F* the quantitie of 6 gr. in the meridian, and setting the one foote in *A*, turne the other to the meridian drawne through *C*, and it shall there cut off the line *D C*, which is the difference of latitude required.

In the same maner, the *Sector* being opened to a right angle, in the lines of *lines*, if in the one line we reckon the difference of longitude from the center to 5 gr. $\frac{1}{2}$, then taking 6 gr. for the distance out of the line of *Meridians*, according to the latitude of the places, we set the one foote in the terme of the given difference, and turne the other foote to the other perpendicular line, we shall finde that it cuts a line from it, which taken and measured in the line of *meridians*, from 50 gr. on forward, doth shew the difference of latitude to be as before 5 gr.

But if this difference of latitude were to be found by the common sea-chart, it would seeme to be only 2 gr. 25 m. which is 2 gr. 35 m. lesse then the truth. Such is the difference betweene both these charts.

THE

THE THIRD BOOKE

Containing the vse of the particular
Lines.

THE lines of *lines*, of *superficies*, of *solids*, of *lines*, with the laterall lines of *tangents* and *meridians*, whereof I haue hitherunto spoken, are those which I principally intended: that little roome on the *Sector* which remaineth, may be filled vp with such particular lines as each one shall think conuenient for his purpose. I haue made choise of such as I thought might be best prickt on without hindring the sight of the former, viz. lines of *Quadrature*, of *Segments*, of *Inscribed bodies*, of *Equated bodies*, and of *Mettals*.

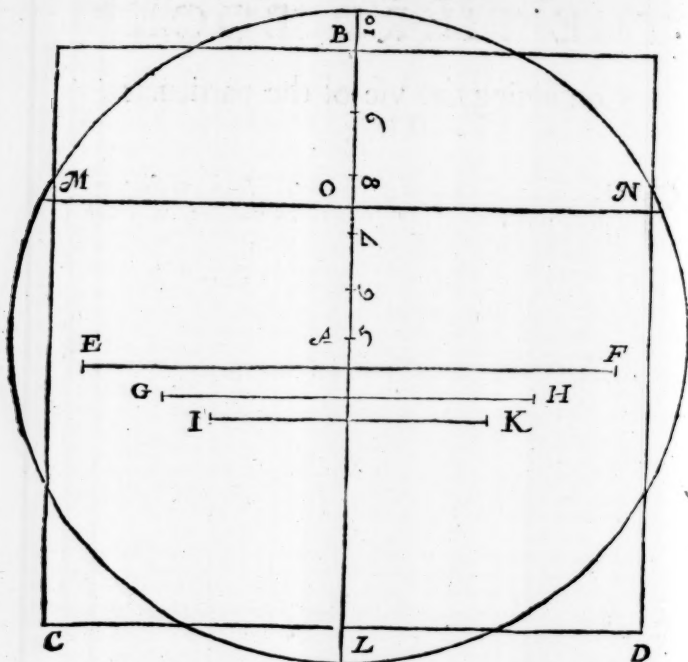
CHAP. I.

Of the lines of Quadrature.

THe lines of *quadrature* may be knowne by the letter *Q*, and by their place betweene the lines of *lines*. *Q* signifieth the side of a *square*; *5* the side of a *pentagon* with five equall sides, *6* of an *hexagon* with six equall sides, and so *7*, *8*, *9*, and *10*. *S* stands for the Semidiameter of a circle, and *90* for a line equall to *90 gr.* in the circumference. The vse of them may be

- 1 *To make a square equall to a circle giuen.*
- 2 *To make a circle equall to a square giuen.*

If the circle be first giuen, take his semidiameter, and to it open the *Sector* in the points at *S*: so the parallell taken from betweene the points at *Q*, shall be the side of the square required.



If the square be given take his side, and to it open the Sector, in the points at Q : so the parallell taken from between the points at S , shall be the Semidiameter of the circle required.

Let the Semidiameter of the circle given be AB , the side of the square equall vnto it shall be found to be CD .

3 To reduce a circle given, or a square into an equall pentagon, or other like sided and like angled figure.

Take the side of the figure given, and fit it over in his due points: so the parallels taken from between the points of the

the other figures, shall be the sides of those figures: which being made vp with equall angles, shall be all equall one to the other.

Let the Semidiameter of the circle giuen be AB , the side of an hexagon equal to this circle, shall by these meanes be found to be GH ; and the sides of an octagon to be IK . Other planes not here set downe, may first be reduced into a square, by the sixt *Prop. Superf.* and then into a circle, or other of these equall figures, as before.

4 To find a right line, equall to the circumference of a circle, or other part thereof.

Take the Semidiameter of the circle giuen, and to it open the Sector in the points at S ; so the parallell taken from betwene the points at 90 in this line, shall be the fourth part of the circumference: which being knowne, the other parts may be found out by the second and third *Prop. of lines*.

Thus if the Semidiameter of the circle giuen be AB , the right line EF shall be found to be the fourth part of the circumference. Therefore the double of EF shall be equall to the circumference of 180 gr; and the halfe of EF shall be the circumference of 45 gr. and so in the rest.

CHAP. II.

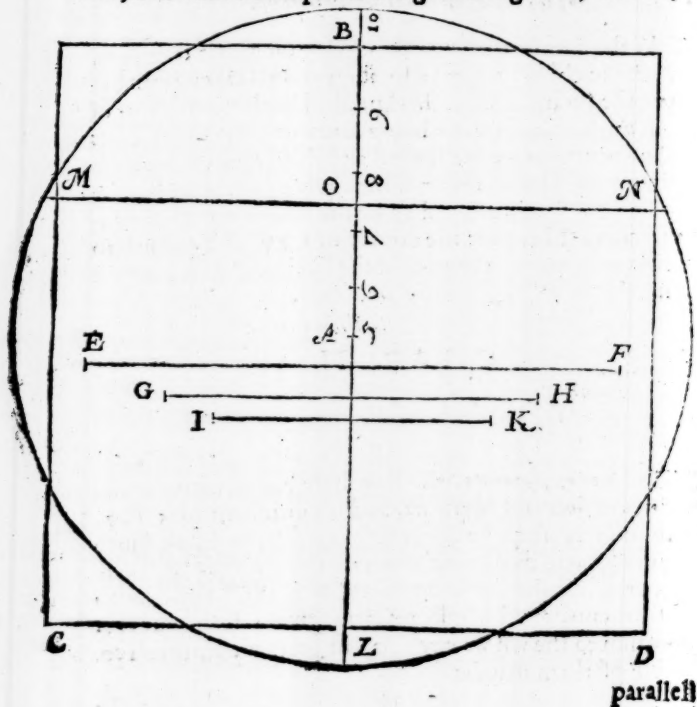
Of the lines of Segments.

THE lines of segments which are here placed between the lines of *lines* and *superficies*, and are numbered by 5, 6, 7, 8, 9, 10, do represent the diameter of a circle, so diuided into a hundred parts, as that a right line drawne through these parts, perpendicular to the diameter, shall cut the circle into two segments, of which the greater segment shall haue that proportion to the whole circle, as the parts cut haue to 100. The vse of them may be

- 1 To divide a circle given into two segments,
according to a proportion given.
- 2 To finde a proportion betweene a circle
and his segments given.

Let the *Sector* be opened in the points of an 100, to the diameter of the circle given: so a parallell taken from the points proportionall to the greater segment required, shall give the depth of that greater segment.

Or if the segments be given, let the *Sector* be opened as before; then take the depth of the greater segment, and carry it



parallell to the diameter: so the number of points wherein they stay, shall shew the proportion to 100.

As if the diameter of the circle giuen were BL , the depth of the greater segment LO being 75, doth shew the proportion of the segment $OMLN$ to the circle to be as 75 to 100. viz. three parts of foure.

Hence I might shew, if there were any vse of it,

To find the side of a square, equall to any knowne segment of a circle.

The side of a square equall to the whole circle, may be found by the former *Cap.* and then hauing the proportion of the segment to the circle, we may diminish the square in such proportion, by that which hath been shewed *Lib. 1. Cap. 3. Prop. 3.*

CHAP. III.

Of the lines of Inscribed bodies.

THe lines of *inscribed bodies* are here placed betweene the lines of *lines*, and may be knowne by the letters, D, S, I, C, O, T ; of which D signifieth the side of a *dodecabedron*, I of an *Icosahedron, C of a *cube*, O of an *octahedron*, and T of a *tetrahedron*, all inscribed into the same sphere; whose semidiameter is here signified by the letter S .*

The vse of these lines may be,

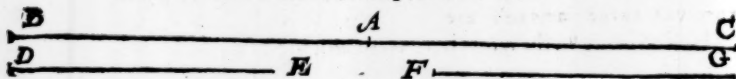
- 1 *The semidiameter of a sphere being giuen, to find the sides of the five regular bodies, which may be inscribed in the said sphere.*
- 2 *The side of any of the five regular bodies being giuen, to find the semidiameter of a sphere, that will circumscribe the said bodie.*

If the sphere be first giuen, take his semidiameter, and to it

S 2

open.

open the *Sector* in the points at *S*: if any of the other bodies be first giuen, take the side of it, and fit it ouer in his due points: so the parallell taken from betweene the points of the other bodies, shall be the sides of those bodies, and may be inscribed into the same sphere.



So if the semidiameter of the sphere be *AC*, the side of the *dodecahedron* inscribed shall be *DE*.

CHAP. IIII.

Of the lines of Equated bodies.

THe lines of *equated bodies* are here placed betweene the lines of *lines* and *solids*, noted with these letters, *D, I, C, S, O, T*, of which *D* stands for the side of a *dodecahedron*, *I* for the side of an *Icosahedron*, *C* for the side of a *cube*, *S* for the diameter of a *sphere*, *O* for the side of an *octahedron*, and *T* for the side of a *tetrahedron*, all equall one to the other. The vse of these lines may be

- 1 The diameter of a sphere being giuen, to find the sides of the five regular bodies equall to that sphere.
- 2 The side of any of the five regular bodies being giuen, to find the diameter of a sphere, and the sides of the other bodies, equall to the first body giuen.

If the sphere be first giuen, take his diameter, and to it open the *Sector* in the points at *S*: if any of the other bodies be first giuen, take the side of it, and fit it ouer in his due points: so the parallels taken from between the points of the other bodies, shall be the sides of those bodies equall to the first body giuen.

Thus in the last diagram, if the diameter of a sphere giuen be *BC*, the side of the *dodecahedron* equall to this sphere, would be found to be *FG*.

CHAP. V.

Of the Lines of Mettalls.

THe lines of *Mettalls* are here ioyned with those before of *equared bodies*, and are nored with these characters \odot , $\&$, h , D , P , S , U . of which \odot stands for gold, $\&$ for quicksiluer, h for leade, D for siluer, P for copper, S for iron, and U for tin. The vse of them is to giue a proportion betweene these seuerall mettalls, in their magnitudo and weight, according to the experiments of *Marinus Obetaldus*, in his booke called *Promotus Archimedes*.

- 1 *In like bodies of seuerall mettalls and equall weight, hauing the magnitudo of the one, to finde the magnitudo of the rest.*

Take the magnitudo giuen out of the lines of *Solids*, and to it open the *Settor* in the points belonging to the mettall giuen: so the parallells taken from between the points of the other mettalls, and measured in the lines of *Solids*, shall giue the magnitudo of their bodies.

Thus hauing cubes or spheres of equall weight, but seuerall mettalls, we shall finde that if those of tin containe 10000 *D*, the others of iron wil containe 9250, those of copper 8222, those of siluer 7161, those of leade 6435, those full of quicksiluer 5453, and those of gold 3895.

- 2 *In like bodies of seuerall mettalls and equall magnitudo, hauing the weight of one to finde the weight of the rest.*

This proposition is the conuerse of the former, the proportion not direct, but recipocall, wherefore hauing two like bodies, take the giuen weight of the one out of the lines of *Solids*, and to it open the *Settor* in the points belonging to

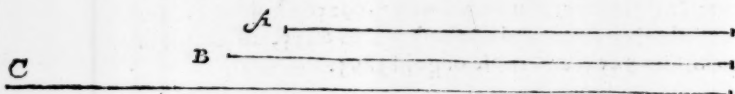
the mettall of the other body: so the parallell taken from the points belonging to the body giuen, and measured in the lines of *Solids*, shall giue the weight of the body required.

As if a cube of gold weighed 38 lb . and it were required to know the weight of a cube of lead hauing equal magnitude. First I take 38 lb . for the weight of the golden cube, out of the lines of *Solids*, & put it ouer in the points of h belonging to lead: so the parallell taken from between the points of \odot standing for gold, and measured in the lines of *Solids*, doth giue the weight of the leaden cube required to be 23 lb .

Thus if a sphere of gold shall weigh 10000, we shall finde that a sphere of the same diameter full of quicksiluer shall weigh 7143, a sphere of lead 6053, a sphere of siluer 5438, a sphere of copper 4737, a sphere of iron 4210, and a sphere of tin 3895.

3 *A bodie being giuen of one mettall, to make another like vnto it, of another mettall, and equall weight.*

Take out one of the sides of the bodie giuen, and put it ouer in the points belonging to his mettall: so the parallell taken from between the points belonging to the other mettall, shall giue the like side, for the bodie required. If it be an irregular bodie, let the other like sides be found out in the same manner.



Let the bodie giuen be a sphere of lead containing in magnitude 16 lb , whose diameter is A , to which I am to make a sphere of iron, of equall waight: If I take out the diameter A , and put it ouer in the points of h belonging to lead, the parallell taken from between the points of g standing for iron, shall be B , the diameter of the iron sphere required. And this compared with the other diameter, in the lines of *Solids*,

Solids will be found to be 23 d. in magnitude.

4 *A body being given of one mettall, to make another like vnto it of another mettall, according to a weight giuen.*

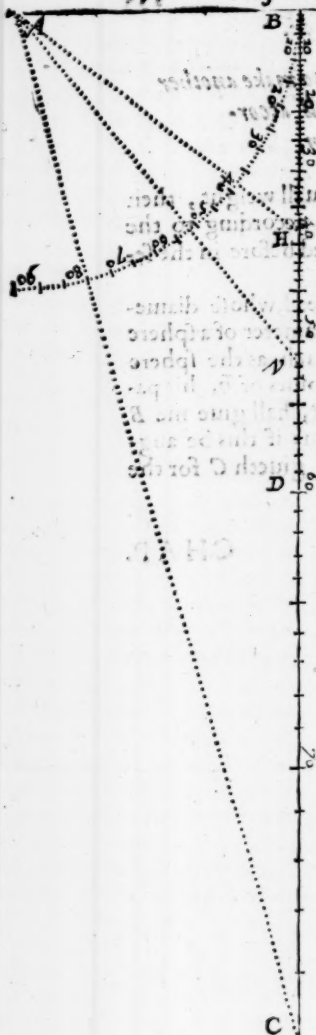
First find the sides of a like bodie of equall weight, then may we either augment or diminish them according to the proportion giuen by that which we shewed before in the second and third *Prop. of Solids*.

As if the bodie giuen were a sphere of lead, whose diameter is *A*, and it were required to find the diameter of a sphere of iron, which shall weigh three times as much as the sphere of lead: I take *A*, and put it ouer in the points of *h*, his parallell taken from betweene the points of *g*, shall giue me *B* for the diameter of an equall sphere of iron: if this be augmented in such proportion as 1 vnto 3, it giueth *C* for the diameter required.

CHAP.

CHAP. VI.

Of the lines on the edges of the Sector.



HAving shewed some vse of the lines on the flat sides of the *Sector*, there remains onely those on the edges. And here one halfe of the outward edge is diuided into inches, and numbred according to their distance from the ends of the *Sector*. As in the *Sector* of fourteene inches long, where we find 1 and 13, it sheweth that diuision to be 1 inch from the nearer end, and 13 inches from the farther end of the *Sector*.

The other halfe containeth a line of lesser *tangents*, to which the gnomon is Radius. They are here continued to 75 gr. And if there be need to produce them farther, take 45 out of the number of degrees required, and double the remainder: so the *tangent* and *secant* of this double remainder being added, shall make vp the *tangent* of the degrees required.

As if *AB* being the Radius, and *BC* the tangent line, it were required to find the tangent of 75 gr. If we take 45 gr. out of 75 gr. the remainder is 30 gr. and the double 60 gr. whose tangent is *BD*, and the secant is *AD*: if then we adde *AD* to *BD*, it maketh *BC* the tangent of 75 gr. which was required. In like sort the secant of 61 gr. added to the tangent of 61 gr. giueth the tangent of 75 gr. 30 m. and the secant of 62 gr. added to the tangent of 62 gr. giueth the tangent of 76 gr. and

and so in the rest. The vse of this line may be

To obserue the altitude of the Sunne.

Hold the *Sector* so as the tangent *BC* may be verticall, and the gnomon *B A* parallell to the horizon; then turne the gnomon toward the Sunne, so that it may cast a shadow vpon the tangent, and the end of the shadow shal shew the altitude of the Sunne. So if the end of the gnomon at *A*, do giue a shadow vnto *H*, it sheweth that the altitude is 38 gr. $\frac{1}{2}$; if vnto *D*, then 60 gr. and so in the rest.

There is another vse of this *tangent* line, for the drawing of the houre lines vpon any ordinary plane, whereof I will set downe these propositions.

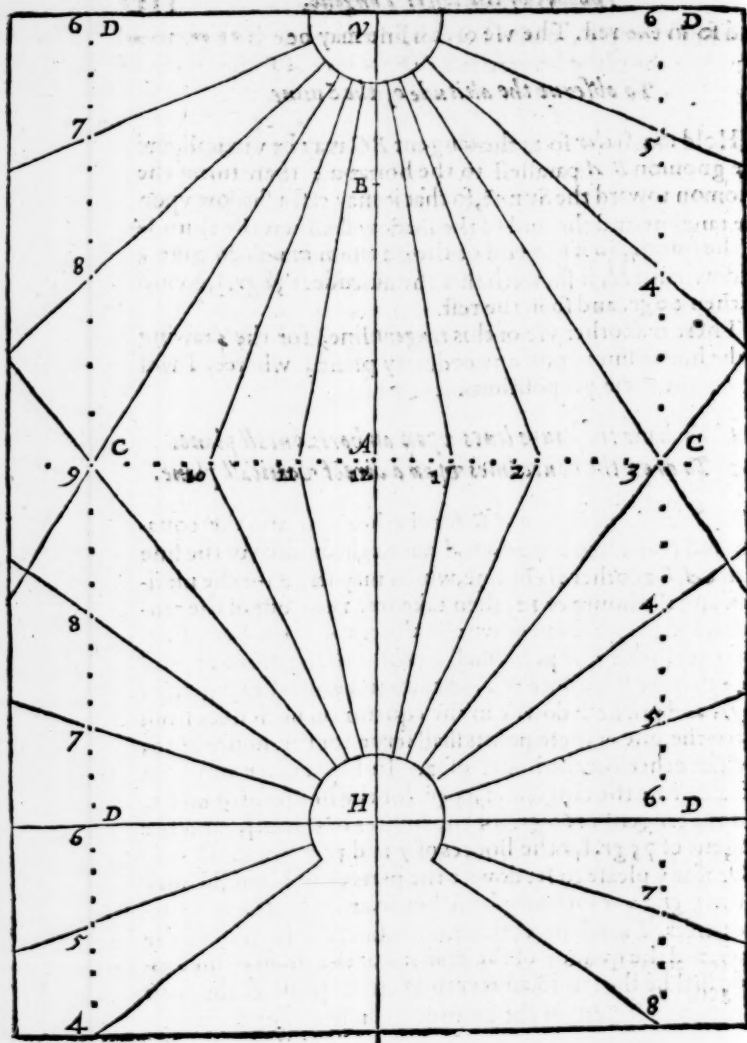
- 1 To draw the houre lines vpon an horizontall plane.
- 2 To draw the houre lines vpon a direct verticall plane.

First draw a right line *AC* for the horizon and the equator, and crosse it at the point *A* about the middle of the line with *AB* another right line, which may serue for the meridian and the houre of 12; then take out 15 gr. out of the tangents, and pricke them downe in the equator on both sides from 12: so the one point shall serue for the houre of 11, and the other for the houre of 1. Againe, take out the tangent of 30 gr. and pricke it downe in the equator on both sides from 12: so the one of these points shall serue for the houre of 10, and the other for the houre of 2. In like maner may you prick downe the tangent of 45 gr. for the houres of 9 and 3, and the tangent of 60 gr. for the houres of 8 and 4, and the tangent of 75 gr. for the houres of 7 and 5.

Or if any please to set downe the parts of an houre, he may allow 7 gr. 30 m. for euery halfe houre, and 3 gr. 45 m. for euery quarter. This done, you are to consider the latitude of the place, and the qualitie of the plane: For the *secant* of the latitude shal be the semidiameter in a verticall plane, & the *secant* of the complement of the latitude in an horizontall plane.

T

For



For example, about London the latitude is $51^{\circ} 30'$ and let the plane be verticall. If you take AV the secant of $51^{\circ} 30'$ out of the *Sector*, and prick it downe in the meridian line from A vnto V , the point V shall be the center: and if you draw right lines from V vnto 11 , and 10 , and the rest of the houre points, they shall be the houre lines required.

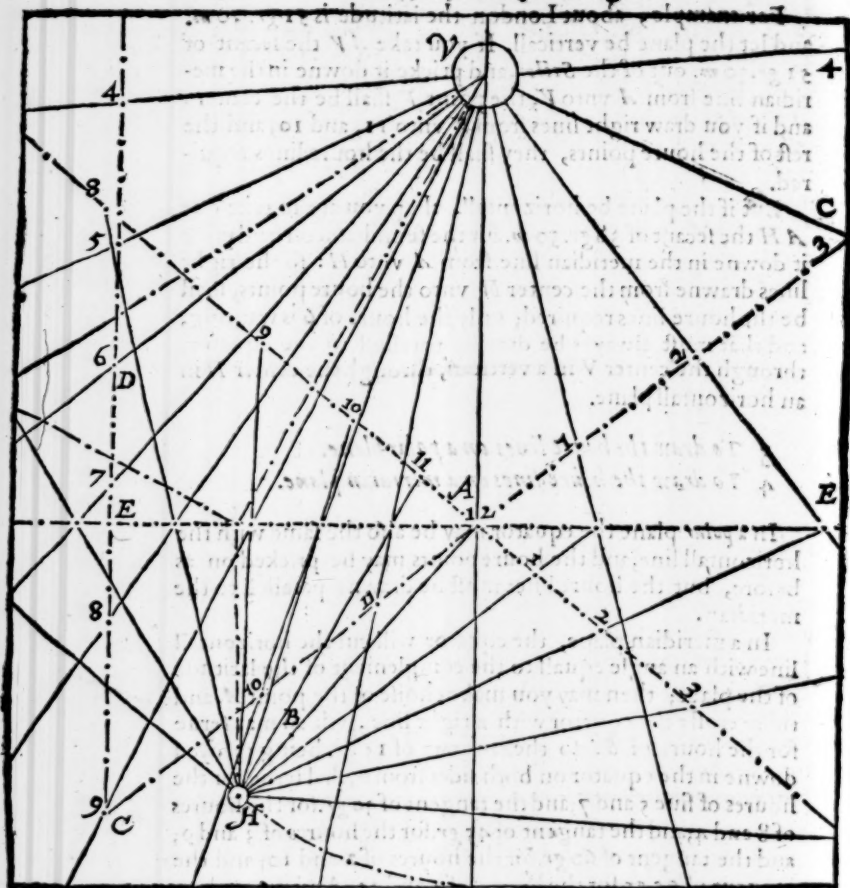
But if the plane be horizontall, then you are to take out AH the secant of $38^{\circ} 30'$ for the semidiameter, and prick it downe in the meridian line from A vnto H : so the right lines drawne from the center H vnto the houre points, shall be the houre lines required; only the houre of 6 is wanting, and that must alwayes be drawne parallell to the equator, through the center V in a verticall, through the center H in an horizontall plane.

3 To draw the houre lines on a polar plane.

4 To draw the houre lines on a meridian plane.

In a polar plane the equator may be also the same with the horizontall line, and the houre points may be pricked on as before, but the houre lines must be drawne parallell to the meridian.

In a meridian plane, the equator will cut the horizontall line with an angle equall to the complement of the latitude of the place; then may you make choise of the point A , and there crosse the equator with a right line, which may serue for the houre of 6: so the tangent of 15° being pricked downe in the equator on both sides from 6, shal serue for the houres of five 5 and 7; and the tangent of 30° for the houres of 8 and 4; and the tangent of 45° for the houres of 3 and 9; and the tangent of 60° for the houres of 2 and 10; and the tangent of 75° for the houres of 1 and 11. And if you draw right lines through these houre points, crossing the equator at right angles, they shall be the houre lines required.



5 To draw the houre lines in a verticall declining place,

First, draw *AV* the meridian, and *AE* the horizontal line, crossing one the other at right angles in the point *A*.

2 Then

2 Then take out AV , the secant of the latitude of the place, which you may suppose to be $51^{\circ} 30'$. and prick it downe in the meridian line from A vnto V .

3 Because it is a declining plane, and you may suppose it to decline 40° . Eastward, you are to make an angle of the declination vpon the center A , below the horizontall line, and to the left hand of the meridian line, because the declination is Eastward, for otherwise it should haue bin to the right hand, if the declination had bin Westward.

4 Take AH , the secant of the complement of the latitude out of the *Sector*, & pricke it downe in the line of declination from A vnto H , as you did before for the semidiameter in the horizontall plane.

5 Draw a line at full length through the point A , which must be perpendicular vnto AH , and cut the horizontall line according to the angles of declination, and it will be as the equator in the horizontall plane.

6 Take the houre points out of the *Tangent* line in the *Sector*, and pricke them downe in this equator on both sides from the houre of 12 at A .

7 Lay your ruler, & draw right lines through the center H , & each of these houre points: so haue you all the houre lines of an horizontall plane, onely the houre of 6 is wanting, and that may be drawne through H perpendicular to HA .

Lastly you are to obserue and marke the interfections, which these houre lines do make with AE the horizontall line of the plane: and then if you draw right lines through the center V , and each of these interfections, they shal be the houre lines required.

6 To pricke downe the houre points another way.

Hauing drawne a right line for the equator as before, and made choice of the point A , for the houre of 12: you may at pleasure cut of two equal lines $A10$, and $A2$. Then vpon the distance betweene 10 and 2, make an equilaterall triangle, and you shall haue B for the center of your equator, and the

line *AB* shall giue the distance from *A* to 9, and from *A* to 3. That done take out the distance betweene 9 and 3, and this shall giue the distance from *B* vnto 8, and from 8 vnto 7, and from 8 vnto 1: and againe from *B* vnto 4, and from 4 vnto 5, and from 4 vnto 11. So haue you the houre points, and if you take out the distance *B* 1, *B* 3, *B* 5, &c. You may finde the points not onely for the halfe houres, but also for the quarters.

But if it so fall out, that some of these houre points fall out of your plane, you may helpe your selfe by the larger *tangent*, both in the verticall, and horizontall planes.

For if at the houre points of 3 and 9, you draw occult lines parallell to the meridian; the distances *DC*, betweene the houre line of 6, and the houre points of 3 and 9, will be equal to the semidiameter *AV* in a verticall, and *AH* in a horizontall plane, and if they be diuided in such sort as the line *AC* is diuided, you shall haue the points of 4, and 5, and 7, and 8, with their halfes and quarters.

As in the horizontall plane, take out the semidiameter *AH*; and make it a parallell Radius by fitting it ouer in the *sines* of 90 and 90: Then take 15 gr. out of the larger *tangent*, and lay them on the lines of *sines*, where they will reach from the center vnto the *sines* of 15 gr. 32 m. therefore take out the parallell line of 15 gr. 32 m. and it shall giue the distance from 6 vnto 5, and from 6 vnto 7, in your horizontall plane. That done take out 30 gr. out of the larger *tangent*, and lay them on the *sines*, from the center vnto the *sines* of 35 gr. 16 m. and the parallell line of 35 gr. 16 m. shall giue you the distance from 6 vnto 4, and from 6 vnto 8, in your horizontall plane. The like may be done for the halfe houres and quarters.

So also in the verticall declining plane. If you first take out the *secant* of the declination of the plane, and prick it downe in the horizontall line from *A* vnto *E*, and through *E* draw right lines parallell to the meridian, which will cut the former houre lines of 3 and 9, or one of them in the point *C*: then take out the semidiameter *AV*, and prick it downe in those

those parallels from C vnto D, and draw right lines from A vnto C, and from V vnto D; the line V D shall be the houre of 6, and if you diuide these lines A C and D C, in such sort as you diuided the like line D C in the horizontall plane, you shall haue all the houre points required.

Or you may find the point D, in the houre of 6, without knowledge either of H or C. For hauing prickt downe A V in the meridian line, and A E in the horizontall line, and drawne parallels to the meridian through the points at E, you may take the *tangent* of the latitude out of the *Sector*, and fit it ouer in the lines of 90 and 90: so the parallell line of the declination measured in the same *tangent* line, shall there shew the complement of the angle DVA, which the houre line of 6 maketh with the meridian; then hauing the point D, take out the semidiameter VA, and prick it downe in those parallels from D vnto C: so shall you haue the lines D C and A C to be diuided as before.

The like might be vsed for the houre lines vpon all other planes. But I must not write all that may be done by the *Sector*. It may suffice that I haue wrote something of the vse of each line, and thereby giuen the ingenuous Reader occasion to thinke of more.

The conclusion to the Reader.

IT is well knowne to many of you, that this *Sector* was thus contriued, the most part of this booke written in Latin, many copies transcribed and dispersed more then sixteene yeares since. I am at the last contented to giue way that it come forth in English. Not that I thinke it worthy either of my labour or the publique view, but partly to satisfy their importunity, who not understanding the Latin, yet were at the charge to buy the instrument, and partly for my owne ease. For as it is painefull for others to transcribe my copie, so it is trouble some for me to giue satisfaction herein to all that desire it. If I finde this to giue you content, it shall incourage me to do the like for my Crosse-staffe, and some other Instruments. In the meane time beare with the Printers faults, and so I rest.

Gresham Coll. 1. Maij. 1623.

E. G.

FINIS.

V
D

Bridgewater ex dono Authoris.
THE

DESCRIPTION
AND VSE OF THE
CROSSE-STAFFE.

For such as are studious of
Mathematicall practise.



LONDON,
Printed by WILLIAM IONES.
and are to be sold by JOHN TAP at Saint
Magnus corner. 1623.

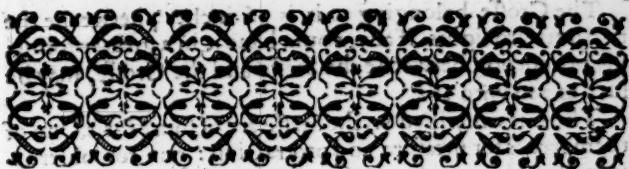
Amphiprion ex *Amphiprion* *Amphiprion*

DESCRIPTION

THE USE OF THE

AMPHIPRION

AMPHIPRION



THE
FIRST BOOKE OF THE
CROSSE-STAFFE.

CHAP. I.

Of the description of the Staffe.



He *Crosse-Staffe* is an instrument wel knowne to our Sea-men, and much vsed by the ancient Astronomers and others, seruing Astronomically for obseruation of altitude and angles of distance in the heauens, Geometrically for perpendicular heights and distances on land and sea.

The description and seuerall vses of it are extant in print, by *Gemma Frisius* in Latin, in English by *Dr. Hood*. I differ something from them both, in the proiection of this *Staffe*, but so, as their rules may be applied vnto it, and all their propositions be wrought by it: and therefore referring the Reader to their bookes, I shal be brieue in the explanation of that which may be applied from theirs vnto mine, and so come to the vse of those lines which are of my addition, not extant heretofore.

The necessary parts of this Instrument are five: the *Staffe*, the *Crosse*, and the three *sights*. The *Staffe* which I made for my owne vse, is a full yard in length, thar so it may serue for measure.

The description of the lines.

The Crosse belonging to it is 26 inches⁷ betweene the two outward sights. If any would haue it in a greater forme, the proportion betweene the Staffe and the Crosse, may be such as 360 vnto 262.

The lines inscribed on the Staffe are of foure sorts. One of them serues for measure and protraction: one for obseruation of angles: one for the Sea-chart; and the foure other for working of proportions in seuerall kinds.

The line of measure is an *inch line*, and may be knowne by his equall parts. The whole yard being diuided equally into 36 inches, and each inch subdiuided, first into ten parts, and then each tenth part into halves.

The line for obseruation of angles may be knowne by the double numbers set on both sides of the line, beginning at the one side at 20, and ending at 90: on the other side at 40, and ending at 180: and this being diuided according to the degrees of a quadrant, I call it the *tangent line on the Staffe*.

The next line is the meridian of a Sea-chart, according to *Mercators* projection from the Equinoctiall to 58 gr. of latitude, and may be knowne by the letter *M*, and the numbers 1. 2. 3. 4. vnto 58.

The lines for working of proportions, may be knowne by their vnequall diuisions, and the numbers at the end of each line.

1 The line of *numbers* noted with the letter *N*, diuided vnequally into 1000 parts, and numbred with 1. 2. 3. 4. vnto 10.

2 The line of *artificiall tangents* is noted with the letter *T*, diuided vnequally into 45 degrees, and numbred both ways, for the Tangent and the complement.

3 The line of *artificiall sines*, noted with the letter *S*, diuided vnequally into 90 degrees, and numbred with 1. 2. 3. 4. vnto 90.

4 The line of *versed sines* for more easie finding the houre and azimuth, noted with *V*, diuided vnequally into about 164 gr. 50 m. numbred backward with 10. 20. 30. vnto 164.

Thus there are seuen lines inscribed on the Staffe: there are fise lines more inscribed on the Crosse.

The inscription of the lines.

1 A Tangent line of 36 gr. 3 m. numbred by 5. 10. 15. vnto 35: the midst whereof is at 20 gr; and therefore I call it the *tangent of 20*; and this hath respect vnto 20 gr. in the Tangent on the Staffe.

2 A Tangent line of 49 gr. 6 m. numbred by 5. 10. 15. vnto 45; the midst whereof is at 30 gr. and hath respect vnto 30 gr. in the Tangent on the Staffe, whereupon I call it the *tangent of 30*.

3 A line of *inches* numbred with 1. 2. 3. vnto 26; each inch equally subdiuided into ten parts, answerable to the inch line vpon the Staffe.

4 A line of seuerall *chords*, one answerable to a circle of twelue inches semidiameter, numbred with 10. 20. 30. vnto 60; another to a semidiameter of a circle of six inches; and the third to a semidiameter of a circle of three inches; both numbred with 10. 20. 30. vnto 90.

5 A continuation of the *meridian* line from 57 gr. of latitude vnto 76 gr; and from 76 gr. to 84 gr.

For the inscription of these lines. The first for measure is equally diuided into inches and tenth parts of inches.

The tangent on the Staffe for obseruation of angles, with the tangent of 20 and the tangent of 30 on the Crosse, may all three be inscribed out of the ordinary *table of tangents*. The Staffe being 36 inches in length; the Radius for the tangent on the Staffe will be 13 inches and 103 parts of 1000: so the whole line will be a tangent of 70 gr. and must be numbred by their complements, & the double of their complements, the tangent of 10 gr. being numbred with 80 and 160.

The Radius for the tangent of 20 on the Crosse, will be 36 inches, and the whole line betweene the sights a tangent of 36 gr. 3 m. according as it is numbred. The Radius for the tangent of 30 gr. on the Crosse, will be 22 inches and 695 parts of 1000: so the whole line betweene the sights will containe a tangent of 49 gr. 6 m. in such sort as they are numbred.

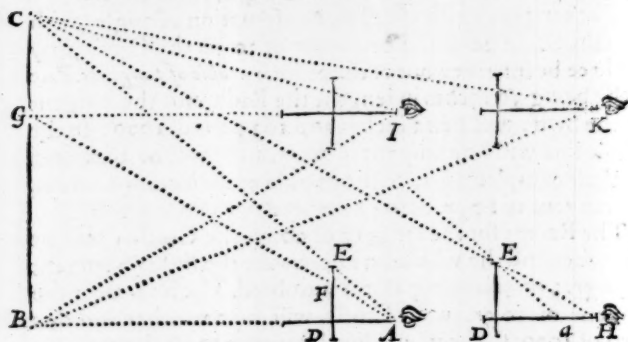
The meridian line may be inscribed out of the Table which I set downe for this purpose in the vse of the Sector.

The line of numbers may be inscribed out of the first Chi-
liad of Mr. *Briggs* Logarithmes: & the rest of the lines of pro-
portion out of my *Canon of artificiall sines and tangents*; and in
recompence thereof this booke will serue as a comment to
explain the vse of my *Canon*.

CHAP. II.

*The vse of the lines of inches for perpen-
dicular heights and distances.*

IN taking of heights and distances, the Staffe may be held
in such sort, that it may be euen with the distance, and
the Crosse parallell with the height: and then if the eye at
the beginning of the Staffe shall see his marks by the inward
sides of the two first sights, there will be such proportion be-
tweene the distance and the height, as is betweene the parts
intercepted on the Staffe and the Crosse. Which may be far-
ther explained in these propositions.



I To find an height at one station, by knowing
the distance.

Set the middle sight vnto the distance vpon the Staffe;
the

the height will be found vpon the Crosse. For

As the segment of the Staffe

vnto the segment on the Crosse:

So is the distance giuen,

vnto the height.

As if the distance AB being knowne to be 256 feete, it were required to find the height BC : first I place the middle sight at 25 inches and 6 parts of 10, then holding the Staffe leuell with the distance, I raise the Crosse, parallell vnto the height, in such sort, as that my eye may see from A the beginning of the inches on the Staffe by the sight E , at the beginning of the inches on the Crosse vnto the mark G which being done, if I find 19 inches and 2 parts of 10 intercepted on the Crosse betweene the sights at E and D , I would say the height BC were 192 feete.

Or if the obseruation were to be made before the distance were measured, I would set the middle sight either vnto 10 inches, or 12, or 16, or 20, or 24, or some such other number as might best be diuided into severall parts, and then worke by proportion. As if in the former example the middle sight were at 24 on the Staffe, and 18 on the Crosse, it should seeme that the height is $\frac{3}{4}$ of the distance; and therefore the distance being 256, the height should be 192.

2 To finde an height, by knowing some part of the same height.

As if the height from G to C were knowne to be 48, and it were required to find the whole height BC : either put the third sight or some other running sight vpon the Crosse betweene the eye and the marke G . For then

As the difference betweene the sights,

vnto the whole segment of the Crosse:

So is the part of the height giuen,

vnto the whole height.

If then the difference betweene the sights E and F , shall

a 3

bc

be 45, and the segment of the Crosse ED 180, the whole height BC will be found to be 192.

3 To find an height at two stations, by knowing the difference of the same stations.

As the difference of segments on the Staffe,
vnto the difference of stations:

So is the segment of the Crosse,
vnto the height.

Suppose the first station being at H , the segment of the Crosse ED were 180, and the segment of the Staffe HD 300: then coming 64 fecte nearer vnto B , in a direct line, vnto a second station at A , and making another obseruation; suppose the segment of the Crosse ED were 180 as before, and the segment of the Staffe AD 240; take 240 out of 300, the difference of segments will be 60 parts. And

As 60 parts vnto 64 the difference of stations:

So DE 180 vnto BC 192 the height required.

In these three *Prop.* there is a regard to be had of the height of the eye. For the height measured, is no more then from the leuell of the eye vpward.

4 To find a distance, by knowing the height.

As the segment of the Crosse,
vnto the segment of the Staffe:

So is the height giuen,
vnto the distance.

So the segment ED being 18, and DA 24, the height CB 192, will shew the distance AB to be 256.

5 To find a distance, by knowing part of the height.

As the difference betweene the sights,
vnto the segment of the Staffe:

So is the part of the height giuen,
vnto the distance.

And thus the difference betweene *E* and *F* being 45, and
the segment *D A* 240, the part of the height *GC* 48, will
giue the distance *AB* to be 256.

*6 To finde a distance at two stations, by knowing
the difference of the same stations.*

As the difference of segments on the Staffe,
vnto the difference of stations:

So is the whole segment,
vnto the distance.

And thus the segment of the Crosse being 180, the seg-
ment of the Staffe at the first station 240, at the second 300,
the difference of the segments 60, & the difference of stations
64, the distance *AB* at the first station will be found to be
256, and the distance *HB* at the second station 320.

*7 To find a bredth by knowing the distance per-
pendicular to the bredth.*

This is all one with the first *Prop.* For this bredth is but
an height turned sideways: and therefore

As the segment of the Staffe,
vnto the segment of the Crosse:

So is the distance
vnto the bredth.

And thus the segment of the Staffe being 24, and the seg-
ment of the Crosse 18, the distance *AB* 256, will giue the
bredth *BC* to be 192.

*8 To find a bredth at two stations in a line perpen-
dicular to the bredth, by knowing the diffe-
rence of the same stations.*

This is also the same with the third *Prop.* and therefore

As the difference of segments on the Staffe,

vnto the difference of Stations:

So the segment on the Crosse betwene the two lights,
vnto the breadth required.

And thus the difference betweene the Stations at *A* and *H*
being 64, the difference of segments on the Staffe 60, the
segment of the Crosse 180, the breadth *BC* will be found
to be 192.

In like maner may we finde the breadth *GC* for hauing
found the breadth *BC* the proportion will hold.

As *DE* is vnto *FE*, so *BC* vnto *GC*. Or otherwise,

As *HA* vnto *HA*, so *FE* vnto *GC*.

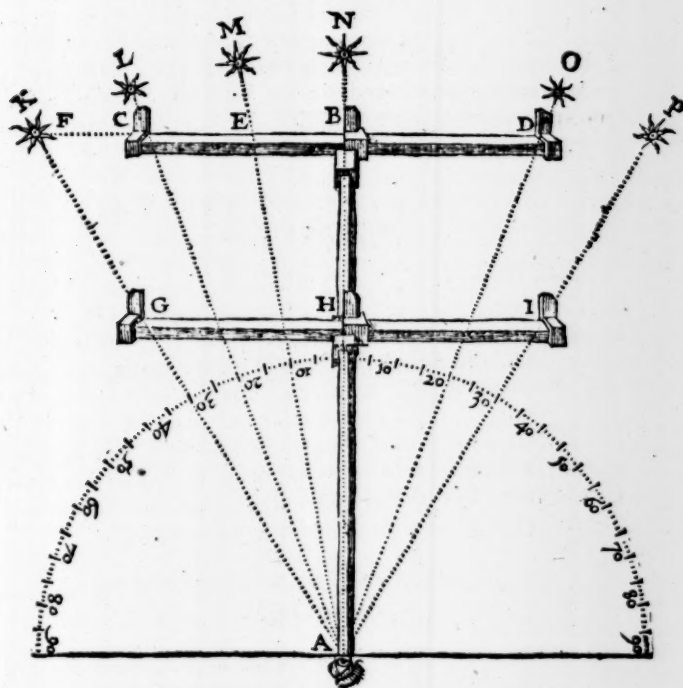
Neither is it materiall whether the two stations be chosen
at the one end of the breadth proposed, or without it, or with-
in it, if the line betweene the stations be perpendicular vnto
the breadth: as may appeare if in stead of the stations at *A* and
H, we make choise of the like stations at *I* and *K*.

There might be other wayes proposed to work these *Prop.*
by holding the Crosse euen with the distance, and the Staffe
parallell with the height: but these would proue more trou-
blesome, and those which are deliuered are sufficient, and
the same with those which others haue set downe vnder the
name of the *Jacobs staffe*.

The use of the Tangent lines.

CHAP. III.

*The use of the Tangent lines
in taking of Angles.*



I *To find an angle by the Tangents
on the Staffe.*

L Et the middle sight be alwayes set to the middle of the
Crosse, noted with 20 and 30, and then the Crosse
drawne

drawne nearer the eye, vntill the marks may be seene close within the sights. For so if the eye at *A* (that end of the Staffe which is noted with 90 and 180) beholding the marks *K* and *N*, betweene the two first sights, *C* and *B*, or the marks *K* and *P* betweene the two outward sights, the Crosse being drawne downe vnto *H*, shall stand at 30 and 60, in the Tangent on the Staffe: it sheweth that the angle *KAN* is 30 gr. the angle *KAP* 60 gr. the one double to the other; which is the reason of the double numbers on this line of the Staffe: and this way wil serue for any angle from 20 gr. toward 90 gr. or from 40 gr. toward 180 gr. But if the angle be lesse then 20 gr. we must then make vse of the Tangent vpō the Crosse.

**2 To find an angle by the Tangent of 20
upon the Crosse.**

Set 20 vnto 20, that is, the middle sight to the middest of the Crosse at the end of the Staffe, noted with 20: so the eye at *A*, beholding the marks *L* and *N*, close betweene the two first sights, *C* and *B*, shall see them in an angle of 20 gr.

If the marks shall be nearer together, as are *M* and *N*, then draw in the Crosse from *C* vnto *E*: if they be farther asunder, as are *K* and *N*, then draw out the Crosse from *C* vnto *F*; so the quantitie of the angle shal be still found in the Crosse in the Tangent of 20 gr. at the end of the Staffe; and this will serue for any angle from 0 gr. toward 35 gr.

**3 To find an angle by the Tangent of 30
upon the Crosse.**

This Tangent of 30 is here put the rather, that the end of the Staffe resting at the eye, the hand may more easily remoue the Crosse: for it supposeth the Radius to be no longer then *AH*, which is from the eye at the end of the Staffe vnto 30 gr. about 22 inches and 7 parts: Wherefore here set the middle sight vnto 30 gr. on the Staffe, and then either draw the Crosse in or out, vntill the marks be seene between the

the two first sights; so the quantitie of the angle will be found in the Tangent of 30, which is here represented by the line *GH*; and this will serue for any angle from 0 gr. toward 48 gr.

4 To obserue the altitude of the Sunne backward.

Here it is fit to haue an horizontall sight set to the beginning of the Staffe, and then may you turne your backe toward the Sun, and your Crosse toward your eye. If the altitude be vnder 45 gr. set the middle sight to 30 on the Staffe, and looke by the middle sight through the horizontall vnto the horizon, mouing the Crosse vpward or downward, vntill the vpper sight doe shadow the vpper halfe of the horizontall sight: so the altitude will be found in the Tangent of 30.

If the altitude shal be more then 45 gr. set the middle sight vnto the midst of the Crosse, and look by the inward edge of the lower sight through the horizontall to the horizon, mouing the middle sight in or out, vntill the vpper sight do shadow the vpper halfe of the horizontall sight: so the altitude will be found in the degrees on the Staffe betweene 40 and 180.

5 To set the Staffe to any angle giuen.

This is the conuerse of the former *Prop.* For if the middle sight be set to his place and degree, the eye looking close by the sights as before, cannot but see his object in the angle giuen.

6 To obserue the altitude of the Sunne another way.

Set the middle sight to the middle of the Crosse, and hold the horizontall sight downward, so as the Crosse may be parallel to the horizon, then is the Staffe verticall; and if the outward sight of the Crosse do shadow the horizontall sight,

the complement of the altitude will be found in the tangent on the Staffe.

7 To obserue an altitude by thread and plummet.

Let the middle sight be set to the middest of the Crosse, and to that end of the Staffe which is noted with 90 and 180; then hauing a thread and a plummet at the beginning of the Crosse, and turning the Crosse vpward, and the Staffe toward the Sunne, the thread will fall on the complement of the altitude about the horizon. And this may be applied to other purposes.

8 To apply the lines of inches to the taking of angles.

If the angles be obserued betweene the two first sights, there will be such proportion between the parts of the Staffe and the parts of the Crosse, as betweene the Radius and the Tangent of the angle.

As if the parts intercepted on the Staffe were 20 inches, the parts on the Crosse 9 inches. Then by proportion as 20 vnto 9, so 100000 vnto 45000 the tangent of 24 gr. 14 m.

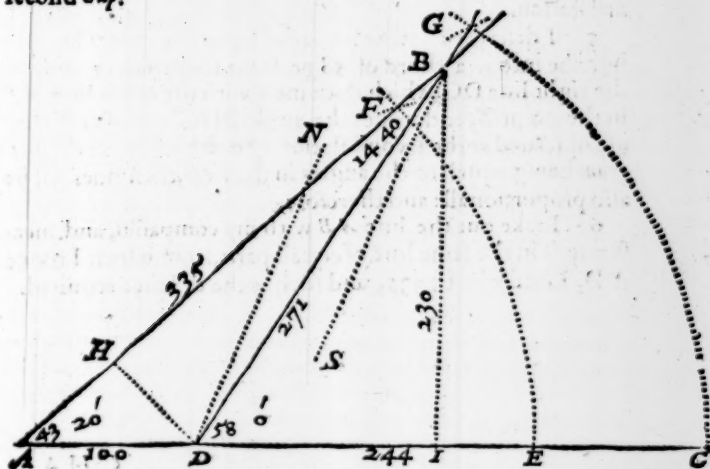
But if the angle shall be obserued betweene the two outward sights, the parts being 20 and 9 as before, the angle will be 48 gr. 28 m. double vnto the former.

In all these there is regard to be had to the parallax of the eye, and his height about the Horizon in obseruations at Sea; to the Semidiameter of the Sun, his parallax and refraction, as in the vse of other staues. And so this will be as much, or more then that which hath been heretofore performed by the Crosse-staffe.

CHAP. IIII.

The use of the lines of equall parts
ioyned with the lines of Chords.

THe lines of equall parts do serue also for protraction, as may appeare by the former *Diagrams*; but being ioyned with the lines of Chords, which I place vpon one side of the Crosse, they will farther serue for the protraction and resolution of right line triangles; whereof I will giue one example in finding of a distance at two stations otherwise then in the second *Cap.*



Let the distance required be AB . At A the first station I make choise of a station line toward C , and obserue the angle BAC by the tangent lines, which may be $43\text{ gr. }20\text{ m}$; then hauing gon an hundred paces toward C , I make my second station at D , where suppose I find the angle BDC to be 58 gr. or

b 3

the

the angle BDA to be 122 gr. ; this being done, I may finde the distance AB in this manner.

1 I draw a right line AC , representing the station line.

2 I take 100 out of the lines of equall parts, and pricke them downe from A the first station vnto D the second.

3 I open my compasses to one of the chords of 60 gr. and setting one foote in the point A , with the other I describe an occult arke of a circle intersecting the station line in E .

4 I take out of the same line of *chords* a chord of $43\text{ gr. } 20\text{ m.}$ (because such was the angle at the first station) and this I inscribe into that occult arke from E vnto F , which makes the angle FAD equall to the angle obserued at the first station.

5 I describe another like arke vpon the center D , and inscribe into it a chord of 58 gr. from C vnto G , and draw the right line DG , which doth meet with the other line AF in the point B , and makes the angle BDC equall to the angle obserued at the second station. So the angles in the *Diagram* being equall to the angles in the field, their sides will be also proportionall: and therefore,

6 I take out the line AB with my compasses, and measuring it in the same line of equall parts, from which I rooke AD , I find it to be 335, and such is the distance required.

CHAP. V.

The use of the Meridian line.

1 **T**He Meridian line, noted with the letter *M*, may serue for the more easie diuision of the plane sea-chart, according to *Mercators* projection. For if you shall draw parallell meridians, each degree being halfe an inch distant from other, the degrees of this meridian line on the Staffe, shall giue the like degrees for the meridians on the chart, from the Equinoctiall toward the Pole: and then if through these degrees you draw streight lines perpendicular to the meridians, they shall be parallels of latitude.

If any desire to haue the degrees of his chart larger then those which I haue put on the Staffe, he may take these and increase them in a double, or treble, or a decuple proportion at his pleasure.

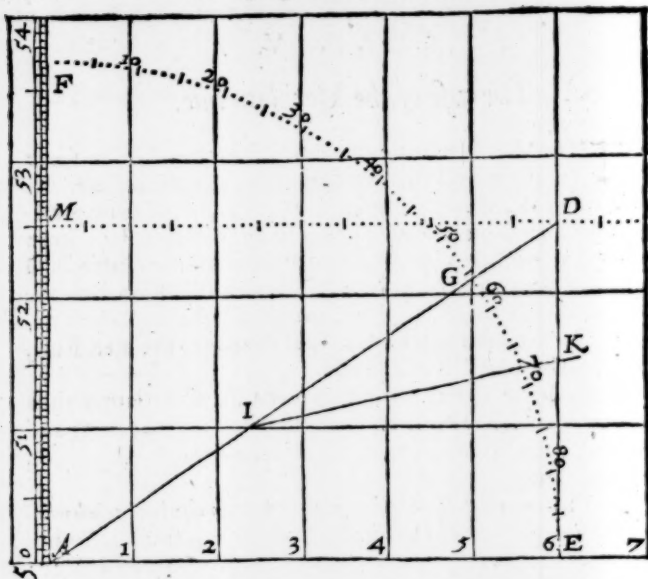
2 This *meridian* line being ioyned with the line of *chords*, may serue for the protraction & resolution of such right line triangles as concerne latitude, longitude, rumb and distance in the practise of nauigation. As may appeare by this example.

Suppose two places giuen, *A* in the latitude of 50 gr. D in the latitude of 52 gr. $\frac{1}{2}$, the difference of longitude betweene them being 6 gr. and let it be required to know, first what Rumb leadeth from the one place to the other; secondly how many degrees distant they are asunder.

1 I draw a right line *AE*, representing the parallell of the place from whence I depart.

2 I take 6 gr. for the difference of longitude, either out of the line of *inches*, allowing halfe an inch for euery degree, or out of the beginning of the *meridian* line; (for there the meridian degrees differ very little from the equinoctiall degrees) and these 6 gr. I pricke downe in the parallell from *A* to *E*.

3 In *A* and *E*, I erect two perpendiculars, *AM* and *ED*, representing the meridians of both places.



4 I take the difference of latitude from 50 gr. to 52 gr. 30 m. out of the *meridian* line, and prick it downe in the meridians from *A* vnto *M*, and from *E* to *D*, and draw the right line *MD* for the parallell of the second place, and the right line *AD* for the line of distance betweene both places: so the angle *MAD* shall giue the Rumb that leadeth from the one place to the other.

5 To finde the quantitie of this angle *MAD*, I may either make vse of the *Protractor*, or else of a line of *chords*, and so I open my compasses vnto one of the chords of 60 gr. and setting one foote in the point *A*, with the other I describe one foot of a circle, intersecting the meridian in *F*, and the line of distance in *G*; then I take the chord *FG* with my compasses, and measuring it in the same line of *chords* as before, I find it 56 gr. $\frac{1}{2}$; and such is the inclination of the

the Rumb to the meridian, which is the first thing that was required.

6 To find the quantitie of the line of distance *A D*, I take it out with my compasses, and measuring it in the meridian line, setting one foote beneath the lesser latitude, and the other foote as much above the greater latitude, I find about $4\text{ gr. } \frac{1}{2}$ intercepted between both feet: and such is the distance vpon the Rumb, which is the second thing that was required.

But if this example were prorracted according to the common Sea-chart, where the degrees of the equinoctiall and meridian are both alike; the Rumb *M A D* would be found to be about 67 gr. and *A D* the distance vpon the Rumb about $6\text{ gr. } \frac{1}{2}$.

Suppose farther, that hauing set forth from *A* toward *D*, vpon the former Rumb of $56\text{ gr. } 15\text{ m. N E b E}$, after the ship had runne 36 leagues; the wind changing, it ran 50 leagues more vpon the seventh Rumb of *E b N*, whose inclination to the meridian is $78\text{ gr. } 45\text{ m.}$ And let it be required to know what longitude and latitude the ship is in, by pricking down the way thereof vpon the Chart.

Hauing drawne a blanke chart as before, with meridians & parallels, according to the latitude of the places proposed,

1 I would make an angle *M A D* of $56\text{ gr. } 15\text{ m.}$ for the Rumb of *N E b E*, which is done after this manner: I open my compasses to one of the chords of 60 gr. and setting one foote in the point *A*, with the other I describe an occult ark of a circle, intersecting the meridian in *F*; then I take $56\text{ gr. } 15\text{ m.}$ out of the same line of chords, and prick them downe from *F* vnto *G*: so the right line *A G* shall be the Rumb of *N E b E*.

2 I would take 36 leagues out of the meridian line, extending my compasses from 50 gr. to $51\text{ gr. } 48\text{ m.}$ or rather from as much below 50 as about 51 ; and prick them downe vpon the Rumb from *A* vnto *I*; so the point *I*, shal represent the place wherein the ship was when the wind changed. And this is in the latitude of $51\text{ gr. } 0\text{ m.}$ and in the longitude of $2\text{ gr. } 21\text{ m.}$ Eastward from the meridian *A M*.

3 By the same reason, I may draw the right line *I K* for the Rumb of *E b N*, and pricke downe the distance of 50 leagues from *I* vnto *K*: so the point *K* shal represent the place whither the ship came, after the running of these 50 leagues: and this is in the latitude of 51 gr. 30 m. and in longitude 6 gr. 16 m. Eastward from the first meridian *A M*, and therefore 16 m. Eastward from the second meridian *E D*.

But if these two courses were to be pricked downe by the common sea-chart, the point *I* would fall in the latitude of 51 gr. 0 m. and the point *K* in the latitude of 51 gr. 30 m. But the longitude of *I* would be onely 1 gr. 30 m. and the longitude of *K* onely 3 gr. 57 m. which is 33 m Westward from the meridian of the place to which the ship was bound.

Such is the difference betweene both these charts.

CHAP. VI.

The use of the line of Numbers.

1 *Having two numbers giuen to find a third in continuall proportion, a fourth, a fift, and so forward.*

Extend the compasses from the first number vnto the second; then may you turne them, from the second to the third, and from the third to the fourth, and so forward.

Let the two numbers giuen be 2 and 4. Extend the compasses from 2 to 4, then may you turne them from 4 to 8, and from 8 to 16, and from 16 to 32, and from 32 to 64, and from 64 to 128.

Or if the one foote of the compasses being set to 64, the other fall out of the line, you may set it to another 64 nearer the beginning of the line, and there the other foot will reach to 128, and from 128 you may turne them to 256, and so forward.

Or if the two first numbers giuen were 10 and 9: extend the compasses from 10 at the end of the line, backe vnto 9, then may you turne them from 9 vnto 8.1, and from 8.1
vnto

vnto 7.29. And so if the two first numbers giuen were 1 and 9, the third would be found to be 81, the fourth 729, with the same extent of the compasses.

In the same maner, if the two first numbers were 10 and 12, you may finde the third proportionall to be 144, the fourth 1728. And with the same extent of the compasses, if the two first numbers were 1 and 12, the third would be found to be 144, and the fourth to be 1728.

*2 Having two extreme numbers giuen, to find
a meane proportionall betwene them.*

Diuide the space betweene the extreme numbers into two equall parts, and the foote of the compasses will stay at the meane proportionall. So the extreme numbers giuen being 8 and 32, the meane betweene them will be found to be 16, which may be proued by the former Prop. where it was shewed, that as 8 to 16, so are 16 to 32.

3 To find the square roote of any number giuen.

The square roote is alwayes the meane proportionall betweene 1 and the number giuen, and therefore to be found by diuiding the space betweene them into two equall parts. So the roote of 9 is 3, and the roote of 81 is 9, and the roote of 144 is 12.

*4 Having two extreme numbers giuen, to find
two meane proportionals between them.*

Diuide the space betweene the two extreme numbers giuen, into three equall parts. As if the extreme numbers giuen were 8 and 27, diuide the space betweene them into three equall parts, the feet of the compasses will stand in 12 and 18.

5 To find the cubique roote of a number giuen.

The cubique roote is alwayes the first of two meane pro-

proportionals betweene 1 and the number giuen, and therefore to be found by diuiding the space betweene them into three equall parts.

Thus the roote of 1728 will be found to be 12. The roote of 17280 is almost 26: and the roote of 172800 is almost 56.

6 To multiply one number by another.

Extend the compasses from 1 to the multiplicator; the same extent applied the same way, shall reach from the multiplicand to the product.

As if the numbers to be multiplied were 25 and 30: either extend the compasses from 1 to 25, and the same extent will giue the distance from 30 to 750; or extend them from 1 to 30, and the same extent shall reach from 25 to 750.

7 To diuide one number by another.

Extend the compasses from the diuisor to 1, the same extent shall reach from the diuidend to the quotient.

So if 750 were to be diuided by 25, the quotient would be found to be 30.

8 Three numbers being giuen to find a fourth proportionall.

This golden rule, the most usefull of all others, is performed with like ease. For extend the compasses from the first number to the second, the same extent shall giue the distance from the third to the fourth.

As for example, the proportion between the diameter and the circumference, is said to be such as 7 to 22: if the diameter be 14, how much is the circumference? Extend the compasses from 7 to 22, the same extent shall giue the distance from 14 to 44: or extend them from 7 to 14, and the same extent shall reach from 22 to 44.

Either of these wayes may be tried on seuerall places of
this

this line; but that place is best, where the scere of the compasses may stand nearest together.

*9 Three numbers being giuen to finde a fourth
in a duplicated proportion.*

This proposition concernes questions of proportion between *lines* and *superficies*; where if the denomination be of lines, extend the compasses from the first to the second number of the same denomination: so the same extent being doubled, shall giue the distance from the third number vnto the fourth.

The diameter being 14, the content of the circle is 154: the diameter being 28, what may the content be? Extend the compasses from 14 to 28, the same extent doubled will reach from 154 to 616. For first it reacheth from 154 vnto 308; and turning the compasses once more, it reacheth from 308 vnto 616: and this is the content required.

But if the first denomination be of the superficiall content, extend the compasses vnto the halfe of the distance, betweene the first number and the second of the same denomination: so the same extent shall giue the distance from the third to the fourth.

The content of a circle being 154, the diameter is 14: the content being 616, what may the diameter be? Diuide the distance betweene 154 and 616 into two equall parts, then set one foote in 14, the other will reach to 28 the diameter required.

*10 Three numbers being giuen to find a fourth
in a triplicated proportion.*

This proposition concerneth questions of proportion between *lines* and *solids*; where if the first denomination be of lines, extend the compasses from the first number to the second of the same denomination: so the extent being tripled, shall giue the distance frō the third number vnto the fourth.

Suppose the diameter of an iron bullet being 4 inches, the weight of it was 9 lb : the diameter being 8 inches, what may the weight be? Extend the compasses from 4 to 8, the same extent being tripled, will reach from 9 vnto 72. For first it reacheth from 9 vnto 18; then from 18 to 36; thirdly from 36 to 72. And this is the weight required.

But if the first denomination shall be of the Solid content, or of the weight, extend the compasses to a third part of the distance betweene the first number and the second of the same denomination: so the same extent shal giue the distance from the third number vnto the fourth.

The weight of a cube being 72 lb , the side of it was 8 inches: the weight being 9 lb , what may the side be? Diuide the distance betweene 72 and 9, into three equall parts; then set one foote to 8, the other will reach to 4, the side required.

CHAP. VII.

The vse of the lines of artificiall Sines.

THIS line of *sines* hath such vse in finding a fourth proportionall, as the ordinary *Canon of Sines*: and the manner of finding it, is alwayes such as in this example.

As the sine of 30 *gr.* vnto the sine of 52 *gr.*

So the sine of 38 *gr.* to a fourth sine.

Extend the compasses in the line of *sines* from 30 *gr.* vnto 52 *gr.*; the same extent shall giue the distance from 38 *gr.* vnto 76 *gr.* Or extend them from 30 *gr.* vnto 38 *gr.* the same extent will reach from 52 *gr.* vnto 76 *gr.* which is the fourth proportionall sine required.

And thus may the rest of all sinical proportions be wroughe two wayes. The minutes which are wanting in the first degree, may be supplied by the line of *Numbers*.

CHAP.

CHAP. VIII.

The Use of the line of artificiall Tangents.

THis line of *Tangents* hath like vse, but commonly ioyned with the line of *sines*: the maner of working by it, may appeare by this example.

As the Tangent of 38 gr. 30 m.
is to the Tangent of 23 gr. 30 m.
So the Sine of 90 gr.
to a fourth Sine.

This *Prop.* and such others vpon two lines, may be wrought two wayes. For extend the compasses from the Tangent of 38 gr. 30 m. to the Tangent of 23 gr. 30 m; the same extent shall giue the distance from the sine of 90 gr. to the sine of 33 gr. 8 m. Or else extend them from 38 gr. 30 m. in the Tangents vnto 90 gr. in the line of *sines*; the same extent from the Tangent of 23 gr. 30 m. shall reach to the sine of 33 gr. 8 m. which is the fourth proportionall sine required.

And this crossework in many cases is the better, in regard the tangents which should passe on from 40 gr. to 50 gr. and so forward, do turne backe at 45 gr. These two lines of *Sines* and *Tangents*, may serue for the resolution of all sphericall triangles, according to those Canons which I haue set downe in the vse of the *Sector*.

Or if at any time one meete with a *secant*, let him account the sine of 80 gr. for a *secant* of 10 gr. and the sine of 70 gr. for a *secant* of 20 gr. and so take the sine of the complement in stead of the *secant*. As if the proposition were,

As the Radius to the secant of 51 gr. 30 m.
So the sine of 23 gr. 30 m. to a fourth sine.

Extend the compasses from the Radius that is the sine of 90 gr. to the sine of 38 gr. 30 m. the same extent will giue the distance from the sine of 23 gr. 30 m. both to the sine of 14 gr.

22 *m.* and to the line of 39 *gr.* 50 *m.* But in this case, the sine of 39 *gr.* 50 *m.* is the fourth required. For the first number being lesse then the second, that is, the Radius lesse then the secant, the sine of 23 *gr.* 30 *m.* which is the third, must also be lesse then the fourth.

CHAP. IX.

The vse of the line of Sines and Tangents ioyned with the line of numbers.

THe lines of *sines* and *tangents* haue another like vse ioyned with the line of *numbers*, especially in the resolution of right line triangles, where the angles are measured by degrees and minutes, and the sides measured by absolute numbers, whereof I will set downe these propositions.

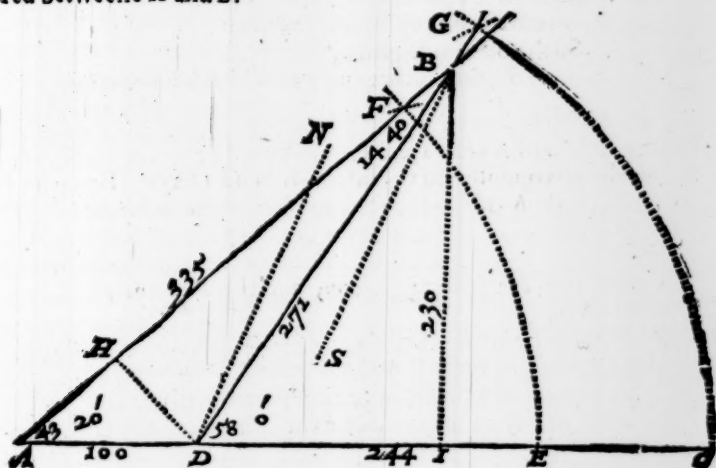
I Having three angles and one side, to find the two other sides.

As the sine of the angle opposite to the side given,
is to the number belonging to that side giuen:
So the sine of the angle opposite to the side required,
to the number belonging to the side required.

As in the example of the fourth *Cap.* of this booke, where knowing the distance betweene two stations at *A* and *D* to be 100 paces, the angle *BAC* to be 43 *gr.* 20 *m.* and the angle *BDC* to be 58 *gr.* it was required to find the distance *AB*.

First having these two angles, I may find the third angle *ABD* to be 14 *gr.* 40 *m.* either by subtraction or by complement vnto 180. Then in the triangle *BAD*, I haue three angles, and one side, whereby I may find both *AB* and *DB*. I know the angle *ABD* opposite to the measured side *AD* to be 14 *gr.* 40 *m.* and the angle *ADB* opposite to the side required, to be 12 2 *gr.* wherefore I extend the compalles in
the

the line of *sines* from 14 gr. 40 m. vnto 122 gr. or (which is all one) to 58 gr. (for after 90 gr. the sine of 80 gr. is also the sine of 100 gr. and the sine of 70 gr. the sine of 110 gr. and so in the rest) so shall I find the same extent to reach in the line of *numbers*, from 100 vnto 335. And such is the distance required between *A* and *B*.



In like manner if I extend my compasses from the sine of 14 gr. 40 m. to the sine of 43 gr. 20 m. the same extent will reach in the line of *numbers* from 100 to 271. And such is the distance between *D* and *B*.

Or in crosse worke, I may extend the compasses from 14 gr. 40 m. in the *sines*, vnto 100 parts in the line of *numbers*: so the same extent will giue the distance from 58 gr. to 335 parts, and from 43 gr. 20 m. to 271 parts.

- 2 Having two sides giuen, and one angle opposite to either of these sides, to find the other two angles and the third side.

As the side opposite to the angle giuen,
is to the sine of the angle giuen:

So the other side giuen,
is to the sine of that angle to which it is opposite.

So in the former triangle, hauing the two sides AB 335 paces, and AD 100 paces, and knowing the angle ADB , which is opposite to the side AB , to be 122 gr. I may find the angle ABD , which is opposite to the other side AD . For if I extend the compasses from 335 to 100 in the line of *numbers*, I shall finde the same extent to reach in the line of *sines* from 122 gr. to $14\text{ gr. }40\text{ m.}$ and therefore such is the angle ABD .

Then knowing these two angles ABD and ADB , I may find the third angle BAD either by subtraction or by complement to 180, to be $43\text{ gr. }20\text{ m.}$ and hauing three angles and two sides, I may well find the third side DB , by the former *Prop.*

This may be done more readily by crosse worke. For if I extend the compasses from 335 parts, in the line of *numbers*, to the sine of 122 gr. the same extent wil reach from 100 parts to the sine of $14\text{ gr. }40\text{ m.}$ and backe from $43\text{ gr. }20\text{ m.}$ to 271 parts; and such is the third side DB .

- 3 Having two sides and the angle between them, to find the two other angles and the third side.

If the angle contained betweene the two sides be a right angle, the other two angles will be found readily by this canon.

As the greater side giuen,
is to the lesser side:

So

So the tangent of 45 gr.
to the tangent of the lesser angle.

So in the rectangle triangle AIB , knowing the side AI to be 244, and the side IB to be 230: if I extend the compasses from 244 to 230 in the line of *numbers*, the same extent will reach from 45 gr. to about 43 gr. 20 m. in the line of *tangents*; and such is the lesser angle BAI , and the complement 46 gr. 40 m. shewes the greater angle ABI . The angles being knowne, the third side AB may be found by the first *Prop.*

So likewise in the example of the third *Cap.* of this booke, concerning taking of angles by the line of *inches*, where the parts intercepted on the Staffe being 20 *inches*, and the parts on the Crosse 9 *inches*, it was required to find the angle of altitude. For I may extend the compasses in the line of *numbers*, from 20 vnto 9, the same extent will reach in the line of *tangents*, from 45 gr. to 24 gr. 14 m. Or in the crosse worke, I may extend the compasses from 20 parts in the line of *numbers* to the tangent of 45 gr; the same extent shall giue the distance from 9 parts vnto the tangent of 24 gr. 14 m. And such is the angle of altitude required.

But if it be an oblique angle that is contained betweene the two sides giuen, the triangle may be reduced into two rectangle triangles, and then resolved as before.

As in the triangle ADB , where the side AB is 335, and the side AD 100, and the angle BAD 43 gr. 20 m: if I let downe the perpendicular DH vpon the side AB , I shal haue two rectangle triangles, AHD , DHB ; and in the rectangle AHD , the angle at A being 43 gr. 20 m. the other angle ADH will be 46 gr. 40 m; and with these angles and the side AD , I may find both AH and DH , by the first *Prop.* Then taking AH out of AB , there remains HB for the side of the rectangle DHB ; and therefore with this side HB and the other side DH , I may find both the angle at B , and the third side DB , as in the former part of this *Prop.*

Or I may find the angles required, without letting downe any perpendicular. For

28 *The use of the line of Sines and Tangents*

As the summe of the sides,

is to the difference of the sides:

So the tangent of the halfe summe of the opposite angles,
to the tangent of half the difference between those angles.

As in the former triangle ADB , the summe of the sides AB, AD , is 435, and the difference betweene them 235; the angle contained 43 gr. 20 m; and therefore the summe of the two opposite angles 136 gr. 40 m. and the halfe summe 68 gr 20 m. Hereupon I extend the compasses in the line of *numbers* from 455 to 235, and I find them to reach in the line of *tangents* from 68 gr. 20 m. vnto 53 gr. 40 m; and such is the halfe difference betweene the opposite angles at B and D . This halfe difference being added to the halfe summe, doth giue 122 gr. for the greater angle ADB : and being subtracted, it leaueth 14 gr. 40 m. for the lesser angle ABD . Then the three angles being knowne, the third side BD may be found by the first *Prop.*

4 *Having the three sides of a right line triangle, to find the perpendicular and the three angles.*

Let one of the three sides giuen be the base, but rather the greater side, that the perpendicular may fall within the triangle; then gather the summe, and the difference of the two other sides, and the proportion will hold.

As the base of the triangle,
is to the summe of the sides:

So the difference of the sides

to a fourth, which being taken forth of the base, the perpendicular shal fall on the middle of the remainder.

As in the former triangle ADB , where the base AB is 335, the summe of the sides AD and DB 371, and the difference of them 171. If I extend the compasses in the line of *numbers* from 335 vnto 371, I shall find the same extent to reach from 171 vnto 189.4. This fourth number I take out of the base

base 335.0, and the remainder is 145.6, the halfe whereof is 72.8, and doth shew the place *H*, where the perpendicular shall fall, from the angle *D*, vpon the base *AB*, diuiding the former triangle *ADB* into two right angle triangles, *DHA* and *DHB*, in which the angles may be found by the former part of the third *Prop.* And this may suffice for right line triangles. But for the more easie protraction of these triangles, I will set downe one proposition more concerning *chords*.

3 *Having the semidiameter of a circle, to find the chords of euery arke.*

As the sine of 30 gr.

to the sine of halfe the arke proposed:

So is the semidiameter of the circle giuen,

to the chord of the same arke.

As if in protracting the former triangle *ADB*, it were required to find the length of a chord of 43 gr. 20 m. agreeing to the semidiameter *AE*, which is knowne to be 3 inches. The halfe of 43 gr. 20 m. is 21 gr. 40 m; wherefore I extend the compasses from the sine of 30 gr. to the sine of 21 gr. 40 m. and I finde the same extent to reach in the line of *numbers* from 3.000 parts to 2.215; which shewes, that the semidiameter being 3 inches, the chord of 43 gr. 20 m. will be 2 inches and 215 parts of 1000.

In like maner the chord of 58 gr. agreeing to the same semidiameter, would be found to be 2 inches and 909 parts. For the halfe of 58 being 29; if I extend the compasses in the line of *sines* from 30 gr. to 29 gr. the same extent will reach in the line of *numbers* from 3.000. vnto 2.909.

Or in crosse worke, if I extend the compasses from the sine of 30 gr. to 3.000 in the line of *numbers*, I shall find the same extent to reach from 21 gr. 40 m. to 2.215 parts, and from 29 gr. to 2.909 parts, and from 7 gr. 20 m. to 765 parts; for the chord of 14 gr. 40 m. for the third angle *ABD*.

CHAP. X.

The use of the line of versed sines.

THis line of *versed sines* is no necessary line. For all triangles, both right lined and sphericall, may be resolved by the three former lines of *numbers*, *sines* and *tangents*; yet I thought good to put it on the Staffe for the more easie finding of an angle having three sides, or a side having three angles of a sphericall triangle given.

Suppose the three sides to be, one of them 110 gr. the other 78 gr. and the third $38\text{ gr. }30\text{ m.}$ and let it be required to find the angle, whose base is 110 gr.

I first add them together, and from halfe the summe subtract the base, noting the difference after this maner.

The base	110 gr. 0 m.
The one side	78 0
The other side	38 30
The summe of all three	226 30
The halfe summe	113 15
The difference	3 15

This done, I come to the Staffe, and extend the compasses from the sine of 90 gr. to the sine of 78 gr. which is one of the sides; and applying this extent from the sine of the other side $38\text{ gr. }30\text{ m.}$ I find it to reach to a fourth sine, about $37\text{ gr. }30\text{ m.}$ From this fourth sine of $37\text{ gr. }30\text{ m.}$ I extend the compasses again, to the sine of the halfe summe $113\text{ gr. }15\text{ m.}$ (which is all one with the sine of $66\text{ gr. }45\text{ m.}$) and this second extent will reach from the sine of the difference $3\text{ gr. }15\text{ m.}$ to the sine of $4\text{ gr. }54\text{ m.}$ Over against this sine you shall find 146 gr. in the line of *versed sines*; and such is the angle required.

THE

THE SECOND BOOKE.

*Of the vse of the former lines of proportiō,
more particularly exemplified
in seuerall kinds.*

THe former booke containing the generall vse of each line of proportion, may be sufficient for all those which know the rule of *Three*, and the doctrine of triangles.

But for others, I suppose it would be more difficult to find either the declination of the Sunne, or his amplitude, or the like, by that which hath been said in the vse of the line of *sines*, vnlesse they may haue the particular proportions, by which such propositions are to be wrought. And therefore for their sakes I haue adioyned this second booke, containing seuerall proportions for propositions of ordinary vse, and set them down in such order, that the Reader considering which is the first of the three numbers giuen, may easily apply them to the Sector, and also resolue them by Arithmetique, beginning with those which require help onely of the line of *numbers*.

CHAP. I.

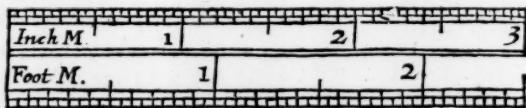
*The vse of the line of Numbers in broad measure, such as boord, glasse,
and the like.*

THe ordinary measure for bredth and length are feete and inches, each foote diuided into 12 inches, and euery inch into halues & quarters,



ters, which being parts of severall denominations, doth breed much trouble both in arithmetique and the use of instruments.

For the avoiding whereof, where I may preuaile I giue this counsell, that such as are delighted in measure would use severall lines, first a line of inch measure, wherein euery inch may be diuided into 10 or 100 parts; secondly a line of foote measure, wherein euery foote may be diuided into 100 or 1000 parts, both which lines may be set on the same side of a two foote ruler, after this or the like maner.



Then if they be to giue the content of any superficies or solid in inches, they may measure the sides of it by the line of inches and parts of inches; but if they be to giue the content in feete, it would be more easie for them to measure those sides by the foote line and his parts.

For example, let the length of a plane be 30 inches, and the bredth 21 inches and $\frac{6}{10}$ of an inch; this length multiplied into the bredth, would giue the content to be 648 inches: but if I were to find the content of the same plane in feet, I would measure the sides of it by the foote line and his parts; so the length would proue to be 2 feete $\frac{10}{100}$, and the bredth 1 foote $\frac{26}{100}$, and the length multiplied by the bredth, cutting off the foure last figures, for the foure figures of the parts, would giue the content to be 4.5000, which is 4 foote and 5000 parts, of a foot being diuided into 10000 parts.

$$\begin{array}{r}
 21.6 \\
 30.0 \\
 \hline
 648.00 \\
 \\
 \begin{array}{r}
 2.50 \\
 1.80 \\
 \hline
 20000 \\
 250 \\
 \hline
 4.5000
 \end{array}
 \end{array}$$

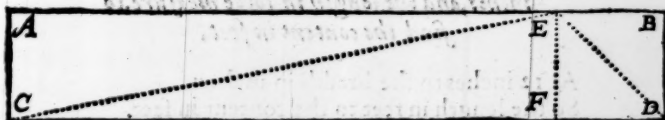
The like reason holdeth for yards and elnes, and all other measures diuided into 10, 100, or 1000 parts.

This being presupposed, the worke will be more easie both by arithmetique and the line of *numbers*, as may appeare by these propositions.

- 1 *Having the bredth and length of any oblong superficies given in inch-measure, to finde the content in inches.*

As 1 inch vnto the bredth in inches:

So the length in inches vnto the content in inches.



Suppose in the plane AD, the bredth AC to be 30 inches, and the length AB to be 183 inches; extend the compasses from 1 vnto 30, the same extent will reach from 183 vnto 5490; or extend them from 1 vnto 183, the same extent will reach from 30 vnto 5490. So both wayes the content required is found to be 5490 inches.

As 1 vnto 30: so are 183 vnto 5490.

- 2 *Having the length and bredth of any oblong superficies given in inches, to finde the content in fecte.*

As 144 inches vnto the bredth in inches;

So the length in inches vnto the content in fecte.

And thus in the former plane AD , working as before, the content will be found to be 38.125, which is 38 foote and $\frac{1}{8}$ of a foote.

As 144 vnto 30: so are 183 vnto 38.125.

- 3 *Having the length and breadth of any oblong superficies given in foote measure, to finde the content in feete.*

As 1 foote vnto the bredth in foote measure:

So the length in feete vnto the content in feete.

And thus in the former plane AD, the bredth will be 2 foote 50 parts; and the length 15 foote 25 parts; then working as before, the content will be found to be 38.125.

As 1 vnto 2.50: so are 15.25 vnto 38.125.

- 4 *Having the breadth of any oblong superficies given in inches, and the length in foote measure, to find the content in feete.*

As 12 inches to the bredth in inches:

So the length in feet to the content in feet.

So also in the former plane, the content will be found to be 38.125.

As 12 vnto 30: so are 15.25 vnto 38.125.

- 5 *Having the breadth of an oblong superficies given in inches, to finde the length of a foot superficiall in inch measure.*

As the bredth in inches, vnto 144 inches:

So 1 foot vnto the length in inch measure,

So the bredth being 30 inches, the length of a foot will be found to be 4 inches 80 parts.

As 30 vnto 144: so are 1 vnto 4.80.

- 6 *Having the breadth and length of an oblong superficies given in feet, to finde the length of a foot superficiall in foot measure.*

As the bredth in foot measure to 1 foot:

So the number of feet to the length in foot measure.

So the breadth being 2 foot 30 parts, the length of a foot will be found to be 40 parts, the length of 2 feet 80 parts, and the length of 3 feet 120 parts, &c.

As 250 vnto 1: so are 1 vnto 0.40

7 Having the length and breadth of an oblong superficies, to find the side of a square equall to the oblong.

Divide the space between the length and the breadth into two equall parts, and the foot of the compasses will stay at the side of the square.

So the length being 183 inches, and the breadth 30 inches, the side of the square will be found to be almost 74 inches and 10 parts of 100.

Or the breadth being 2 foot and 30 parts, the length 15 foot and 25 parts, the side of the square will be found to be about 6 feet and 17 parts.

As 30 vnto 74.10: so are 74.10 vnto 183.027.

And as 2.30 vnto 6.174: so are 6.174 vnto 15.247.

8 Having the diameter of a circle, to find the side of a square equall to that circle.

As 10000 to the diameter:

So 8862 vnto the side of the square.

So the diameter of a circle being 15 inches, the side of the square will be found about 13 inches and 29 parts.

As 10000 vnto 8862: so are 15 vnto 13.29.

9 Having the circumference of a circle to find the side of a square equall to the same circle.

As 10000 to the circumference:

So 2821 to the side of the square.

So the circumference of a circle being 47 inches 13 parts, the side of the square will be about 13 inches 29 parts.

As 10000 vnto 2821: so are 47.13 vnto 13.29.

10 *Having the diameter of a circle, to find the circumference.*

11 *Having the circumference of a circle, to find the diameter.*

As 1000 to the diameter:

So 3142 to the circumference.

So the diameter being 15 inches, the circumference will be found about 47 inches 13 parts: or the circumference being 47.13, the diameter will be 15.

CHAP. II.

The use of the line of Numbers in the measure of land by perches and acres.

1 *Having the breadth and length of an oblong superficies given in perches, to find the content in perches.*

As 1 perch to the breadth in perches:

So the length in perches to the content in perches.

So in the former plane *AD*, if the breadth *AC* be 30 perches, and the length *AB* 183 perches, the content will be found to be 5490 perches.

2 *Having the length and breadth of an oblong superficies given in perches, to find the content in acres.*

As 160 to the breadth in perches:

So the length in perches to the content in acres.

So in the former plane *AD*, the content will be found to be 34 acres, and 31 centesims or parts of an 100.

As 160 vnto 30: so are 183 vnto 34.31.

3 *Having the length and breadth of an oblong superficies given in chaines, to find the content in acres.*

It being troublesome to divide the content in perches by 160, we may measure the length and breadth by chaines, each chaine being 4 perches in length, and divided into 100 links, then will the worke be more easie in arithmetique.

For

breadth
As 10 to the ~~content~~ in chaines:

So the length in chaines to the content in acres.

And thus in the former plane AD, the breadth AC will be 7 chaines 50 links, and the length AB 45 chaines 75 links; then working as before, the content will be found as before, 34 acres 31 parts.

4 *Having the perpendicular and base of a triangle given in perches, to find the content in acres.*

If the perpendicular go for the breadth, and the base for the length, the triangle will be the halfe of the oblong. As the triangle CED is the halfe of the oblong AD, whose content was found in the former Prop. Or without halving,

As 320 to the perpendicular:

So the base to the content in acres.

So in the triangle CED, the perpendicular being 30, and the base 183, the content will be found to be about 17 acres and 15 parts.

5 *Having the perpendicular and base of a triangle given in chaines, to find the content in acres.*

As 20 to the perpendicular:

So the base to the content in acres.

And so in the triangle CED, the perpendicular EF being

ing 7.50, and the base CD 45.75, the content will be found as before to be about 17 acres 15 parts.

6 *Having the content of a superficies after one kind of perch, to find the content of the same superficies according to another kind of perch.*

As the length of the second perch
to the length of the first perch;

So the content in acres to a fourth number;
and that fourth to the content in acres required.

Suppose the plane AD measured with a chaine of 66 feet, or with a perch of 16 feete and an halfe, contained 34 acres 31 parts; and it were demanded how many acres it would containe if it were measured with a chaine of 18 foot to the perch: these kind of propositions are wrought by the backward rule of *three*, after a duplicated proportion. Wherefore I extend the compasses from 16.5 vnto 18.0, and the same extent doth reach backward, first from 34.31 to 31.45, and then from 31.45 to 28.84, which shewes the content to be 28 acres 84 parts.

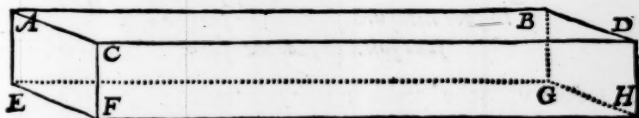
7 *Having the plot of a plane with the content in acres, to find the scale by which it was plotted.*

Suppose the plane AD contained 34 acres 31 centesmes; if I should measure it with a scale of 10 in the inch, the length AB would be 38 chaines and about 12 centesmes, and the breadth AC 6 chaines and 25 centesmes; and the content would be found by the third *Prop.* of this Chapter, to be about 23 acres 82 parts, whereas it should be 34 acres 31 parts.

Wherefore I diuide the distance betweene 23.82, and 34.31, vpon the line of *numbers* into two equall parts; then setting one foote of the compasses vpon 10, my supposed scale, I find the other to extend to 12, which is the scale required.

CHAP. III.

*The use of the line of Numbers in solid measure,
such as stone, timber, and the like.*



- 1 *Having the side of a square equall to the base of any solid given in inch measure, to find the length of a foot solid in inch measure.*

THe side of a square equall to the base of a solid, may be found by diuiding the space between the length and bredth into two equall parts, as in the 7 *Prop.* of broad measure. Then

As the side of the square in inches to 41.57:

So is 1 foot to a fourth number;
and that fourth to the length in inches.

So in the solid *AH*, the side of the square equall to the base *EC*, being about 25 inches 45 parts, the length of a foot solid will be found about 2 inches 67 parts, and the length of two foot solid 5 inches 33 parts.

As 25.45 vnto 41.57 : so 1.00 vnto 1.63:

and so are 1.63 vnto 2.67.

- 2 *Having the side of a square equall to the base of any solid given in foot measure, to find the length of a foot solid in foot measure.*

As the side of the square in feet vnto 11

So is 1 vnto a fourth number;

And that fourth to the length in foot measure:

So in the solid *AH*, the side of the square equall to the
base

base EE , being about 2 foote 120 parts, the length of a foot solid will be found about 222 parts of a foot.

As 2.120 vnto 1.000: so 1.000 vnto 0.471:
and so are 471 vnto 222.

3 *Having the bredth and depth of a squared solid given in foot measure, to find the length of a foot solid in foot measure.*

As 1 vnto the bredth in foot measure:
So the depth in feet to a fourth number;
which is the content of the base in foot measure. Then
As this fourth number vnto 1:
So 1 vnto the length in foot measure.

So in the solid AH , the bredth being 2 foot 50 parts, the depth 1 foot 80 parts, the content of the base EC will be found 4 foot 50 parts, and the length of one foot solid about 222 parts, the length of two foot solid about 444 parts of 1000.

As 1.00 vnto 2.50: so are 1.80 vnto 4.50.
As 4.50 vnto 1.00: so 1.000 vnto 0.222.

4 *Having the bredth and depth of a squared solid given in inches, to find the length of a foot solid in inch measure.*

As 1 hath to the bredth in inches:
So the depth in inches to a fourth number;
which is the content of the base in inches. Then
As this fourth number vnto 1728:
So 1 vnto the length of a foot in inch measure.

So in the solid AH , the bredth AC being 30 inches, and the depth AE 21 inches 60 parts, the content of the base EC will be found to be 648 inches, and the length of a foot solid about 2 inches 67 parts.

As 1 vnto 21.6: so 30 vnto 648:

As 648 vnto 1728: so 1 vnto 2.667.

Or as 12 to the bredth in inches:

So the depth in inches to a fourth number.

As this fourth number to 144:

So 1 vnto the length of a foot solid in inch measure.

So in the solid *AH*, the bredth being 30 inches, the depth 21 inches 6 parts, the fourth number will be found to be 54, and the depth of a foot solid 2 inches 67 parts.

As 12 vnto 21.6: so 30 vnto 54.

As 54 vnto 144: so 1 vnto 2.667.

- 5 *Having the side of a square equall to the base of any solid, and the length thereof given in inch measure, to find the content thereof in feet.*

As 41.57 to the side of the square in inches:

So the length in inches to a fourth number;

and that fourth to the content in foot measure.

So in the solid *AH*, the length *AB* being 183 inches, and the side of the square equall to the base *EC* about 25 inches 45 parts, the fourth number will be found about 112, and the whole solid content about 68 feet 62 parts.

As 41.57 vnto 25.45: so 183 vnto 112:

and so are 112 vnto 68.62.

- 6 *Having the side of a square equall to the base of any solid, and the length thereof given in foot measure, to find the content thereof in feet.*

As 1 to the side of the square in foot measure:

So the length in feet to a fourth number;

and that fourth to the content in foot measure.

So in the former solid *AH*, the side of the square equall to the base *AE*, being about 2 foot 12 parts, and the length *AB* 15 foot 25 parts, the content will be found to be about 68 foot 62 parts.

As 1 vnto 2.12: so 15.25 vnto 32.35:
and so are 32.35 vnto 68.62.

- 7 *Having the side of a square equall to the base of any solid given in inch measure, & the length of the solid given in foot measure, to find the content thereof in feet.*

As 12 to the side of the square given in inches:
So the length in feet to a fourth number;
and that fourth to the content in foot measure.

So in the former solid *AH*, the side of the equall square being 25 inches 45 parts, the content will be found to be about 68 feet 62 parts.

As 12 vnto 25.45: so 15.25 vnto 32.35:
and so are 32.35 vnto 68.62.

- 8 *Having the length, bredth and depth of a squared solid given in inches, to find the content in inches.*

As 1 vnto the bredth in inches:
So the depth in inches vnto the base in inches. Then
As 1 vnto the base:
So the length in inches vnto the solid content in inches.

So in the solid *AH*, whose bredth *AC* is 30 inches, the depth *AE* 21 inches and 6 parts of 10, and length *AB* 183, the content of the base *EC* will be found 648 inches, and the whole solid content about 118500 inches.

As 1 vnto 21.6: so are 30 vnto 648:
As 1 vnto 648: so are 183 to 118584.

9 *Having the length, bredth and depth of a squared solid giuen in inches, to find the content in feet.*

As 1 to the bredth in inches:

So the depth in inches to the base in inches.

As 1728 to that base:

So the length in inches to the content in feet.

So in the solid *AH*, the content will be found to be about 68 feet 62 parts.

As 1 vnto 21.6: so 30 vnto 648:

As 1728 vnto 648: so 183 to 68.62.

Or as 12 to the bredth in inches:

So the depth in inches to a fourth number.

As 144 to that fourth number:

So the length in inches to the content in feet.

And so also in the same solid *AH*, the content will be found to be about 68 feet 62 parts.

As 12 vnto 21.6: so 30 vnto 54:

As 144 vnto 54: so 183 vnto 68.62.

10 *Having the length, bredth and depth of a squared solid giuen in foot measure, to finde the content in feet.*

As 1 vnto the bredth in foot measure:

So the depth in feet to the base in feet.

As 1 vnto that base:

So the length in feet to the content in feet.

And thus in the former solid *AH*, the bredth *AC* will be 2 foot 50 parts, the depth *AE* 1 foot 80 parts, and the length *AB* 15 foot 25 parts; then working as before, the content of the base *AF* will be found 4 feet 50 parts, and the whole solid content about 68 foot 62 parts, which of all others may

very easily be tried by arithmetique.

As 1 vnto 2.50: so 1.80 vnto 4.50.

As 1 vnto 4.50: so 15.25 vnto 68.625.

II *Having the bredth and depth of a squared solid given in inches, and the length in foot measure, to find the content thereof in feet.*

As 1 vnto the bredth in inches:

So the depth in inches vnto a fourth number:
which is the content of the base in inches.

As 144 hath vnto that fourth number:

So the length in feet to the content in feet.

And so in the same solid AH, the content will be found to be about 68 feet 62 parts.

As 1 vnto 21.6: so 30 vnto 648.

As 144 vnto 15.25: so 648 vnto 68.62.

Or as 144 vnto the bredth in inches:

So the depth in inches vnto a fourth number:
which is the content of the base in feet.

As 1 hath vnto that fourth number:

So the length in feet to the content in feet.

And so in the same solid AH, the content will be found to be about 68 feet 62 parts.

As 144 vnto 21.6: so 30 vnto 4.50:

As 1 vnto 4.50: so 15.25 vnto 68.62.

Or as 12 vnto the bredth in inches:

So the depth in inches vnto a fourth number.

As 12 vnto this fourth number:

So the length in feet to the content in feet.

And so also in the same solid AH, the content will be found to be about 68 feet 62 parts.

As 12 vnto 21.6: so 30 vnto 54.

As 12 vnto 54: so 15.25 vnto 68.62.

All these varieties (and such like not here mentioned)

do

do follow vpon making of the bafe of the folid, to be *EC*; there would be as many more if any fhall begin with the bafe *EH*, and fo likewise if they make the bafe to be *FD*.

12 *Having the diameter of a cylinder given in inch meafure, to find the length of a foot folid in inches.*

As the diameter in inches vnto 46.90:
So is 1 vnto a fourth number:
and that fourth to the length in inches.

So the diameter of a cylinder being 15 inches, the fourth number will be about 3.12, and the length of a foote folid 9 inches 78 parts.

As 15 vnto 46.90: fo 1 vnto 3.127:
and fo are 3.127 vnto 9.778.

13 *Having the diameter of a cylinder given in foot meafure, to find the length of a foot folid in foot meafure.*

As the diameter in feet vnto 1.128:
So is 1 vnto a fourth number;
and that fourth to the length in foot meafure.

So the diameter being 1 foot 25 parts, the length of a foot folid will be found about 8.14 parts of 1000.

As 1.25 vnto 1.128: fo 1.00 to 0.9027:
and fo are 9027 vnto 8148.

14 *Having the circumference of a cylinder given in inches, to find the length of a foot folid in inch meafure.*

As the circumference in inches to 147.36:
So is 1 to a fourth number;
and that fourth to the length in inches.

So the circumference being 47 inches 13 parts, the length of a foot folid will be found about 9 inches 78 parts.



As 47.13 vnto 147.36: so 1.00 to 3.13:
and so are 3.13 vnto 9.78.

- 15 *Having the circumference of a cylinder giuen in foot measure, to find the length of a foot solid in foot measure.*

As the circumference in fecte to 3.545:
So is 1 to a fourth number;
and that fourth to the length in foot measure.

So the circumference being 3 foot 927 parts, the length of a foot solid will be found to be about 815 parts.

As 3.927 vnto 3.545: so 1.000 vnto 0.903:
and so are 903 vnto 815.

- 16 *Having the side of a square equall to the base of a cylinder, to find the length of a foot solid.*

The side of a square equall to the circle, may be found by the eighth *Prop.* of broad measure, and then this *Prop.* may be wrought by the first and second *Prop.* of solid measure.

- 17 *Having the diameter of a cylinder, and the length giuen in inches, to find the content in inches.*

As 1.128 vnto the diameter in inches:
So the length in inches to a fourth number;
and that fourth number to the content in inches.

So the diameter being 15 inches, and the length 105, the content of the cylinder will be found to be about 18560 inches.

As 1.1284 vnto 15: so are 105 vnto 1395.87:
and so are 1395.87 vnto 18555.34.

- 18** *Having the diameter and length of a cylinder in foot measure, to find the content in feet.*

As 1.128 to the diameter in feet:
So the length in feet to a fourth number;
and that fourth to the content in feet.

So the diameter being 1 foote 25 parts, and the length 8 foot and 75 parts, the content of the cylinder will be found about 10 foot 74 parts.

As 1.128 vnto 1.25: so 8.75 vnto 9.69:
and so are 9.69 vnto 10.737.

- 19** *Having the diameter of a cylinder, and the length given in inches, to find the content in feet.*

As 46.90 to the diameter in inches:
So the length in inches to a fourth number;
and that fourth to the content in feet.

So the diameter being 15 inches, and the length 105, the content will be found about 10 foot 74 parts.

As 46.906 vnto 15: so 105 vnto 33.58:
and so are 33.58 vnto 10.737.

- 20** *Having the diameter of a cylinder given in inches and the length in feet, to find the content in feet.*

As 13.54 to the diameter in inches:
So the length in feet to a fourth number;
and that fourth to the content in feet.

So the diameter being 15 inches, and the length 8 foote 75 parts, the content will be found about 10 foot 74 parts.

As 13.54 vnto 15: so 8.75 vnto 9.69:
and so are 9.69 vnto 10.74.

- 21 *Having the circumference and the length of a cylinder given in inches, so find the content in inches.*

As 3.545 to the circumference in inches:
So the length in inches to a fourth number;
and that fourth to the content in inches.

So the circumference being 47 inches 13 parts, and the length 105 inches, the content will be found about 18560 inches.

As 3.545 vnto 47.13: so 105 vnto 1396:
and so are 1396 vnto 18555.

- 22 *Having the circumference and length of a cylinder given in inches, so find the content in feet.*

As 147.36 to the circumference in inches:
So the length in inches to a fourth number;
and that fourth to the content in feet.

So the circumference being 47 inches 13 parts, and the length 105 inches, the content will be found about 10 foote 74 parts.

As 147.36 vnto 47.13: so 105 vnto 33.58:
and so are 33.58 vnto 10.74.

- 23 *Having the circumference and length of a cylinder given in foot measure, so find the content in feet.*

As 3.545 to the circumference in feet:
So the length in feet to a fourth number;
and that fourth to the content in feet.

So the circumference being 3 foote 927 parts, and the length 8 foot 75 parts, the content will be found to be 10 foot 74 parts.

As 3.545 vnto 3.927: so 8.75 vnto 9.69:
and so are 9.69 vnto 10.74.

- 24 *Having the circumference of a cylinder given in inches, and the length in foot measure, to find the content in feets.*

As 42.54 to the circumference in inches:
So the length in feet to a fourth number;
and that fourth to the content in feet.

So the circumference being 47 inches 13 parts, and the length 8 foot 75 parts, the content will be found as before, 10 foot 74 parts.

As 42.54 vnto 47.13: so 8.75 vnto 9.69:
and so are 9.69 vnto 10.74.

CHAP. IIII.

The vse of the line of Numbers in gauging of vessels.

THe vessels which are here measured, are supposed to be cylinders, or reduced vnto cylinders, by taking the mean between the diameter at the head and the diameter at the bungue, after the vsuall maner.

- 1 *Having the diameter and the length of a vessell with the content thereof, to find the gauge point.*

Extend the compasses in the line of *numbers* to halfe the distance between the content and the length of the vessell, the same extent will reach from the diameter to the gauge point.

I put this proposition first, because these kind of measures are not alike in all places. Here at London it is said that a wine vessell being 66 inches in length, and 38 inches the diameter, would containe 324 gallons: which if it be true, we

31 *The use of the line of Numbers in gauging.*

may diuide the space betweene 324 and 66 into two equall parts, and the middle will fall about 146, and the same extent which reacheth from 324 to 146, wil reach from the diameter 38 vnto 17.15 the gauge point for a gallon of wine or oyle after London measure. The like reason holdeth for the like measures in all other places.

2 *Having the meane diameter and the length of a vessell, to find the content.*

Extend the compasses from the gauge point to the meane diameter, the same extent being doubled, shall giue the distance from the length to the content.

So the meane diameter of a wine vessell being 20 inches, and the length 25 inches, the content will be found to be 34 gallons after London measure. For extend the compasses from 17.15 vnto 20, the same extent wil reach from 25 vnto 29.15, and from 29.15 vnto 34.

In like maner if the meane diameter were 16 inches, and the length 23, the content would be found to be about 20 gallons. For the same extent which reacheth back from 17.15 vnto 16, will reach from 23 to 21.45, and from 21.45 vnto 20.

So that if the meane diameter shall be 17 inches and 15 centesmes or parts of 100, the number of inches in the length of the vessell, will giue the number of gallons contained in the same vessell: if the diameter shall be more or lesse then 17.15, the content in gallons will be accordingly more or lesse then the length in inches.

3 *Having the diameter and content, to find the length.*

Extend the compasses from the diameter to the gauge point, the same extent being doubled shall giue the distance from the content to the length of the vessell.

So the gauge point standing as before, if the diameter shall be 38 inches, and the content 324 gallons wine measure, the
length

length of the vessels will be found about 66 inches.

- 4 *Having the length of a vessell and the content,
to find the diameter.*

Extend the compasses to halfe the distance betweene the length and the content, the same extent shall reach from the gauge point to the diameter.

So the length being 66 inches, and the content 324 gallons wine measure, the gauge point standing as before, the diameter of the vessell will be found to be about 38 inches.

CHAP. V.

*Containing such Astronomicall propositions
as are of ordinary use in the pra-
ctise of Navigation.*

- 1 *To find the altitude of the Sunne by the shadow
of a gnomon set perpendicular
to the horizon.*

As the parts of the shadow
are to the parts of the gnomon:

So the tangent of 45 gr.
to the tangent of the altitude.

Extend the compasses in the line of *numbers*, from the parts of the shadow to the parts of the *gnomon*; the same extent will giue the distance from the tangent of 45 gr. to the tangent of the Sunnes altitude.

So the *gnomon* being 36, and the shadow 27, the altitude will be found to be 36 gr. 52 m. Or the *gnomon* being 27, and the shadow 36, the altitude will be found to be 53 gr. 8 m. Or the shadow being 20, and the *gnomon* 9, the altitude will be found to be 24 gr. 14 m. as in the eighth *Prop.* of the use of the *tangent* line.

2 *Having the distance of the Sunne, from the next equinoctiall point, to find his declination.*

As the Radius is in proportion
to the sine of the Sunnes greatest declination:
So the sine of the Suns distance from the next equinoctiall point,
to the sine of the declination required.

Extend the compasses in the line of *sines*, from 90 gr. to 23 gr. 30 m. the same extent will giue the distance from the Sunnes place vnto his declination.

So the Sunne being either in 29 gr. of φ , or 1 gr. of ω , or 1 gr. of Ω , or 29 gr. of \mathfrak{m} , that is 59 gr. distant from the next equinoctiall point, the declination will be found about 20 gr.

If the Sunne be so neare the equinoctiall point that his declination fall to be vnder 1 gr. it may be found by the line of *numbers*. As if the Sunne were in 2 gr. 5 m. of φ , that is, 125 m. from the equinoctiall point, the former extent of the compasses from the sine of 90 gr. to the sine of 23 gr. 30 m. will reach in the line of *numbers* from 125 vnto 50, which shewes the declination to be about 50 m.

3 *Having the latitude of the place, and the declination of the Sun, to find the time of the Suns rising and setting.*

As the cotangent of the latitude
to the tangent of the Suns declination:
So is the Radius

to the sine of the ascensionall difference betweene the
houre of 6 and the time of the Suns rising or setting.

Extend the compasses from the tangent of the complement of the latitude, to the tangent of the declination: the same extent will reach from the sine of 90 gr. to the sine of the ascensionall difference.

Or extend the compasses from the cotangent of the latitude

to the sine of 90 gr, the same extent will reach from the tangent of the declination to the sine of the ascensionall difference.

So the latitude being 51 gr. 30 m. Northward, and the declination 20 gr. the difference of ascension will be found to be 27 gr. 14 m. which resolved into houres and minures, doth giue 1 houre and almost 49 m. for the difference betweene the Sunnes rising or setting, and the houre of 6, according to the time of the yeare.

4 *Having the latitude of the place, and the distance of the Sun from the next equinoctiall point, to find his amplitude.*

As the cosine of the latitude
to the sine of the Suns greatest declination:
So the sine of the place of the Sun,
to the sine of the amplitude,

So the latitude being 51 gr. 30 m. and the place of the Sun in 1 gr. of π , that is 59 gr. distant from the next equinoctiall point, the amplitude will be found about 33 gr. 20 m. For extend the compasses in the line of sines, from 38 gr. 30 m. the sine of the complement of the latitude, vnto 23 gr. 30 m. the sine of the Suns greatest declination; the same extent will reach from 59 gr. vnto 33 gr. 20 m. Or extend them from 38 gr. 30 m. vnto 59 gr. the same extent will reach from 23 gr. 30 m. vnto 33 gr. 20 m. as before.

5 *Having the latitude of the place, and the declination of the Sun, to find his amplitude.*

As the cosine of the latitude
is to the Radius:
So the sine of the declination,
to the sine of the amplitude.

Extend the compasses from the cosine of the latitude to

54 *The use of the lines of sines and tangents*

the sine of 90 gr. the same extent will reach from the sine of the Suns declination to the sine of the amplitude.

Or extend them from the cosine of the latitude to the sine of the declination, the same extent will reach from the sine of 90 gr. to the sine of the amplitude.

So the latitude being 51 gr. 30 m. and the declination 20 gr. the amplitude will be found to be 33 gr. 20 m.

6 *Having the latitude of the place, and the declination of the Sun, to find the time when the Sun cometh to be due East or west.*

As the tangent of the latitude,
is to the tangent of the declination:

So the Radius
to the cosine of the houre from the meridian.

Extend the compasses from the tangent of the latitude to the tangent of the declination; the same extent will reach from the sine of 90 gr. to the sine of the complement of the houre.

Or extend them from the tangent of the latitude to the sine of 90 gr; the same extent will reach from the tangent of the declination to the sine of the complement of the houre.

So the latitude being 51 gr. 30 m. and the declination 20 gr. the Sunne will be 73 gr. 10 m: that is 4 houres and 53 m. frō the meridian, when he cometh to be in the East or West.

7 *Having the latitude of the place, and the declination of the Sun, to find what altitude the Sun shall have, when he cometh to be due East or west.*

As the sine of the latitude
is to the sine of the declination:

So the Radius
to the sine of the altitude.

Extend the compasses in the line of *Sines* from the latitude

to the sine of the declination, the same extent will reach from the sine of 90 gr. to the sine of the altitude.

Or extend them from the sine of the latitude to the sine of 90 gr; the same extent will reach from the sine of the declination to the sine of the altitude.

So the latitude being 51 gr. 30 m. and the declination 20 gr. the altitude will be found about 25 gr. 55 m.

8 *Having the latitude of the place, and the declination of the Sun, to find what altitude the Sun shall have at the hour of six.*

As the Radius is in proportion
to the sine of the ~~sine of the~~ declination:

So the sine of the latitude
to the sine of the altitude.

Extend the compasses in the line of *sines*, from 90 gr. to the declination; the same extent will reach from the latitude to the altitude.

Or extend them from 90 gr. to the latitude, the same extent will hold from the declination to the altitude.

So the latitude being 51 gr. 30 m. and the declination of the Sunne 20 gr. the altitude of the Sun will be found to be about 15 gr. 30 m.

9 *Having the latitude of the place, and the declination of the Sun, to find what azimuth the Sun shall have at the hour of six.*

As the cosine of the latitude
is to the Radius:

So the cotangent of the Suns declination,
to the tangent of the azimuth from the North part
of the meridian.

So the latitude being 51 gr. 30 m. and the declination 20 gr. the azimuth will be found to be 77 gr. 14 m. For extend the compasses in the line of *sines*, from 38 gr. 30 m. to 90 gr. the same extent will reach from the tangent of 70 gr. to the tangent of 77 gr. 14 m.

10 *Having the latitude of the place, and the declination of the Sun, and the altitude of the Sun, to find the azimuth.*

First consider the declination of the Sunne, whether it be toward the North or the South, so have you his distance from your pole: then adde this distance, the complement of his altitude, and the complement of your latitude, all three together, and from halfe the summe subtraet the distance from the pole, and note the difference.

- 1 As the Radius is in proportion
to the cosine of the altitude:
So the cosine of the latitude,
to a fourth sine.
- 2 As this fourth sine
is to the sine of the halfe summe:
So the sine of the difference,
to a seuenth sine.

Then find a meane proportionall betweene this seuenth sine and the Radius, this meane shall be the sine of the complement of halfe the azimuth from the North part of the meridian.

(Suppose the declination of the Sun being knowne by the time of the yeare to be 20 gr. Southward, the altitude above the horizon found by obseruation 12 gr. and the latitude Northwards 51 gr. 30 m. it were required to find the azimuth. The declination is Southward, and therefore the distance from the pole 110 gr; then turning the altitude and latitude vnto their complements, I adde them all three together, and from halfe the summe subtraet the distance from the pole,
noting

noting the difference after this maner.

Declin. South	20 gr. 0 m.	The distance	110 gr. 0 m.
Altitude	12 0	The complement	78 0
Latitude N.	51 30	The complement	38 30
The summe of all three			226 30
The halfe summe			113 15
The difference			3 19

This done, I come to the Staffe, and extend the compasses from the line of 90 gr. to the line of 78 gr. and find the same extent to reach from the line of 38 gr. 30 m. vnto 37 gr. 30 m. Or if I extend them from 90 gr. to 38 gr. 30 m. the same extent doth reach from 78 gr. vnto 37 gr. 30 m. which is the fourth line required.

Then I extend the compasses againe, from this fourth line of 37 gr. 30 m. vnto the line of the halfe summe 113 gr. 15 m. that is to the line of 66 gr. 45 m. (for after 90 gr. the line of 30 gr. doth stand for a line of 100 gr. and the line of 70 gr. for a line of 110 gr. and so the rest for those which are their complements to 180 gr.) and this second extent doth reach from the line of the difference 3 gr. 15 m. to the line of 4 gr. 54 m. Or if I extend them from the fourth line of 37 gr. 30 m. to the line of the difference 3 gr. 15 m. the same extent will reach from the line of the halfe summe 113 gr. 15 m. vnto 4 gr. 54 m. which is the seuenth line required.

Lastly, I diuide the space betweene this seuenth line of 4 gr. 54 m. and the line of 90 gr. into two equall parts, and I find the meane proportionall line to fall on 17 gr. whose complement is 73 gr; the double of 73 gr. is 146 gr. and such is the azimuth required.

Or hauing found the seuenth line to be 4 gr. 54 m. I might looke ouer against it, in the line of *versed sines*, and there I should find 146 gr. for the azimuth from the North part of the meridian; and the complement of 146 gr. to a semicircle being 34 gr. will giue the azimuth from the South part of the meridian.

But if it were required to find the azimuth in the same latitude of 51 gr. 30 m. Northward, with the same altitude of

12 gr. and like declination of 20 gr. to the Northward, it would be found to be onely 72 gr. 52 m. though the manner of worke be the same as before.

Declin. North	20 gr. 0 m	The distance is	70 gr. 0 m.
Altitude	12 0	The complement	78 0
Latitud. North	51 30	The complement	38 30
The summe of all three			186 30
The halfe summe			93 15
The difference			23 15

Here as the Radius is to the sine of 78 gr. so the sine of 38 gr. 30 m. to the sine of 37 gr. 30 m. which is the fourth sine, and the same as before.

Then as this fourth sine of 37 gr. 30 m. is to the sine of 93 gr. 15 m. so the sine of 23 gr. 15 m. to the sine of 40 gr. 20 m. which is the seventh sine.

The halfe way betweene this seventh sine and the sine of 90 gr. doth fall at 53 gr. 34 m. whose complement is 36 gr. 26 m; and the double of that is 72 gr. 52 m. the azimuth required.

Or I may find this same azimuth in the line of *versed sines*, ouer against the seventh sine of 40 gr. 20 m.

II Having the latitude of the place, the declination of the Sun, and the altitude of the Sun, to find the houre of the day.

Adde the complement of the Suns altitude, and the distance of the Sun from the pole, and the complement of your latitude, all three together, and from halfe the summe subtract the complement of the altitude, and note the difference.

- 1 As the Radius is in proportion
to the sine of the Suns distance from the pole:
So the sine of the complement of the latitude,
to a fourth sine,

2 As this fourth line
is to the sine of the halfe summe:

So the sine of the difference
to a seuenth line.

The meane proportionall betweene this seuenth line and
the sine of 90 gr. will be the sine of the complement of halfe
the houre from the meridian.

Thus in our latitude of 51 gr. 30 m. the declination of the
Sun being 20 gr. Northward, and the altitude 12 gr. I might
find the Sun to be 95 gr. 52 m. from the meridian.

Altitude	12 gr. 0 m.	The complement is	78 gr. 0 m.
Declin. North	20 0	the dist. from the pole	70 0
Latitude	51 30	the complement is	38 30
The summe of all three			186 30
The halfe summe			93 15
The difference			15 15

Here as the Radius is to the sine of 70 gr.

So the sine of 38 gr. 30 m. to the sine of 35 gr. 48 m.

As this sine of 35 gr. 48 m. is to the sine of 93 gr. 15 m.

So the sine of 15 gr. 15 m. to the sine of 26 gr. 40 m.

The halfe way between this seuenth sine of 26 gr. 40 m. and
the sine of 90 gr. doth fall at 42 gr. 4 m. whose complement is
47 gr. 56 m. and the double of that, 95 gr. 52 m. which con-
uerted into houres, doth giue 6 houres and almost 24 m. from
the meridian.

Or I might find these 95 gr. 52 m. in the line of *versed sines*,
ouer against the seuenth sine of 26 gr. 40 m.

12 Having the azimuth, the Suns altitude, and the
declination, to find the houre of the day.

As the cosine of the declination
is to the sine of the azimuth:

So the cosine of the altitude
to the sine of the houre.

h 2

Thus

Thus the declination being 20 gr. Southward, the altitude 12 gr. and the azimuth found by the tenth *Prop.* 146 gr. I might find the time to be 35 gr. 36 m. that is 2 houres 31 m. from the meridian.

13 *Having the houre of the day, the Sunnes altitude, and the declination, to find the azimuth.*

As the cosine of the altitude
is to the sine of the houre:
So the cosine of the declination,
to the sine of the azimuth.

So the altitude of the Sun being 12 gr. and the declination 20 gr. Southward, and the angle of the houre 35 gr. 36 m. I should find the azimuth to be 34 gr. And so it is if it be reckoned from the South; but 146 gr. if it be taken from the North part of the meridian.

14 *Having the distance of the Sun from the next equinoctiall point, to find his right ascension.*

As the Radius
to the cosine of the greatest declination:
So the tangent of the distance,
to the tangent of the right ascension.

So the Sun being in the first degree of π , that is 59 gr. distant from the next equinoctiall point, and the greatest declination 23 gr. 30 m. the right ascension will be found to be 56 gr. 46 m. short of the beginning of γ ; and therefore 303 gr. 14 m.

15 *Having the declination of the Sun, to find his right ascension.*

As the tangent of the greatest declination
is to the tangent of the declination giuen

So the Radius

to the sine of the right ascension.

So the greatest declination being $23\text{ gr. }30\text{ m.}$ and the declination of the Sun giuen 20 gr. the right ascension will be found about $56\text{ gr. }50\text{ m.}$

These are such Astronomically propositions as I take to be usefull for Sea-men. For the first and second will help them to find their latitude; the third to find the Suns rising and setting; the 4. 5. 6. 7. 8. 9. 10. 11. 12. *Prop.* to finde the variation of their compasse; the 11 and 12 *Prop.* to find the houre of the day; and the two last toward the finding of the houre of the night. For hauing the latitude of the place, with the declination and altitude of any starre, they may find the houre of the starre from the meridian, as in the 11 *Prop.* Then comparing the right ascension of the starre with the right ascension of the Sunne, they may haue the houre of the night.

All these propositions and such others may be wrought also by the tables of *sines* and *tangents*. For where foure numbers do hold in proportion; as the first to the second, so the third to the fourth; there if we multiply the second into the third, and diuide the product by the first, the quotient will giue the fourth required. As in the example of the last *Prop.* where the declination being giuen, it was required to find the right ascension. The tangent of 20 gr. the declination giuen is 3639702, which being multiplied by the Radius, the product is 36397020000000, and this diuided by 4348124 the tangent of $23\text{ gr. }30\text{ m.}$ the quotient is 8370741 the sine of $56\text{ gr. }50\text{ m.}$ for the right ascension required.

Or if any will vse my tables of *artificiall sines* and *tangents*, they may adde the second and the third together, and from the summe subtract the first, the remainder will giue the fourth required. And so my tangent of 20 gr. is 9561.0658, which being added to the Radius, makes 19561.0658; from this if they subtract 9638.3019 the tangent of $23\text{ gr. }30\text{ m.}$ they shall find the remainder to be 9922.7639, which in my

Canon is the sine of 56 gr. 49 m. 56 seconds; and such is the right ascension required, if it be reckoned from the next equinoctiall point.

The like reason holdeth for all other Astronomicall propositions, as I will farther shew by those two examples which I gave before for the finding of the azimuth in the 10 *Prop.* because they are thought to be harder then the rest, and require three operations.

In the first example.

Declin. South	20 gr. 0 m.	The distance	110 gr. 0 m.
Altitude	12 0	the complement	78 0
Latitude Nor.	51 30	the complement	38 30
The summe of all three			226 30
The halfe summe			113 15
The difference			3 15

The first operation will be to finde the fourth line; and that is done by adding the sine of the complement of the altitude to the sine of the complement of the latitude, and subtracting the Radius: so adding 9990.4044 the sine of 78 gr. vnto 9794.1495 the sine of 38 gr. 30 m. the summe will be 19784.5539. And the Radius being subtracted, the remainder 9784.5539 is the fourth sine, and belongeth to 37 gr. 30 m.

The second operation will be to find the seventh sine; and that is done by adding the sine of the halfe summe to the sine of the difference, and subtracting the fourth sine. So the halfe summe being 113 gr. 15 m. I take his complement to a semicircle, and so find his sine to be 9963.2168, to which I adde 8753.5278, the sine of the difference 3 gr. 15 m; and the summe is 18716.7446. From this I take the fourth sine 9784.5539, and the remainder will be 8932.1907, which is the seventh sine, and belongeth to 4 gr. 54 m.

The third operation will be to finde the meane proportionall sine betwene the seventh sine and the Radius. This in common arithmetique is done by multiplying the two extremes, and taking the square roote of the product. As in finding

finding a meane proportionall betweene 4 and 9, we multiply 4 into 9, and the product is 36, whose square root is 6; the meane proportionall between 4 and 9. But here is done by adding the sine and the Radius, and taking the halfe of them. So the summe of the last seuenth sine and the Radius is 18932.1907 and the halfe of that 9466.0953, which is the meane proportionall sine required, and belongeth to 17 gr. whose complement is 73 gr. and the double of that 146 gr. the same azimuth as before.

In the second example.

Declin. North	20 gr. 0 m.	The distance	70 gr. 0 m.
Altitude	12 gr. 0 m.	the complement	78 gr. 10 m.
Latitud. North	51 gr. 30 m.	the complement	38 gr. 30 m.
The summe of all three			186 gr. 30 m.
The halfe summe			93 gr. 15 m.
The difference			23 gr. 15 m.

The first operation will be to finde the fourth sine; and that is here 9784.5539, as in the former example.

The second operation will be to find the seuenth sine; and so here the sine of the halfe summe 93 gr. 15 m. being the same with the sine of 86 gr. 45 m. his complement to 180 gr. I find it to be 9999.3009, to which I adde 9596.3153 the sine of the difference 23 gr. 15 m. and the summe is 19595.6162. From this I take the fourth sine 9784.5539, and the remainder will be 9811.0623 for the seuenth sine, and belongeth to 40 gr. 20 m.

The third operation will be to find the meane proportionall sine betweene the seuenth sine and the Radius. And so here the Radius being added to the seuenth sine, the summe will be 19811.0623, and the halfe of that 9905.5311 doth giue the meane proportionall sine belonging to about 53 gr. 34 m. whose complement is 36 gr. 26 m. & the double of that 72 gr. 52 m. the same azimuth as before.

I haue set downe these three examples thus particularly, that I might shew the agreement between the *Staffe* and the *Canon*. But otherwise I might deliver both the precept and the

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the worke, for the two last, more compendiously. For generally in all sphericall triangles, where three sides are knowne, and an angle required, make that side which is opposit to the angle required, to be the base; and gather the summe, the halfe summe, and the difference as before.

As the rectangle contained vnder the sines of the sides, is to the square of the whole sine:

So the rectangle contained vnder the sines of the halfe summe and the difference, to the square of the cosine of halfe the angle.

Then for the worke, we may for the most part leaue out the two last figures; and if they be about 50, put an vnitie to the sixth place, after this maner.

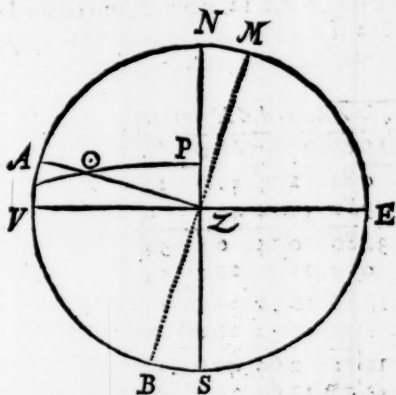
The second example.

70 gr. 0 m			
78	0	9990	40
38	30	9794	15
186	30	19784	55
93	15	9999	30
23	15	9596	32
		20000	00
		39595	62
		19811	07
36	26	9905	53
72	52		
			53 gr. 34 m.
		107	8

Or for such numbers as are to be subtracted, I may take them out of the Radius, and write downe the residue, and then adde them together with the rest. As in the same second example, the sines of 78 gr. and of 38 gr. 30 m. being the numbers to be subtracted; if I take 9990.4044 the sine of 78 gr. out of the Radius 10000 0000, the residue is 9.5956: and so the residue of 9794.1495 is 205.8505. Wherefore in stead of subtracting those sines, I may adde these residues after this maner:

70 gr. 0 m.		
78 0	9 59	
38 30	205 85	
186 30		
93 15	9999 30	
23 15	9596 32	
	19811 06	
36 26	9905 53	53 gr. 34 m.
72 52	107 8	

Having these meanes to find the Sunnes azimuth, we may compare it with the magneticall azimuth, and so finde the variation of the needle.



For let the circle AMB , drawne on the center Z , be a plane, parallel to the horizon; A the point whereon the Sun beareth from vs, M the North point of the magneticall needle, and the angle AZM the magneticall azimuth. If we find the Sunnes azimuth as before, to be $72\text{ gr. } 52\text{ m.}$ from the North to the Westward, we may allow so many degrees from A vnto N , and so we haue the true North point of the meridian, and consequently the East, South, and West points of the horizon; and the distance betweene N and M shall be the

the variation of the needle. So that if the magneticall azimuth AZM shall be $84^{\circ} 7'$ and the Suns azimuth AZN $72^{\circ} 52'$. then must NZM the difference betwene the two meridians, giue the variation to be $11^{\circ} 15'$ as Mr. *Bourough* heretofore found it by his obseruations at *Limehouse* in the year 1580. But if the magneticall azimuth AZM shall be $79^{\circ} 7'$. and the Suns azimuth AZN $72^{\circ} 52'$. then shall the variation NZM be only $6^{\circ} 15'$. as I haue sometimes found it of late. Herevpon I enquired after the place where Mr. *Bourough* obserued, and went to *Limehouse* with some of my friends, and tooke with vs a quadrant of 3 foote semidiameter, and two needles, the one about 6 inches, and the other 10 inches long, where I made the semidiameter of my horizontall plane AZ 12 inches; and toward night the 13 of Iune 1622, I made obseruation in seuerall parts of the ground, and found as followeth.

Alt.	◉	AZM	AZN	Varia
Gr. M.	Gr. M.	Gr. M.	Gr. M.	
19	082	275	526	10
18	580	5074	446	6
17	3480	074	65	54
17	079	1573	205	55
16	1878	1272	325	40
16	077	5072	105	40
10	1071	264	496	13
9	5270	1264	255	47

CHAP. VI.

*Containing such nauticall questions, as are
of ordinary vse, concerning longitude,
latitude, Rumb, and distance.*

1 *To keep an account of the ships way.*

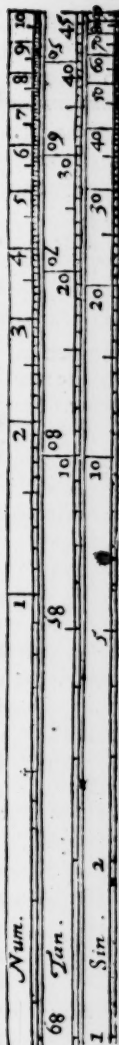
THe way that the ship maketh, may be knowne to an old sea-man by experience, by others it may be found for some small portion of time, either by the log line, or by the distance of two knowne markes on the ships side. The time in which it maketh this way, may be measured by a watch, or by a glasse, Then as long as the wind continueth at the same stay, it followeth by proportion,

As the time giuen is to an houre:
So the way made, to an houres way.

Suppose the time to be 15 seconds, which make a quarter of a minute, and the way of the ship 88 feet: then because there are 3600 seconds in an houre, I may extend the compasses in the line of *numbers*, from 15 vnto 3600, and the same extent will reach from 88 vnto 21120.

Or I may extend them from 15 vnto 88, and this extent will reach from 3600 vnto 21120; which shewes that an houres way came to 21120 feete.

But this were an vnecessary businesse, to hearken after feet or fadoms. It sufficeth our sea-men to find the way of their ship in leagues or miles. And they say that there are 5 feet in a pace, 1000 paces in a mile, and 60 miles in a degree, and therefore 300000 feete in a degree, Yet compa-



ring severall obseruations, and their measures with our feete vsuall about *London*, I find that we may allow 352000 feete to a degree; and then if I extend the compasses in the line of *numbers* from 352000 vnto 21120, I shall find the same extent to reach from 20 leagues the measure of one degree, to 1.2, and from 60 miles to 3.6; which shewes the houres way to be 1 league and 2 tenths of a league, or 3 miles and 6 tenths of a mile.

But to auoid these fractions and other tedious reductions, I suppose it would be more easie to keep this account of the ships way (as also of the difference of latitude, and the difference of longitude) by degrees and parts of degrees, allowing 100 parts to each degree, which we may therefore call by the name of *centesimes*. Neither would this be hard to conceiue. For if 100 such parts do make a degree, then shall 50 parts be equall to 30 minutes, as 30 minutes are equall to 10 leagues. And 5 parts shall be equall to 3 minutes, as 3 minutes are equall to 1 league. And so the same extent as before, will reach from 100 parts vnto 6; which shewes that the houres way required is 6 *cent.* such as 100 do make a degree, and 5 do make an ordinary league.

This might also be done at one operation. For vpon these suppositions, diuide 44 feet into 4; lengths, and set as many of them as you may conueniently betwene two markes on the ships side, and note the seconds of time in which the ship goeth these lengths: so the lengths diuided by the time, shall giue the *cent.* which the ship goeth in an houre.

Suppose the distance betwene the two markes to be 60 lengths (which are 58 feet and 8 inches) and let the time be 12 seconds: extend the compasses from 12 to 1, in the line of *numbers*; so the same extent will reach from 60 vnto 5. Or extend them from 12 vnto 60, and the same extent will reach from 1 vnto 5. This shewes that the ships way is according to 5 *cent.* in an houre.

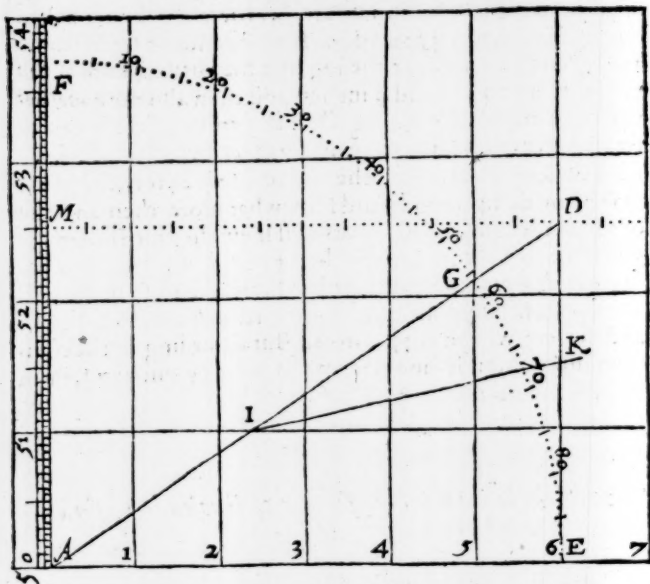
This may be found yet more easily, if the log line shall be fitted to the time. As if the time be 45 seconds, the log line may haue a knot at the end of euery 44 feete; then doth the ship

ship run so many *cent.* in an houre, as there are knots vered out in the space of 45 seconds. If 30 seconds do seeme to be a more conuenient time, the log line may haue a knot at the end of euery 29 feet and 4 inches; and then also the *centesmet* will be as many as the knots. Or if the knots be made to any set number of feet, the time may be fitted vnto the distance. As if the knots be made at the end of euery 24 feet, the glasse may be made 24 seconds and somewhat more then an halfe of a second; and so these knots will shew the *cent.* If there be 5 knots vered out in a glasse, then 5 *cent.*; if 6 knots, then the ship goeth 6 *cent.* in the space of an houre; and so in the rest. For vpon this supposition, the proportion between the time and the feet will be as 45 vnto 44. But according to the common supposition it should seeme to be as 45 vnto 37 $\frac{1}{2}$, or in lesser termes as 6 vnto 5. Those which are vpon the place, may make prooffe of both, and follow that which agrees best with their experience.

2 *By the latitude and difference of longitude, to find the distance vpon a course of East and West.*

Extend the compasses from the sine of 90 *gr.* vnto the sine of the complement of the latitude; the same extent shal reach in the line of *numbers* from the difference of longitude to the distance.

So the measure of one degree in the equator, being 100 *cent.* the distance belonging to one degree of longitude in the latitude of 51 *gr.* 30 *m.* will be found about 62 *cent.* and $\frac{1}{2}$. Or if the measure of a degree be 60 miles, the distance will be found about 37 miles and $\frac{1}{2}$. If the measure be 20 leagues, then almost 12 leagues and $\frac{1}{2}$. If the measure be 17 $\frac{1}{2}$, as in the Spanish charts, then somewhat lesse then 11 leagues sailing vpon this parallell, will giue an alteration of one degree of longitude.



3 By the latitude and distance upon a course of East or west, to find the difference of longitude.

Extend the compasses from the sine of the complement of the latitude, to the sine of 90 gr; the same extent will reach in the line of *numbers* from the distance to the difference of longitude.

So the distance upon a course of East or West, in the latitude of 51 gr. 30 m. being 100 cent. the difference of longitude will be found 1.60, which make one degree and 60 *centesimes* or 1 gr. 36 m.

Or if it be 60 miles, the difference of longitude will be 96, which also make 1 gr. 36 m. as before.

- 4 *The longitude and latitude of two places being given,
so find the Rumb leading from the one
to the other.*

Extend the compasses in the line of *numbers* from the difference of latitudes to the difference of longitudes; the same extent will give the distance from the tangent of 45 gr. unto the tangent of the Rumb, according to the projection of the common sea-chart.

So the latitude of the first place being 50 gr. the latitude of the second $52\text{ gr. }30\text{ m.}$ and the difference of longitude 6 gr. the Rumb will be found to be about $67\text{ gr. }23\text{ m.}$ which is neare the inclination of the sixth Rumb to the meridian. But this Rumb so found, is alwayes greater then it should be, and therefore to be limited; which may be done sufficiently for the Sea-mans vse, after this manner:

Extend the compasses either from the sine of 90 gr. unto the sine of the complement of the middle latitude, the same extent will reach from the tangent of the Rumb before found, to the tangent of the Rumb limited.

Or else extend them from the sine of 90 gr. unto the tangent of the Rumb before found; the same extent will reach from the sine of the complement of the middle latitude, unto the tangent of the Rumb limited.

So the middle latitude between 50 gr. and $52\text{ gr. }30\text{ m.}$ being $51\text{ gr. }15\text{ m.}$ and the Rumb before found $67\text{ gr. }23\text{ m.}$ the Rumb limited will be found to be about $56\text{ gr. }20\text{ m.}$ which is but five minutes more then the inclination of the fifth Rumb to the meridian.

2 This Rumb may be found by the help of the *meridian line* upon the Staffe. For if I take the difference of latitude out of the *meridian line* from 50 gr. unto $52\text{ gr. }30\text{ m.}$ and measure it in his equinoctiall, or at the beginning of the *meridian line*, I shall find it there to be equal to 4 gr. Wherefore I work as if the difference of latitude were 4 gr. and extend the compasses in the line of *numbers* from $4.$ unto $6.$ so shall I finde
the

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the same extent to reach from the tangent of 45 gr. vnto the tangent of 56 gr. 20 m. and this is the inclination of the Rumb required.

5 *By the Rumb and both latitudes, to find the distance vpon the Rumb.*

Extend the compasses from the sine of the complement of the Rumb, vnto the sine of 90 gr. the same extent in the line of *numbers* shall reach from the difference of latitude vnto the distance vpon the Rumb.

So the latitude of the first place being 50 gr. the latitude of the second 52 gr. 30 m. and the Rumb the fift from the meridian. If I extend the compasses from 33 gr. 45 m. vnto the sine of 90 gr. I shall find the same extent in the line of *numbers* to reach from 2 gr. 50 cent. to 4 gr. 50 cent. and such is the distance required.

6 *By the distance and both latitudes to find the Rumb.*

Extend the compasses in the line of *numbers* from the distance vnto the difference of latitudes; the same extent will reach in the line of *sines*, from 90 gr. vnto the complement of the Rumb.

So the one place being in the latitude of 50 gr. the other in the latitude of 52 gr. 30 m. and the distance between them 4 gr. 50 cent. If I extend the compasses from 4.50 vnto 2.50 in the line of *numbers*, I shall find the same extent to reach from the sine of 90 gr. vnto the complement of 56 gr. 15 m. and such is the inclination of the Rumb required.

7 *By one latitude, Rumb, and distance, to find the difference of latitudes.*

Extend the compasses in the line of *sines*, from 90 gr. vnto the complement of the Rumb; the same extent in the line of *numbers*

numbers, will reach from the distance, vnto the difference of latitudes.

So the lesser altitude being 50 *gr.* and the distance 4 *gr.* 50 *cent.* vpon the fifth Rumb from the meridian: if I extend the compasses from the line of 90 *gr.* to 33 *gr.* 45 *m.* I shall finde the same extent to reach from 4.50 in the line of *numbers*, vnto 2.50; and therefore the second latitude to be 52 *gr.* 30 *m.*

8 *By the Rumb and both latitudes, to find the difference of longitude.*

Extend the compasses from the tangent of 45 *gr.* vnto the tangent of the Rumb; the same extent will reach in the line of *numbers* from the difference of latitudes vnto the difference of longitude, according to the projection of the common sea-chart.

So the first latitude being 50 *gr.* and the second 52 *gr.* 30 *m.* and the Rumb the fifth from the meridian: if I extend the compasses from the tangent of 45 *gr.* vnto 56 *gr.* 15 *m.* I shall find the same extent to reach from 2.50 in the line of *numbers* to about 3.75, which make 3 *gr.* 45 *m.* But this difference of longitude so found, is alwayes lesser then it should be, and therefore to be enlarged, which may be done sufficiently for the sea-mens vse, after this maner:

Extend the compasses from the line of the complement of the middle latitude, vnto the line of 90 *gr.* the same extent will reach in the line of *numbers* from the difference of longitude before found, vnto the difference of longitude enlarged.

So the middle latitude in this example being 51 *gr.* 15 *m.* and the difference of longitude before found 3 *gr.* 75 *cent.* the difference of longitude enlarged will be found about 5 *gr.* 99 *cent.* which are neare 6 *gr.*

2 This difference of longitude may be found by help of the *meridian line* vpon the Staffe. For if I take the proper difference of latitude out of the meridian line, and measure it in his equinoctiall, or at the beginning of the meridian line,

I shall find it to be equall to foure of those degrees. Wherefore hauing extended the compalles as before from the tangent of 45 gr. vnto the tangent of 56 gr. 15 m. the same extent will reach from 4.00 in the line of *numbers*, vnto 5.99: which shewes the difference of longitude to be about 5 gr. 99 cent. or about halfe a minute short of six degrees.

9 *By the Rumb and both latitudes, to finde the distance belonging to the chart of Mercators projection.*

Take the proper difference of latitudes out of the meridian line of the chart, and measure it in his equinoctiall, or one of the parallels, and it will there giue the difference of latitudes enlarged. Then extend the compalles from the sine of the complement of the Rumb vnto the sine of 90 gr. the same extent will reach in the line of *numbers*, from the latitude enlarged, vnto the distance required. Or extend them from the complement of the Rumb to the latitude enlarged, the same extent will reach from 90 gr. vnto the distance.

For example, let the place giuen be *A* in the latitude of 50 gr. *D* in the latitude of 52 gr. 30 m. *AM* the difference of latitudes, and the Rumb *MAD* the fifth from the meridian. First I take out *AM* the difference of latitudes, and measure it in *AE* one of the parallels of the equinoctiall; I find it to be very neare 4 gr: this is the difference of latitudes enlarged. Then if I extend the compalles from the sine of 33 gr. 45 m. the complement of the fifth Rumb vnto the sine 90 gr. I shall find the same extent to reach in the line of *numbers* from 4.00 vnto 7.20. And this is the distance belonging to the chart. Wherefore I take out these 7 gr. 20 cent. out of the scale of the parallell *AE*, and pricke it downe vpon the Rumb from *A* vnto *D*, where it meeteth with the parallell of the second latitude. Lastly, I measure it in the *meridian line*, setting one foote of the compalles as much below the lesser latitude as the other aboue the greater latitude, and find it to be 4 gr. 50 cent: which is the same distance that I found before in the 5. *Prop.*

As the sine of the angle opposite to the knowne side,
is to that knowne side

So the sine of the angle opposite to the side required,
is to the side required.

Wherefore I extend the compasses from 14 gr. 40 m. in the sines, to 10 in line of numbers; and this extent doth reach from 58 gr. to $33\frac{1}{2}$, and such is the distance between *A* and *B*, and it reacheth from 43 gr. 20 m. vnto 27 in the line of numbers; and such is the distance from *D* to *B*.

These two distances being knowne, I may set out the land vpon the chart. For hauing set downe the way of the ship from *A* to *D* by that which I shewed before in the vse of the meridian line, I may by the same reason set off the distance *AB* and *DB*, which meeting in the point *B*, shall there resemble the land required.

II *By knowing the distance between two places on the lands, and how they beare one from the other, and hauing the angles of position at the ship to find the distance betweene the ship and the land.*

If it may be conueniently, let the angle of position be obserued at such time as the ship cometh to be right ouer against one of the places. As if the places be East and West, seeke to bring one of them South or North from you, and then obserue the angle of position: so shall you haue a right line triangle, with one side and three angles, whereby to find the two other sides. First you haue the angle of position at the ship; then a right angle at the place that is ouer against you; and the third angle at the other place is the complement to the angle of position. Wherefore

As the sine of the angle of position,
is to the distance betweene the two places:

So the cosine of the angle of position,
to the distance betweene the ship and the nearer place.
And

And so is the sine of 90 gr.
to the distance from the ship to the farther place.

So the places being 15 cent. or three leagues one from the other, and the angle of position 29 gr.; the nearer distance will be found about 27 cent. and the farther distance about 31 cent.

Or howsoever the angle of position were observed, the distance between the ship and the land may be found generally as in this example:

Suppose *A* and *D* were two head lands knowne to be East Northeast, and West Southwest, 10 cent. or two leagues one from the other; and that the ship being at *B*, I observed the angle of the ships position *DBA*, and found it to be 14 gr. 40 m. and that *D* did beare 9 gr. 30 m. and *A* 24 gr. 10 m. from the meridian *BS*; this example would be like the former. For if the angle *SBD* be 9 gr. 30 m. from the South to the Westward, then shall *NDB* be 9 gr. 30 m. from the North to the Eastward. Take these 9 gr. 30 m. out of the angle *NDE* which is 67 gr. 30 m. because the two head lands lie East Northeast, and there will remaine 58 gr. for the angle *BDE*, and the inward angle *BDA* shall be 122 gr. Take these two angles *ABD* and *BDA* out of 180 gr. and there will remaine 43 gr. 20 m. for the third angle *BAD*. Wherefore here also are three angles and one side, by which I may find the two other sides, as in the last *Prop.*

These propositions thus wrought by the Staffe, are such as I thought to be vsfull for sea-men, and those that are skilfull may apply the example to many others. Those that begin, and are willing to practise, may busie themselves with this which followeth.

Suppose foure ports, *L, N, O, P*; of which *L* is in the latitude of 50 gr. *N* is North from *L* 200 leagues or 1000 centesmes; *O* West from *L* 1000 cent. and *P* West from *N* 1000 cent: so that *L* and *O* will be in the same latitude of 50 gr. *N* and *P* both in the latitude of 60 gr. Then let two ships depart from *L*, the one to touch at *O*, the other at *N*,

and then both to meet at *P*, there to lade, and from thence to returne the neareſt way vnto *L*. Here many questions may be propoſed.

- 1 What is the longitude of the port at *O*?
- 2 What is the longitude of *P*? And why *O* and *P* ſhould not be in the ſame longitude?
- 3 What is the Rumb from *O* vnto *P*?
- 4 What is the diſtance from *O* vnto *P*? And why the way ſhould be more from *L* vnto *P*, going by *O*, then by *N*?
- 5 What is the Rumb from *P* vnto *L*?
- 6 What is the diſtance from *P* vnto *L*?
- 7 What is the Rumb from *N* vnto *O*?
- 8 What is the diſtance from *N* vnto *O*? And why it ſhould not be the like Rumb and diſtance from *N* vnto *O*, as from *P* vnto *L*?

Theſe questions well conſidered, and either reſolued by the Staffe, or pricked downe on the chart, and compared with the globe and the common Sea-chart, will giue ſome light to the direction of a courſe, and reduction of places to their due longitude, which are now ſouly diſtorted in the common Sea-charts.

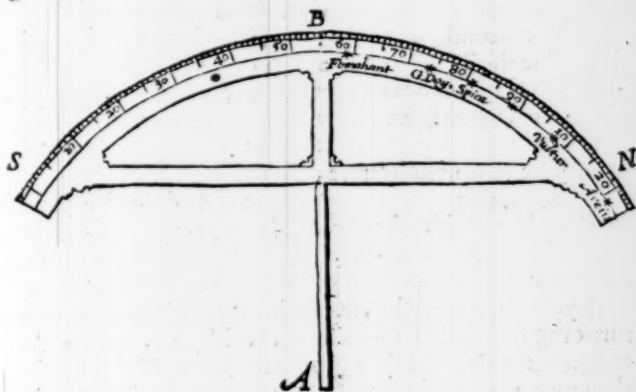
An

An Appendix concerning

*The description and use of an instrument, made
in forme of a Crosse-bow, for the more easie
finding of the latitude at Sea.*

THe former *Prop.* suppose the latitude to be knowne,
I will here shew how it may be easily obserued.

Vpon the center *A*, and semidiameter *AB*, describe an ark
of a circle *SBN*. The same semidiameter will set of 60 *gr.*
from *B* vnto *S* for the South end, and other 60 *gr.* from *B* vnto
N for the North end of the Bow: so the whole Bow will
containe 120 *gr.* the third part of a circle. Let it therefore be
diuided into so many degrees, and each degree subdiuided
into six parts, that each part may be ten minutes: but let the
numbers set to it be 5. 10. 15. vnto 90 *gr.* and then againe
5. 10. 15. vnto 25. that 55 may fall in the middle, as in this fi-
gure.



The Bow being thus diuided and numbred, you may see
the moneths and dayes of each moneth vpon the backe, and
such

such starres as are fit for obseruation vpon the side of the Bow.

If you desire to make vse of it in North latitude, you may number 23 gr. 30 m. from 90 towards the end of the Bowe at N, and there place the tenth day of Iune. And 23 gr. 30 m. from 90 toward S; and there at 66 gr. 30 m. place the tenth day of December. And so the rest of the dayes of the yeare, according to the declination of the Sunne at the same dayes. The starres may be placed in like maner according to their declinations.

Arcturus at 21 gr. 10 m.

The Bulls eye 15 42

The Lions heart 13 45

The Vultures heart 7 58

The little dog 6 9 from 90 toward the

North end of the Bow at N. Then for Southerne starres, you may number their declination from 90 toward the South end of the Bow at S. As first the three starres in *Orions* girdle,

The first at 0 gr. 37 m.

The second 1 28

The third 2 11

The Hydra's heart 7 5

The virgins spike 9 10

The great dog 16 12

The Scorpions heart 25 30

Fomahant 31 30 And so the South

crowne, the triangle, the clouds, the crossiers, or what other starres you think fit for obseruation. This I call the fore side of the Bow.

If you desire to make vse of it in South latitude, you may turne the Bow, and diuide the backe side of it, and number it in like maner; and then put on the moneths and dayes of the yeare, placing the tenth of December at the South end, and the tenth of Iune toward the middle of the Bow, and the rest of the dayes according to the Sunnes declination as before.

The chiefeft of the Northerne ftarres may here be placed in like maner according to their declination, Anno 1625.

The pole ftarre at	87	gr. 20 m.
The first guard	75	45
The fecond guard	73	25
The great Beares backe	63	45
In the great	}	first 58 2
Beares taile		second 57 55
		third 51 15
The fide of Perfeus	48	28
The goate	45	33
The taile of the fwan	44	0
The head of Medufa	39	30
The harp	38	30
Caftor	32	38
Pollux	28	52
The North crowne	28	0
The Rams head	21	40
Arcturus	21	10
The Bulls eye	15	42
The Lions heart	13	45
The Vultures heart	7	58
Orions right foulder	7	17
Orions left foulder	5	57

And fo any other ftarre, whose declination is knowne vnto you, which being done. The vse of this Bow may be

- 1 *The day of the moneth being knowne, to finde the declination of the Sunne.*
- 2 *The declination being giuen, to finde the day of the moneth.*

These two Prop. depend on the making of the Bow. If the day be knowne, looke it out in the backe of the Bow: to the decl nation will appeare in the fide. Or if the declination be knowne, the day of the moneth is fet ouer againft it. As if the day of the moneth were the 14 of Iuly: looke for

this day in the backe of the Bow, and you shall find it ouer against ~~20~~ *20* gr. of North declination. If the declination giuen be 20 gr. to the Southward, you shall find the day to be either the eleuenth of Nouember, or the eleuenth of Ianuary.

3 To find the altitude of the Sunne or starres.

Here it is fit to haue two running sights, which may be easily moued on the backe of the Bow. The vpper sight may be set either to 60 gr. or to 70 gr. or to 80 gr. as you shall find to be most conuenient: the other sight may be set on, to any place betweene the middle and the other end of the Bow. Then with the one hand hold the center of the Bow to your eye, so as you may see the Sunne or starre by the vpper sight, and with the other hand moue the lower sight vp or downe vntill you haue brought one of the edges of it to be euen with the horizon: (as when you obserue with the Crosse-staffe:) so the degrees contained betweene that edge and the vpper sight, shall shew the altitude required.

Thus if the vpper sight shal be at 80 gr. and the lower sight at 50 gr. the altitude required is 30 gr.

4 To find any North latitude, by knowing either the day of the moneth, or the declination of the Sunne.

As oft as you are to obserue in North latitude, place both the sights on the fore side of the Bow, the vpper sight to the declination of the Sunne, or the day of the moneth at the North end, and the lower sight toward the South end. Then when the Sunne cometh to the meridian, turne your face to the South, and with the one hand hold the center of the Bow to your eye, so as you may see the Sunne by the vpper sight; with the other hand moue the lower sight, vntill you haue brought one of the edges of it to be euen with the horizon: so that edge of the lower sight shall shew the latitude of the place in the fore side of the Bow.

Thus being in North latitude vpon the ninth of October

if

if I set the vpper sight to this day, at the fore side and North end of the Bow, I shall find it to fall to the Southward of 90 vpon 80 gr. and therefore at 10 gr. of South declination. Then the Sunne coming to the meridian, I may set the center of the Bow to mine eye, as if I went to find the altitude of the Sunne, holding the North end of the Bow vpperward, with the vpper sight betweene mine eye and the Sunne, and mouing the lower sight, vntill it come to be euen with the horizon. It here the lower sight shall stay at 50 gr. I may well say, that the latitude is 50 gr. For the meridian altitude of the Sunne is 30 gr. by the last *Prop.* and the Sun hauing 10 gr. of South declination, the meridian altitude of the equator would be 40 gr., and therefore the obseruation was made in 50 gr. of North latitude.

By the same reason, if the lower side had stayed at 51 gr. 30 m. the latitude must haue been 51 gr. 30 m. and so in the rest.

5 To find any North latitude, by the meridian altitude of the starres to the Southward.

Let the vpper sight be set to the starre, which you intend to obserue, here placed in the fore side of the Bow. Then hold the North end of the Bow vpperward, and turning your face to the South, obserue the meridian altitude as before: so the lower sight shall shew the latitude of the place in the fore side of the Bow.

Thus if in obseruing the meridian altitude of the great Dog-starre, the lower sight shall stay at 50 gr. it would shew the latitude to be 50 gr. For this starre being here placed at 73 gr. 48 m. if we take thence 50 gr. his meridian altitude would be 23 gr. 48 m. to this if we adde 16 gr. 12 m. for the South declination of this starre, it would shew the meridian altitude of the equator to be 40 gr. and therefore the latitude to be 50 gr.

6 *To find any North latitude, by the meridian altitude of the starres to the Northward.*

Let the vpper sight be set to the starre which you intend to obserue, here placed on the backe side of the Bow. Then hold the North end of the Bow vpward, and turning your face to the North, obserue the altitude of the starre when he cometh to be in the meridian and vnder the pole: so the lower sight shall shew the altitude of the pole in the back side of the Bow.

Thus the former guard coming to be in the meridian vnder the pole, if you obserue and find the lower sight to stay at 50 gr. such is the eleuation of the pole, and the latitude of the place to the Northward. For the distance betweene the two sights will shew the altitude to be 35 gr. 45 m. & the star is 14 gr. 15 m. distant from the North pole. These two doe make vp 50 gr. for the eleuation of the North-pole, and therefore such is the North latitude.

7 *To find any South latitude, by knowing either the day of the moneth, or the declination of the Sunne.*

When you are come into South latitude, turne both your sights to the backside of the Bow: the vpper sight to the declination of the Sun, or the day of the moneth at the South end, and the lower sight toward the North end of the Bow. Then the Sunne coming to the meridian, turne your face to the North, and holding the South end of the Bow vpward, obserue the meridian altitude as before: so the lower sight shall shew the latitude of the place in the backe side of the Bow.

Thus being in South latitude, vpon the tenth of May if you obserue and finde the lower sight to stay at 30 gr. on the backe side of the Bow, such is the latitude. For the declination is 20 gr. Northward, the altitude of the Sunne betweene the two sights 40 gr. the altitude of the equator 60 gr. and there-

therefore the latitude 30 gr.

8 *To find any South latitude, by the meridian altitude of the starres to the Northward.*

Let the vpper sight be set to the starre which you intend to obserue, here placed on the backe side of the Bow. Then hold the South end of the Bow vpward, and turning your face to the North, obserue the meridian altitude as before: so the lower sight shall shew the latitude of the place in the back side of the Bow.

Thus being in South latitude, and the former guard coming to be in the meridian ouer the pole. If you obserue and finde the lower sight to stay at 5 gr. such is the latitude. For this starre is 14 gr. 15 m. from the North pole, the altitude of the starre betweene the two sights 9 gr. 15 m. the North pole depressed 5 gr. and therefore the latitude 5 gr. to the Southward.

9 *To obserue the altitude of the Sunne backward.*

Set the vpper sight either to 60, or 70, or 80 gr. as you shall find it to be most conuenient, the lower sight on any place betweene the middle and the other end of the Bow, and haue an horizontall sight to be set to the center. Then may you turne your backe to the Sunne, and the back of the Bow toward your selfe, looking by the lower sight through the horizontall sight, and mouing the lower sight vp & downe, vntill the vpper sight doe cast a shadow vpon the middle of the horizontall sight: so the degrees contained betweene the two sights on the Bow, shall giue the altitude required.

Thus if the vpper sight shall be at 80 gr. and the lower sight at 50 gr. the altitude required is 30 gr. as in the third Prop.

- 10 *To find any North latitude by a backe observation, knowing either the day of the moneth, or the declination of the Sunne.*

When you obserue in North latitude, place your three sights on the fore side of the Bow: the vpper sight to the declination of the Sun, or the day of the moneth, at the North end; the lower sight toward the South end of the Bow; and the horizontall sight to the center. Then the Sunne coming to the meridian, turne your face to the North, & holding the North end of the Bow vpward, the South end downeward, with the back of it toward your selfe, obserue the shadow of the vpper sight as in the former *Prop.* so the lower sight shall shew the latitude of the place in the fore side of the Bow.

Thus being in North latitude vpon the ninth of October, if you obserue and find the lower sight to stay at 50 gr. on the fore side of the Bow, such is the latitude. For the declination is 10 gr. Southward, and the altitude of the Sunne betwene the two sights 30 gr. the altitude of the equator 40 gr. and therefore the latitude 50 gr. as in the fourth *Prop.*

- 11 *To find any South latitude by a backe observation, knowing either the day of the moneth, or the declination of the Sunne.*

When you obserue in South latitude, place your three sights on the backe side of the Bow: the vpper sight to the declination of the Sunne, or the day of the moneth at the South end; the lower sight toward the North end of the Bow, and the horizontall sight to the center. Then the Sun coming to the meridian, turne your face to the South, and holding the South end of the Bow vpward, with the backe of it toward your selfe, obserue the shadow of the vpper sight as before: so the lower sight shall shew the latitude of the place in the backe side of the Bow.

Thus being in the South latitude vpon the tenth of May,

if

if you obserue and find the lower sight to stay at 30 gr. on the backe of the Bow, such is the altitude. For the declination is 20 gr. Northward, the altitude of the Sunne betweene the two sights 40 gr. the altitude of the equator 60 gr. and therefore the latitude 30 gr. as in the seuenth Prop.

13 To find the day of the moneth, by knowing the latitude of the place, and obseruing the meridian altitude of the Sunne.

Place your three sights according to your latitude; the horizontall sight to the center, the lower sight to the latitude, and the vpper sight among the moneths. Then when the Sunne cometh to the meridian, obserue the altitude, looking by the lower sight through the horizontall, and keeping the lower sight still at the latitude, but mouing the vpper sight vntill it giue shadow vpon the middle of the horizontall sight: so the vpper sight shall shew the day of the moneth required.

Thus in our latitude if you set the lower sight to 51 gr. 30 m. and obseruing finde the altitude of the Sunne betweene that and the vpper sight to be 28 gr. 30 m. this vpper sight will fall vpon the ninth of October, and the twelfth of Februarie. And if yet you doubt which of them two is the day, you may expect another meridian altitude; and then if you find the vpper sight vpon the tenth of October, and the eleuenth of Februarie, the question will be soone resolued.

13 To find the declination of any vnkownne starre, and so to place it on the Bow, by knowing the latitude of the place, and obseruing the Meridian altitude of the Starre.

When you find a starre in the Meridian that is fit for obseruation. Set the center of the Bow to your eye, the lower sight to the latitude, and moue the vpper sight vp or downe vntill you see the horizon by the lower sight, and the starre
by

by the vpper sight, then will the vpper sight stay at the declination and place of the starre.

Thus being in 20 gr. of North latitude, if you obserue and find the meridian altitude of the head of the Crozier to be 14 gr. 50 m. The vpper sight will stay at 34 gr. 50 m. and there may you place this starre. For by this obseruation the distance of this starre from the South pole should be 34 gr. 50 m. and the declination from the equator 55 gr. 10 m. And so for the rest.

The starres which I mentioned before, do come to the meridian in this order, after the first point of *Aries*.

	Ho.	Mo.		Ho.	Mo.
The pole starre at	0	29	The lions hart	9	48
The rams head	1	46	The great beares backe	10	40
The head of Medusa	2	44	First in gr. beares taile	12	37
The side of Perleus	2	58	The Virgins pike	13	5
The Bulls eye.	4	15	Second in gr. bea. taile	13	9
The goate	4	49	Third in gr. beares taile	13	33
Orions left shoulder	5	5	Arcturus	13	58
Orions { the first	5	13	The foremost guard	14	52
girdle { the second	5	17	The North crowne	15	19
{ the third,	5	21	The hindmost guard	15	25
Orions right shoulder	5	35	Scorpions hart	16	7
The great dog	6	39	The harpe	18	24
Castor	7	10	Vulturs hart	19	33
The little dog	7	20	Swans taile	20	29
Pollux	7	22	Fomahant	22	36
The Hydra's hart	9	9			

The end of the second Booke.

